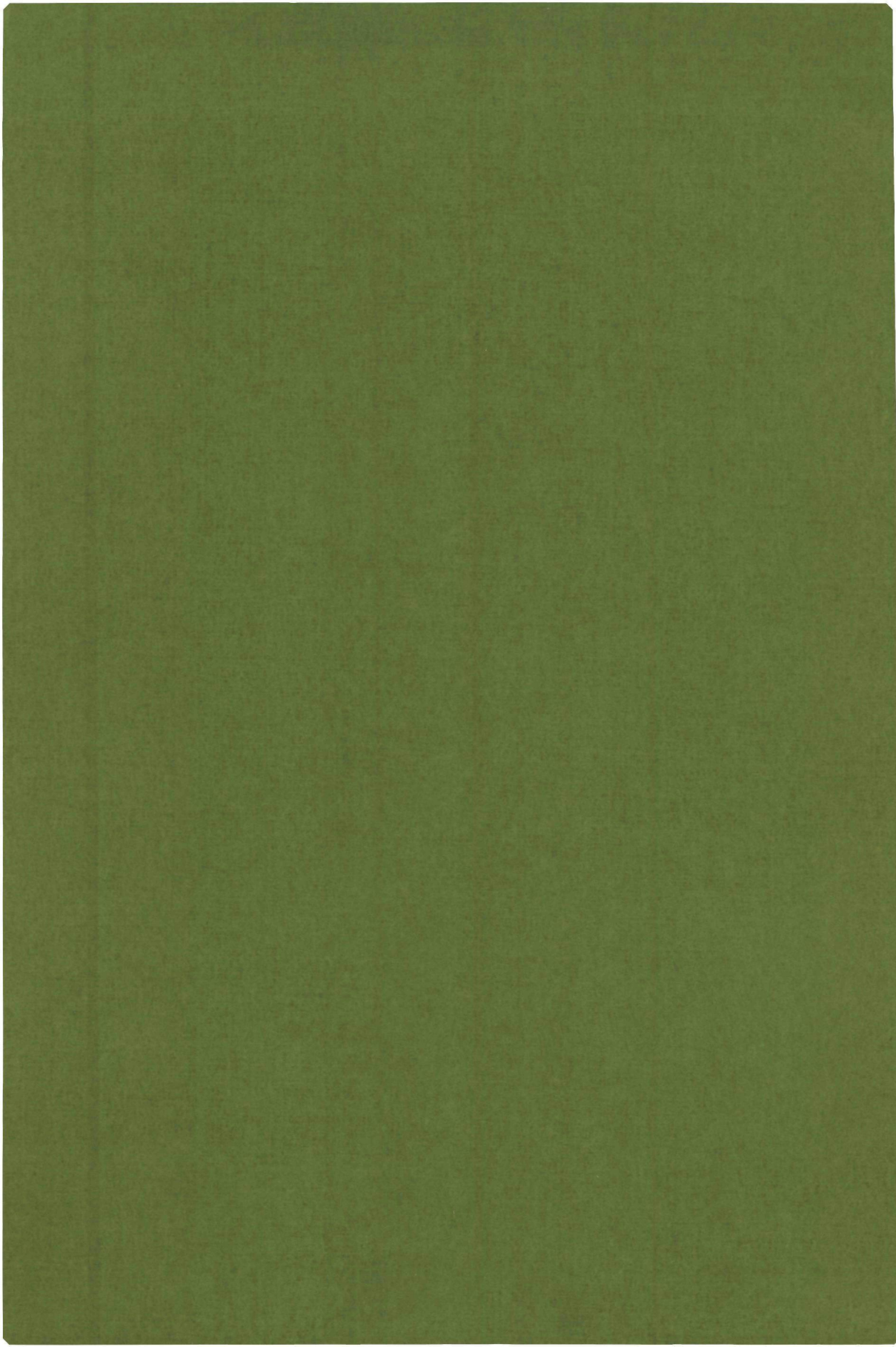


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**A FAR-INFRARED MICHELSON INTERFEROMETER  
AND ITS APPLICATION TO THE STUDY OF  
PHOTOCONDUCTIVITY IN ULTRA-PURE GERMANIUM**

**H.W.H.M. JONGBLOETS**



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**PROMOTOR:  
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**A FAR-INFRARED MICHELSON INTERFEROMETER AND ITS APPLICATION TO THE  
STUDY OF PHOTOCONDUCTIVITY IN ULTRA-PURE GERMANIUM**

**PROEFSCHRIFT**

TER VERKRIJGING VAN DE GRAAD VAN DOCTOR IN DE  
WISKUNDE EN NATUURWETENSCHAPPEN  
AAN DE KATHOLIEKE UNIVERSITEIT TE NIJMEGEN, OP GEZAG VAN  
DE RECTOR MAGNIFICUS PROF. DR. P.G.A.B. WIJDEVELD,  
VOLGENS BESLUIT VAN HET COLLEGE VAN DECANEN  
IN HET OPENBAAR TE VERDEDIGEN  
OP DONDERDAG 13 MAART 1980  
DES NAMIDDAGS TE 4.00 UUR

door

HENDRIKUS WILHELMUS HUBERTUS MARIA JONGBLOETS  
geboren te Nijmegen

1980  
Druk: Krips Repro Meppel

The investigations described in this thesis have been carried out in the group "Experimentele Natuurkunde IV" of the Research Institute for Materials of the Faculty of Science at the Catholic University of Nijmegen under the direction of Prof. Dr. P. Wyder.

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In solid state physics, the relevant energies of many interesting effects correspond with temperatures of the order of 1 - 1000 degrees Kelvin. This range covers the binding energies of impurities in semiconductors, the ordering temperatures in ferromagnetism, antiferromagnetism and ferroelectricity, the transition temperatures of superconductors, phonon- and Debye energies, etc. To study this interesting region, sometimes high magnetic fields can be applied. Field strengths up to 25 Tesla, corresponding to 16.7 K, are available at the Nijmegen High Magnetic Field installation.

In this framework it is obviously very much worthwhile to have electromagnetic radiation available of comparable energy. This is the far infrared (FIR) or submillimetre wave region (wave length range 50 - 1000  $\mu\text{m}$  or wave number range 10 - 200  $\text{cm}^{-1}$ ), which connects the optical and microwave parts of the electromagnetic spectrum. This is a most awkward regime because of the lack of broadband sources of sufficient power. It is true that nowadays FIR lasers are available with output power up to 0.4 W<sup>(1,2)</sup>, but these have the disadvantage of not being continuously tunable nor being very stable. The FIR region can also be approached from the microwave side by the technique of harmonic generation which offers a higher stability and so a higher resolution. With this method the range of 2 - 25  $\text{cm}^{-1}$  can easily be spanned with an output power decreasing from  $10^{-2}$  to  $10^{-8}$  W with increasing wave number.

The most widely used technique is that where the FIR region is penetrated from the optical side by employing a broadband source (medium pressure mercury arc lamp) in combination with a dispersive optical device, such as a grating monochromator<sup>(4,5,6)</sup>, or an interferometric device such as a lamellar grating interferometer<sup>(7,8)</sup> or a Michelson interferometer<sup>(9)</sup>. These instruments deliver only a radiation power which at low wave numbers is less than  $10^{-9}$  W within a bandwidth of 1  $\text{cm}^{-1}$ , but they cover the whole FIR region.

In chapter II of this thesis the hardware of a modern commercial Grubb Parsons FIR Michelson interferometer system is described, originally developed by Chantry et al.<sup>(9)</sup>. This system has been modified and improved in our laboratory by the author with the help of the instrument and electronics workshops of our Faculty. This instrument now covers the range 5 - 350  $\text{cm}^{-1}$ . A short summary of the theory underlying the technique of Fourier spectroscopy is given in chapter III (section III.1). This more conventional part of the

theory does not take into account the radiation which returns to the interferometer due to reflection of the sample. It is shown that this can lead to significant distortions in the spectra, as is analyzed both theoretically and experimentally in chapter III (section III.2) and Ref. 10. In an appendix to this thesis the computer programs, developed for the processing of the data and the execution of the Fourier transform needed to obtain a frequency spectrum, are collected.

The application of the Michelson interferometer to the study of the properties of impurity states in semiconductors is described in chapter IV. This study has been stimulated greatly by the very advanced crystal growing techniques which exist nowadays. It has become possible to fabricate ultra-pure crystals with impurity concentrations smaller than  $10^{10}$  atoms/cm<sup>3</sup>. In other words, only one out of every  $10^{12}$  atoms is an impurity. It is extremely difficult to identify these impurities by conventional methods, e.g. Hall effect measurements. However, photothermal ionization spectroscopy (PTIS) in the far infrared region, first developed by Russian scientists<sup>(11,12)</sup>, offers a possibility to study the impurities both qualitatively and quantitatively. With this technique FIR spectroscopy can be used as a tool for chemical analysis.

It is well known that the group V donor and group III acceptor impurities in group IV semiconductors (e.g. in germanium, silicon) exhibit "hydrogen-like" energy level schemes for the single bound electron/hole, with binding energies in the order of 10 meV. The PTIS technique is based on the two-step photothermal ionization process where the electron/hole is excited from the impurity groundstate to a higher level by the FIR radiation, and then subsequently thermally ionized into the band continuum, where it contributes to the electrical conductivity. Therefore, in the photoconductivity line spectra, which reflect the discrete energy level structure of the impurity, the line intensity will be temperature dependent.

Up to now little was known about the detailed mechanism of this thermal ionization, nor was the temperature dependence exactly known. The ad hoc assumption often mentioned in the literature<sup>(13-16)</sup> of a simple exponential behaviour seems not to fit the experimental results. Therefore the statistical problem was studied in detail (i.e. the transition from a localized impurity state to a delocalized bandstate); we have been able to show that a detailed theoretical understanding is possible and gives very good agreement with our experimental observations. An expression was derived for the temperature

dependence, which includes not only the energy difference between band edge and excited level, but also the degeneracy of the relevant level and the band. Moreover it contains a factor proportional to the impurity concentration.

Preliminary results are given in chapter IV, section IV.2 and Ref. 17. In section IV.3 (Ref. 18) very accurate values for the ground state energies of the Al, B and P impurities in germanium are derived from the temperature dependence. These values are in excellent agreement with existent theories (19,20). In addition, the ratio of the concentrations of the two acceptors present in our sample, is derived. This ratio is in accordance with the value obtained from the concentration profile in the original crystal from which our sample was cut. The determination of this profile is explained in section IV.4. In section IV.5 (Ref. 21) a detailed theoretical derivation of the temperature dependence is given, not only for the line intensity in a photo-thermal conductivity spectrum, but also for the signal strength in a photo-conductivity experiment.

Photoconductivity measurements in magnetic fields are described in section IV.6 (Ref. 22). The linear and quadratic Zeeman terms are determined for all lines of the boron excitation spectrum. In the literature<sup>(23,24)</sup> only values for the lowest levels are available; the corresponding values of our investigation agree with these. Some of the lines showing up in a magnetic field can be identified as states associated with light-hole Landau levels. Values for the magnetic field dependence of these Landau levels, derived from the measurements, agree very well with values from the literature<sup>(25)</sup>.

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### 1. Introduction.

For all far infrared spectroscopic experiments described in this thesis the technique of Fourier transform spectroscopy (FTS) was used. The principle of FTS involves the recording of interference fringes formed by radiation which has propagated through an interferometer. This interferogram is a function of path difference between two optical beams formed in the interferometer, and is uniquely related to the spectral distribution of the radiation. This information is recovered by the process of Fourier transformation (see chapter III).

The FTS technique offers several advantages over conventional dispersive techniques:

1. The multiplex advantage, first pointed out by Fellgett<sup>(1)</sup>, that each spectral element is observed for the whole duration of the experiment. The signal-to-noise ratio of an interferometric system will therefore be superior to that of a dispersive system (at least as long as detector noise is the dominant source of noise).

2. Jacquinot<sup>(2)</sup> showed that the information throughput of cylindrically symmetrical systems such as an interferometer, could be much higher than in systems where a slit has to be used.

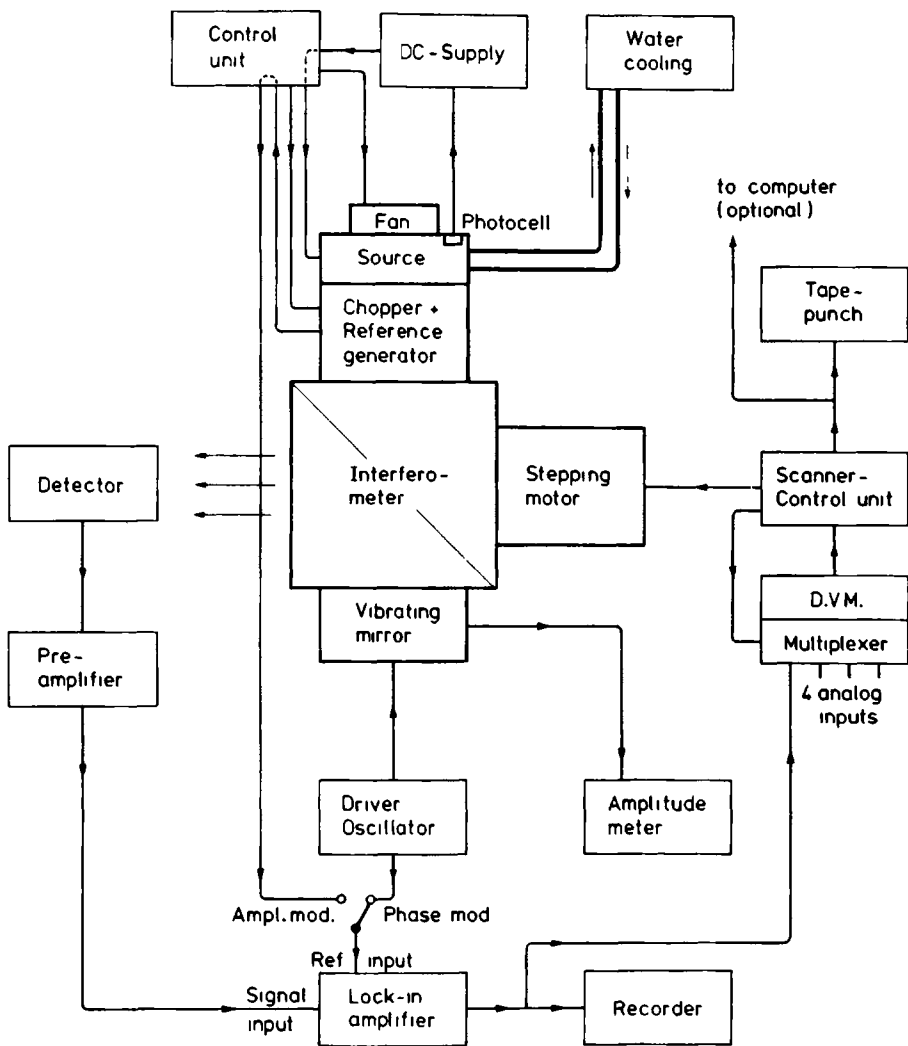
3. An interferometer is mechanically extremely simple as compared with a grating instrument. And it has the added advantage that the problem of stray light is not very severe, because only radiation that has undergone interference contributes to the interferogram.

Once the problem of transforming interferograms into spectra was overcome with the advance of modern computers, the advantages mentioned before made the FTS technique the most widely applied technique in far infrared spectroscopy. Extensive reviews of FTS have been given by Chantry<sup>(3)</sup> and Bell<sup>(4)</sup>; a short summary of the most important aspects is given in chapter III (section III.1).

### 2. Spectrometer hardware

A typical far infrared interferometric spectrometer consists of the following sub-systems:

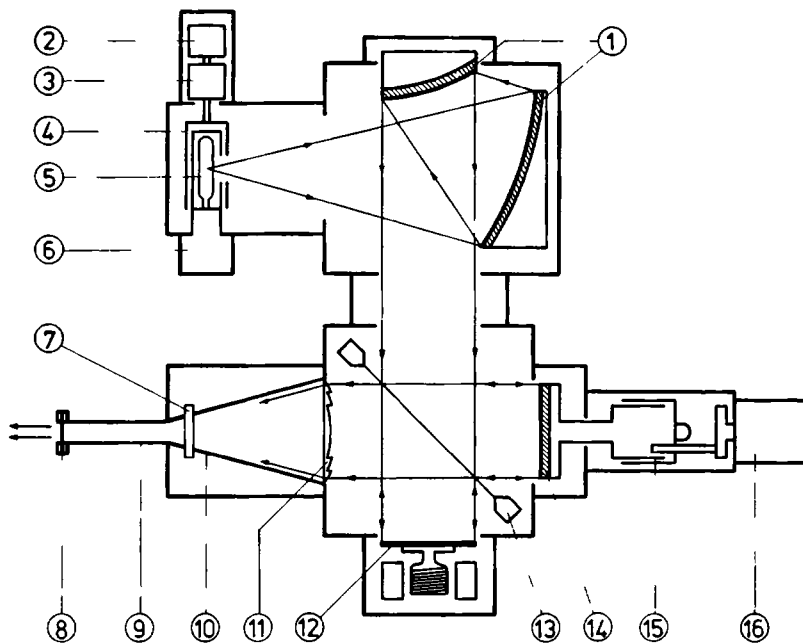
- a). A broad band source of far infrared radiation, and optics for beam collimation.



SCHEMATIC DIAGRAM OF FOURIER SPECTROMETER

*Figure 1*

# F.I.R. MICHELSON INTERFEROMETER



- |                                |                              |
|--------------------------------|------------------------------|
| ① Collimator mirrors           | ⑨ Lightpipe (polished brass) |
| ② Reference signal generator   | ⑩ Brass cone                 |
| ③ Chopper motor                | ⑪ TPX Fresnel lens           |
| ④ Chopper                      | ⑫ Vibrating mirror           |
| ⑤ Medium pressure mercury lamp | ⑬ Beamsplitter (mylar film)  |
| ⑥ Water-cooled lamp mount      | ⑭ Moving mirror              |
| ⑦ Transmission filter          | ⑮ Micrometer                 |
| ⑧ Black polyethylene window    | ⑯ Stepping motor             |

*Figure 2*

- b). The interferometer itself to create the necessary interferogram-function from the broad band radiation.
- c). A far infrared detector and signal amplifying system for recording analogue interferogram signals.
- d). A Scanner-Control Unit to drive the interferometer and containing a digital data system to digitize interferogram signals and store the data, so that they can be processed by a computer.

These subsystems can easily be recognized in the schematic diagram (figure 1) of the Fourier spectrometer, which was used for the experiments described in this thesis. The interferometer is of the Michelson type, with a thin film dielectric beamsplitter (figure 2). This modular instrument was developed at the National Physical Laboratory (Teddington, England) <sup>(5)</sup> and manufactured by Grubb Parsons Company <sup>(6)</sup>. Several details of this basic machine have been modified and improved in our laboratory.

#### a) Source

The medium pressure mercury arc lamp is still the best continuous broad band far infrared source available today, particularly at wave numbers below  $100 \text{ cm}^{-1}$  where emission from the hot electron plasma predominates. Unfortunately its power output is rather low. In the whole  $10 - 300 \text{ cm}^{-1}$  band it delivers  $\sim 5 \text{ mW.cm}^{-2} \cdot \text{sterad}^{-1}$ . Above  $100 \text{ cm}^{-1}$ , strong absorption and reemission from the fused quartz outer mantle takes place, which makes the total radiative output extremely sensitive to effects of convective cooling. For high stability, the lamp operates on d.c., under constant vacuum and in a water-cooled jacket. The special d.c. power supply for the Philips HPK 125 W mercury lamp was custom-built by Heinzinger <sup>(7)</sup>; it uses a photo-cell feed-back circuit to stabilize the radiation output of the source.

The radiation can be chopped by a cylindrical chopper around the lamp (amplitude modulation; see chapter III, section III.1), and the emergent beam is collimated by a combination of a concave and convex mirror.

#### b) Michelson interferometer

The radiation from the source is divided by the beamsplitter into two beams, which are sent to two plane mirrors. One mirror can be translated over a distance of 2 cm by a micrometer screw driven by a stepping motor (Slo-Syn SS25-1002). The beams are reflected by the mirrors and, as they are mutually coherent, interference will occur as they are recombined at the beamsplitter

and travel to the detector. The emergent beam is focussed on the entrance of a polished brass lightpipe by a TPX Fresnel lens. A low-pass transmission filter and a vacuum-tight black polyethylene window remove the unwanted part of the Planck-spectrum of the source.

The division of the incident wavefront into two is achieved by a free-standing dielectric film beamsplitter; Mylar films of thicknesses between  $6\text{ }\mu\text{m}$  and  $100\text{ }\mu\text{m}$  are available, covering the spectral range of  $5\text{-}350\text{ cm}^{-1}$ . (For a description of the properties of thin film beamsplitters: see Ref. 3, chapter 3).

An alternative way of modulating the radiation is the so-called phase modulation (see chapter III, section III.1), achieved by a sinusoidal variation of the optical path difference. For this purpose, one of the plane mirrors is mounted on a vibration generator manufactured by Ling Dynamic Systems Ltd<sup>(8)</sup>.

### c) Far infrared detector

The important criteria for a detector for use in combination with a far infrared interferometer are that it should have a high sensitivity (low noise equivalent power, NEP) and high responsivity, a uniform spectral response over a wide wave number range, and a high linearity of output over a wide dynamic range of input signals. In practice these criteria are only met by thermal detectors, which use the heating by the radiation to measure the absorbed power.

Most commercial far infrared interferometers use a room temperature Golay detector<sup>(9)</sup>, which has the disadvantage of a moderate NEP ( $\sim 10^{-10}\text{ W.Hz}^{-\frac{1}{2}}$ ) and slow response time ( $\sim 30\text{ msec}$ ). Low<sup>(10)</sup> was the first to describe a germanium bolometer operated at liquid helium temperatures. Nowadays both germanium and silicon bolometers<sup>(11)</sup> are used, and often a  $^3\text{He}$  cooling system is employed to reach a temperature of  $0.35\text{ K}$ <sup>(12)</sup>, yielding a NEP of  $10^{-12} - 10^{-14}\text{ W.Hz}^{-\frac{1}{2}}$ . An extensive review of these cryogenic semiconductor bolometers was given by Pankratov and Korotkov<sup>(13)</sup>.

Figure 3 shows a typical example of a basic detection system; various modifications for specific experiments are used in our laboratory. The FIR radiation enters the system via a black polyethylene window, which blocks out all visible and near infrared radiation. A lightpipe directs the radiation through a cooled low-pass filter (removing room temperature radiation) and a conical lightpipe to an integrating cavity wherein the bolometer has been

# CRYOSTAT WITH FAR INFRARED DETECTOR

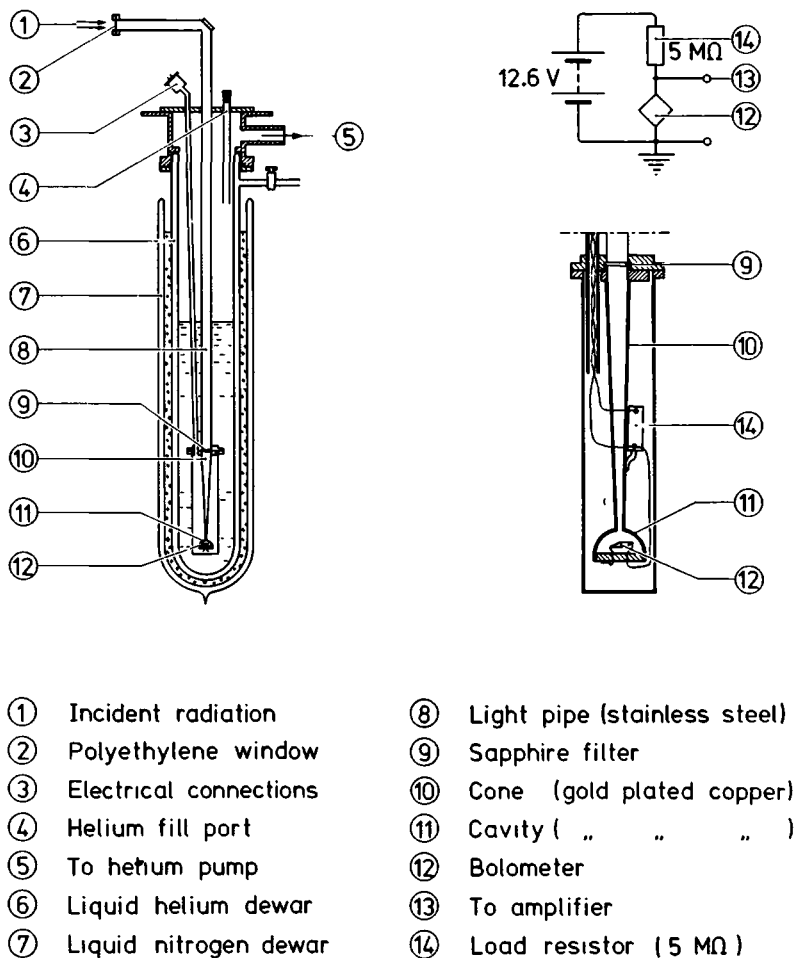
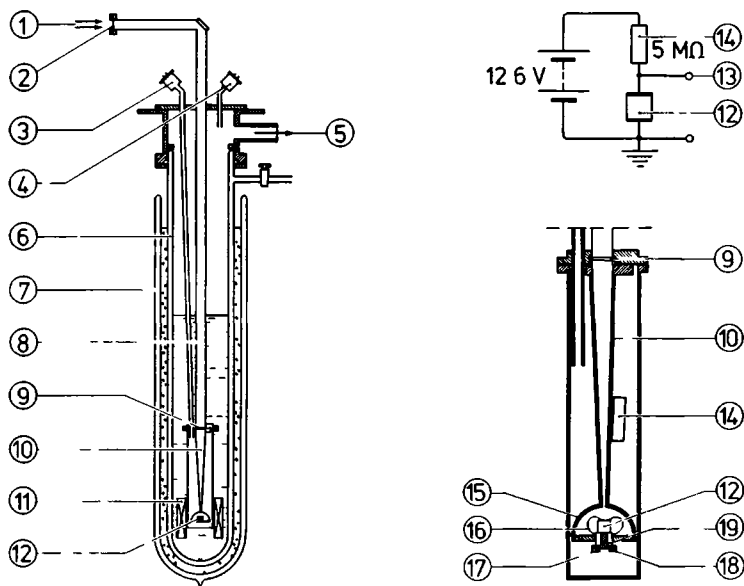


Figure 3



- |  |                               |
|--|-------------------------------|
| ① Incident radiation                     | ⑪ Superconducting magnet      |
| ② Polyethylene window                    | ⑫ Germanium sample            |
| ③ Electrical connections for Ge-sample   | ⑬ To amplifier                |
| ④ Electrical connections for s.c. magnet | ⑭ Load resistor (5MΩ)         |
| ⑤ To helium pump                         | ⑮ Cavity (gold plated copper) |
| ⑥ Liquid helium dewar                    | ⑯ Clamp                       |
| ⑦ Liquid nitrogen dewar                  | ⑰ Thermometer                 |
| ⑧ Light pipe (stainless steel)           | ⑱ Heater                      |
| ⑨ Sapphire filter                        | ⑲ Tufnol insulation           |
| ⑩ Cone (gold plated copper)              |                               |

Figure 4

mounted. The absorbed radiation causes an increase in temperature of the bolometer element; this temperature change manifests itself as a change in the electrical conductance. The bolometer is connected to an electric circuit which delivers a bias current provided by a 12.6 Volt mercury battery via a  $5\text{ M}\Omega$  cooled load resistor. The variation in the voltage drop across the bolometer, due to the chopped radiation, is measured with a lock-in amplifier.

#### d) Scanner-Control Unit

The Scanner-Control Unit contains a preset indexer to drive the stepping motor a preset number of steps for discrete sampling of the interferogram. The step-length (i.e. the number of steps between two consecutive interferogram points) is chosen appropriate to the spectrum high wave number cutoff (see also chapter III). The waiting time between sampled points is about twice the integration time of the detector lock-in amplifier. A digital volt meter ( $4\frac{1}{2}$  digit), preceded by a 4-channel multiplexer, digitizes the analogue output of the lock-in amplifier, and eventually the data are punched on papertape. The data are processed on a PDP-12 (DEC) laboratory computer, linked to a Houston incremental plotter. The relevant computer program is shown in an appendix to this thesis.

### 3. Experimental arrangement for photo-conductivity measurements.

For the photoconductivity experiments, described in chapter IV, no bolometer is needed, as the sample itself is the detector. Figure 4 shows the cryostat insert. As can be seen, the construction is almost identical to that of the bolometer system (figure 3). Only the semispherical cavity at the end of the conical lightpipe is enlarged so that it can contain the cubic sample of dimensions  $1 \times 1 \times 1\text{ cm}^3$ .

The sample is clamped onto the base of the cavity with only a small heat leak to the base. The clamp and the sample can be heated above the temperature of the surrounding helium bath by a resistance wire wrapped around the mounting stud. The temperature of the sample is measured with an Allen Bradley carbon resistor and can be stabilized within 0.1 K. Electric contacts are made by pressing flat copper disks against two sides of the sample which are wetted with a mixture of In-Hg (50-50). The sample is connected to an electric circuit, identical to the bolometer circuit, so the variation of the conductivity due to chopped radiation can be measured with a lock-in amplifier.



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## III.1 GENERAL INTRODUCTION

1. The interferogram-function.

In a Michelson interferometer radiation from the source is divided into two beams, one is sent to a fixed mirror, and the other is sent to another mirror which can be driven along one arm of the interferometer (Fig. 1). Assume the source to be monochromatic, emitting radiation of amplitude  $\sqrt{S}$  (intensity  $S$ ) and of wave number  $\sigma(=1/\lambda)$ . This radiation falls onto a dielectric beam divider to which can be assigned a complex reflectivity  $\hat{r}$  and transmissivity  $\hat{t}$ ; these functions are complex because of the phase shifts introduced into the optical beam by the phenomena that occur in the

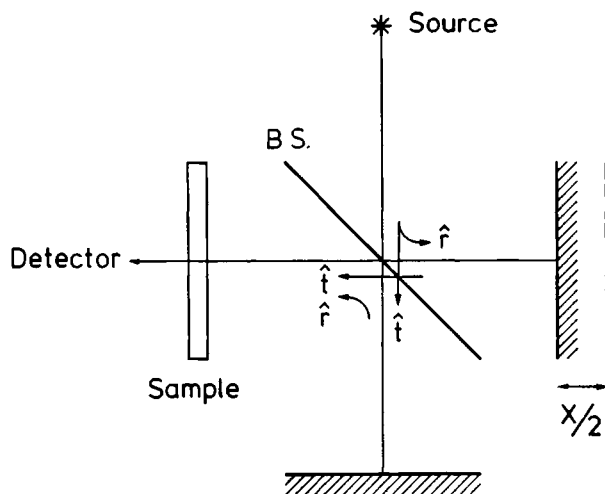


Fig. 1 : The division of complex amplitudes in a Michelson interferometer;  $\hat{r}$  and  $\hat{t}$  are the complex reflectivity and transmissivity of the dielectric film beam splitter (B.S.).

beam divider. (See e.g. Chantry<sup>(1)</sup>, chapter 3.) Therefore  $\hat{r}$  and  $\hat{t}$  can be written as

$$\begin{aligned}\hat{r} &= r \cdot e^{i\theta} \\ \hat{t} &= t \cdot e^{i\varphi} .\end{aligned}\tag{1}$$

A displacement  $x/2$  of the moving mirror in one arm of the interferometer produces a phase delay  $\Delta$  between the two beams, given by

$$\Delta = 2\pi\sigma x .\tag{2}$$

From Fig. 1 it can easily be seen that the amplitude  $\hat{a}_{\text{det}}$  of the recombined beam going to the detector, and the amplitude  $\hat{a}_{\text{source}}$  of the beam going back to the source, are given by

$$\begin{aligned}\hat{a}_{\text{det}} &= \sqrt{S}(\hat{t}\hat{t} + \hat{r}\hat{r} \cdot e^{i\Delta}) \\ \hat{a}_{\text{source}} &= \sqrt{S}(\hat{t}^2 + \hat{r}^2 \cdot e^{i\Delta}) .\end{aligned}\tag{3}$$

These equations illustrate the dissymmetry of the two beams. For the one travelling towards the detector, the beams traversing the two arms each suffer one transmission and one reflection at the beam splitter; for the other case one beam is reflected twice and the other transmitted twice. Using equation (1) this leads to

$$\begin{aligned}\hat{a}_{\text{det}} &= \sqrt{S}rt(1 + e^{i\Delta})e^{i(\theta+\varphi)} \\ \hat{a}_{\text{source}} &= \sqrt{S}(t^2 \cdot e^{i2\varphi} + r^2 \cdot e^{i2\theta} \cdot e^{i\Delta}) .\end{aligned}\tag{4}$$

The intensity of the beams is given by the square of the modulus of the complex amplitude. Therefore, one gets at the detector and at the source respectively

$$\begin{aligned}I_{\text{det}} &= 2r^2t^2S(1 + \cos\Delta) \\ I_{\text{source}} &= S[r^4 + t^4 + 2r^2t^2 \cdot \cos(2\theta - 2\varphi + \Delta)] .\end{aligned}\tag{5}$$

In the case of a non-absorbing beam splitter it can be shown (Chantry<sup>(1)</sup>, chapter 3) that

$$2\theta - 2\varphi = (2n+1)\pi \quad n = 1, 2, 3, \dots \quad (6)$$

This leads to

$$I_{\text{source}} = S(r^4 + t^4 - 2r^2 t^2 \cdot \cos \Delta) \quad (7)$$

The properties of the beam splitter can be combined in two parameters A and B given by

$$\begin{aligned} A &= r^4 + t^4 \\ B &= 2r^2 t^2 \end{aligned} \quad (8)$$

The quantity B is often called the efficiency of the beam splitter. If a sample with energy transmission-coefficient T is placed into the beam going to the detector (see Fig. 1), the intensity of the relevant beams can finally be expressed as

$$\begin{aligned} I_{\text{det}}(\Delta) &= TSB(1 + \cos \Delta) \\ I_{\text{source}}(\Delta) &= S(A - B \cdot \cos \Delta) \end{aligned} \quad (9)$$

So far the source was assumed to be monochromatic with wave number  $\sigma$  and intensity S. If the source emits a broad spectrum with distribution  $S(\sigma)$ , the total intensity  $I'(x)$  reaching the detector can be found by integrating  $I_{\text{det}}(\Delta)$  over all wave numbers  $\sigma$ . It has to be noted that the quantities A, B and T are functions of  $\sigma$  as well. The interferogram-function  $I'(x)$  for a broad spectrum is thus found to be

$$\begin{aligned} I'(x) &= \int_0^{\infty} T(\sigma) \cdot S(\sigma) \cdot B(\sigma) \cdot (1 + \cos 2\pi \sigma x) \cdot d\sigma \\ &= 2 \int_0^{\infty} Q(\sigma) \cdot (1 + \cos 2\pi \sigma x) \cdot d\sigma, \end{aligned} \quad (10)$$

where  $Q(\sigma) \equiv \frac{1}{2} T(\sigma) \cdot S(\sigma) \cdot B(\sigma)$ .

## 2. Amplitude and phase modulation.

Since the intensity of the source in the far infrared region is very weak, the voltage output of the detectors is very small. Therefore it is necessary to modulate the radiation and to use phase sensitive detection (PSD) techniques. In a Michelson interferometer this modulation can be done in two ways: 1) Amplitude Modulation (AM), where the light beam emerging from the source is chopped; 2) Phase Modulation (PM), where the optical path difference between the two interfering beams is modulated. This is achieved by a vibration with small amplitude  $a$  of one of the mirrors in the interferometer. The modulation frequency  $\omega$  for AM and PM is optimized with respect to the response time of the detector.

In the previous section the unmodulated interferogram-function  $I'(x)$  was calculated and is given by equation (10). When some form of periodic modulation is used, the interferogram-function becomes time-dependent and the intensity reaching the detector can be resolved into Fourier components. Since phase sensitive detection selects out only the fundamental Fourier-component of the signal, the ideal form of modulation is a pure sine-function of time. Obviously, the resulting interferograms and spectra will depend on which of the two modulation techniques, AM or PM, is used.

When AM is used the source intensity  $S(\sigma)$  in equation (10) becomes time-dependent. If the modulation has the ideal sine-form, the intensity  $I'_{AM}(x,t)$  reaching the detector becomes

$$I'_{AM}(x,t) = I'(x) \cdot (1 + \sin \omega t) . \quad (11)$$

After PSD, the registered interferogram-function will be proportional to

$$\begin{aligned} I'_{AM}(x) &= \frac{1}{2} I'(x) \\ &= \frac{1}{2} \int_0^{\infty} 2Q(\sigma) \cdot (1 + \cos 2\pi \sigma x) \cdot d\sigma . \end{aligned} \quad (12)$$

This proportionality-factor is determined by the detector sensitivity and the amplification of the phase sensitive detector (lock-in amplifier). If a more practical square function is used for the modulation, the factor  $\frac{1}{2}$  in equation (12) has to be replaced by  $2/\pi$ . The actual spectral information is contained in the  $x$ -dependent part of equation (12), called the interferogram.

This interferogram  $I_{AM}(x)$  is obtained from expression (12) by subtracting the constant level  $\int Q(\sigma) \cdot d\sigma$ , leading to

$$I_{AM}(x) = \frac{1}{2} \int_0^{\infty} 2Q(\sigma) \cdot \cos(2\pi\sigma x) \cdot d\sigma . \quad (13)$$

Notice that  $I_{AM}(x)$  is an even function of  $x$ !

When PM is applied the optical path difference  $x$  is modulated. If a sinusoidal modulation of amplitude  $a$  is used, in equation (10)  $x$  has to be replaced by  $x+a \cdot \sin\omega t$ . Using the Bessel series expansions of  $\cos(z \cdot \sin\theta)$  and  $\sin(z \cdot \sin\theta)$ , the cos-term in equation (10) can be written as

$$\cos[2\pi\sigma(x+a \cdot \sin\omega t)] = \sum_{m=-\infty}^{\infty} J_m(2\pi\sigma a) \cdot \cos(2\pi\sigma x + m\omega t) . \quad (14)$$

Since phase sensitive detection is used, only the component with the fundamental frequency has to be considered. The intensity  $I'_{PM}(x, t)$  of this component is given by

$$I'_{PM}(x, t) = \int_0^{\infty} 2Q(\sigma) \cdot (-2J_1(2\pi\sigma a)) \cdot \sin 2\pi\sigma x \cdot \sin\omega t \cdot d\sigma , \quad (15)$$

where the relation  $J_{-m}(z) = (-1)^m J_m(z)$  has been used. After PSD, the registered interferogram is therefore proportional to

$$I_{PM}(x) = \int_0^{\infty} 2Q(\sigma) \cdot (-2J_1(2\pi\sigma a)) \cdot \sin 2\pi\sigma x \cdot d\sigma . \quad (16)$$

Notice that this interferogram is an odd function of  $x$  and that this expression has no constant term as in equation (12).

### 3. Spectra calculated by Fourier transformation.

Expressions for the observed interferograms were given in equations (13) and (16) for the case of AM and PM respectively. An extension can be made to negative wave numbers by stating that  $Q(-\sigma) = Q(\sigma)$ . Since  $J_1(-2\pi\sigma a) = -J_1(2\pi\sigma a)$  it follows that  $-2J_1(2\pi\sigma a) \cdot Q(\sigma)$  is an odd function of  $\sigma$ . With this

extension equations (13) and (16) can be rewritten in a complex form as

$$I_{AM}(x) = \frac{1}{2} \int_{-\infty}^{\infty} Q(\sigma) \cdot e^{i2\pi\sigma x} \cdot d\sigma$$

$$I_{PM}(x) = \frac{-2}{i} \int_{-\infty}^{\infty} Q(\sigma) \cdot J_1(2\pi\sigma a) \cdot e^{i2\pi\sigma x} \cdot d\sigma .$$
(17)

From these equations it is clear that the interferograms  $I_{AM}(x)$  and  $I_{PM}(x)$  have a Fourier-type of relationship with the spectral functions  $\frac{1}{2} Q(\sigma)$  and  $\frac{-2}{i} Q(\sigma) \cdot J_1(2\pi\sigma a)$  respectively. The latter two functions can be recovered from  $I_{AM}(x)$  and  $I_{PM}(x)$  by an inverse complex Fourier transform. Since in practise a complex Fourier transform yields a complex spectrum, one usually calculates the modulus of this spectrum, i.e.

$$P(\sigma) = \left| \int_{-\infty}^{\infty} I(x) \cdot e^{-2\pi i \sigma x} \cdot dx \right| ,$$
(18)

yielding

$$P_{AM}(\sigma) = \frac{1}{2} Q(\sigma) \quad \text{and}$$

$$P_{PM}(\sigma) = 2 |J_1(2\pi\sigma a)| \cdot Q(\sigma) .$$

From these considerations, some advantages of PM are immediately evident:

1. In order to obtain the spectral information from the interferogram-function  $I'_{AM}(x)$  (equation (12)), the constant term in this expression has to be subtracted. Fluctuations in the source intensity and detector sensitivity affect this constant term and therefore can falsify the spectrum. Since the expression for  $I_{PM}(x)$  (equation (16)) contains no constant term, greater immunity from fluctuations in source intensity and detector sensitivity is obtained with PM.

2. The ratio of  $P_{PM}(\sigma)$  and  $P_{AM}(\sigma)$  is  $4 |J_1(2\pi\sigma a)|$ . This factor is wave number dependent and is  $> 1$  for wave numbers between  $0.08/a$  and  $0.51/a$ . It reaches zero at  $\sigma_0 = 0.611/a$  and has a maximum of 2.3 at  $\sigma_{max} = 0.29/a$ . This means that in case of PM the signal-to-noise ratio exceeds that of AM over 75% of the range between 0 and  $\sigma_0$  and is twice as high near  $\sigma_{max}$ . A physical

way of looking at this is that there is no chopper to interrupt the optical beam, so that the detector views the source for a greater length of time with PM, thus leading to an increase in the signal-to-noise ratio.

3. The wave number dependent multiplicative factor  $4|J_1(2\pi\sigma a)|$  can be used to act as a non-absorptive filter. The Bessel function falls to zero when  $\sigma = 0.611/a$ . By an appropriate choice of the modulation amplitude  $a$ , a certain spectral region can be suppressed. Thus in combination with suited low-pass transmission filters an effective suppression of radiation with high wave numbers can be achieved.

A more detailed description of AM and PM was given by Chamberlain<sup>(2,3,4)</sup> and Chantry<sup>(1)</sup>.

#### 4. Effects due to sampling of the interferogram over a finite length.

The optical path difference  $x$  is usually assumed to be a continuous variable ranging from  $-\infty$  to  $+\infty$ . Actually, however, interferograms are recorded in discrete steps  $\Delta x$ , over a finite range. If  $N$  datapoints are recorded symmetrically around zero path difference, i.e.  $x$  ranging from  $-L$  to  $+L$  with  $L = (N/2) \cdot \Delta x$ , then expression (18) for the modulus spectrum has to be modified into

$$P(\sigma) = \Delta x \cdot \sum_{n=-N/2}^{+N/2} I(n\Delta x) \cdot e^{-2\pi i \sigma n \Delta x} \quad (19)$$

This expression relates  $P(\sigma)$  uniquely with  $I(n\Delta x)$  for a limited range of values for  $\sigma$  only. This can be seen by considering the source to be monochromatic with wave number  $\sigma_0 = (1/\lambda_0)$ . According to equation (9), this results in a cosine-like interferogram. To determine the value of  $\sigma_0$  from the interferogram, the sampling theorem states that at least two points have to be sampled within one wave length. Therefore, this leads, for a given value of  $\Delta x$ , to a limit for the highest wave number  $\sigma_{\max}$ , given by  $\sigma_{\max} = 1/(2\Delta x)$ . However, it is obvious that an infinite number of waves with higher wave numbers will also fit through the points in the range of observation. Therefore, the information contained in the interferogram can be attributed to higher wave numbers as well; this phenomenon is called "aliasing" (Chantry<sup>(1)</sup>, chapter 3). If low pass filters are used, which block out the radiation with wave numbers above  $\sigma_{\max}$ , using expression (19) the original spectrum can be



recovered uniquely from the recorded interferogram.

In addition, the limitation of recording the interferogram over a finite range of  $x$ -values,  $-L \leq x \leq L$ , also confines the spectral resolution. It can easily be shown (Chantry<sup>(1)</sup>, chapter 3), that the power spectrum of a truncated interferogram of a monochromatic line is proportional to

$$\frac{\sin 2\pi(\sigma_0 - \sigma)L}{2\pi(\sigma_0 - \sigma)L} \quad (20)$$

This function has a half-width of approximately  $1/(2L)$ , and has large and slowly decaying sidelobes. If the resolution is taken as the minimum wave number distance at which another line of equal intensity can just be observed in the transformed discrete spectrum, then one gets a resolution of  $1/(2L)$ . Therefore, spectra with a higher resolution can be obtained simply by moving the mirror over a longer distance, parallel to its original position. In practice, the resolution is restricted by the signal-to-noise-ratio, the finite aperture of the source and detector, and by tilt adjustments (i.e. the image of one mirror in the beam splitter is not perfectly parallel to the surface of the other).

The sidelobes mentioned above can easily be mistaken for real spectral features. If the recorded interferogram is multiplied with a so-called apodization-function, these sidelobes can be suppressed effectively. Usually a cosine squared function (which is 1 at  $x = 0$  and 0 at  $x = \pm L$ ) is chosen; this is also done in our computer program, described in the Appendix. It should be noted however, that this procedure broadens the central line, degrading the resolution by a factor 1.5.

An example of an interferogram of a monochromatic line, recorded with amplitude modulation, is shown in figure 2a. This line originated from a home-built grating monochromator, described in refs. 5 and 6. Figure 2b shows the corresponding power spectrum, calculated with equation (19). From figure 2a it can be seen that the interference pattern decays within the interferogram length; therefore figure 2b represents the actual lineshape. In fact, the lineshape is determined by the width of the exit-slit of the grating monochromator. From the interferogram, the wavelength  $\lambda_0$  is seen to have a value of 0.370 mm, while the central position  $\sigma_0$  of the line in the spectrum is situated at  $27 \text{ cm}^{-1}$ . So for these values the relation  $\sigma_0 = 1/\lambda_0$  holds.

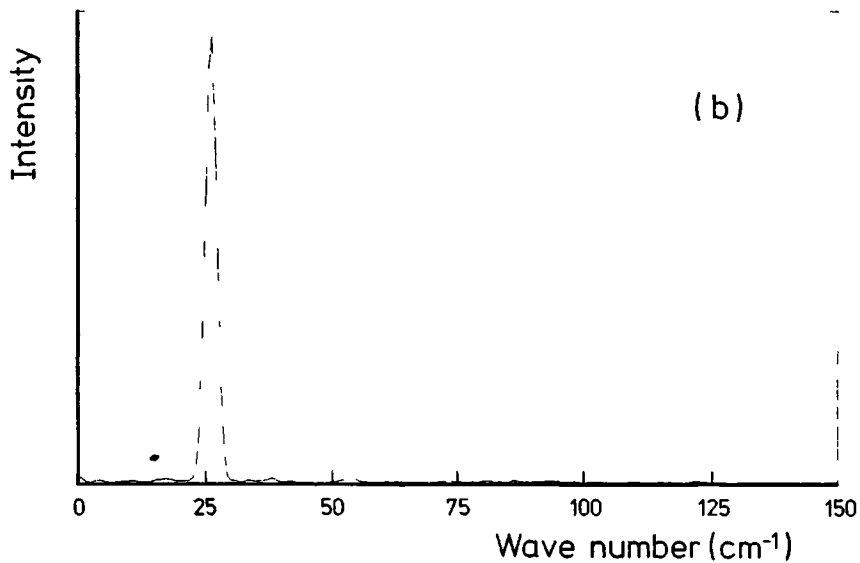
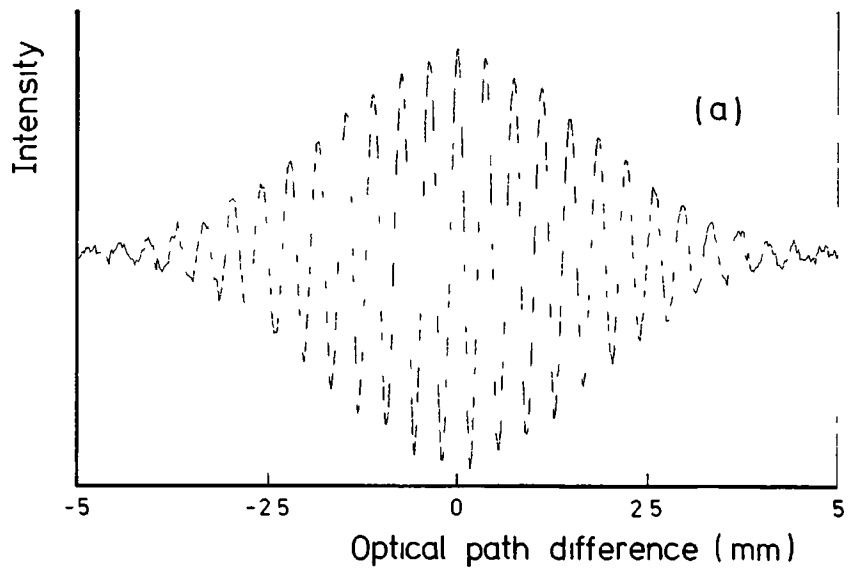


Fig. 2 : Figure (a) and (b) show respectively the interferogram and spectrum of a monochromatic line, originating from a grating monochromator.

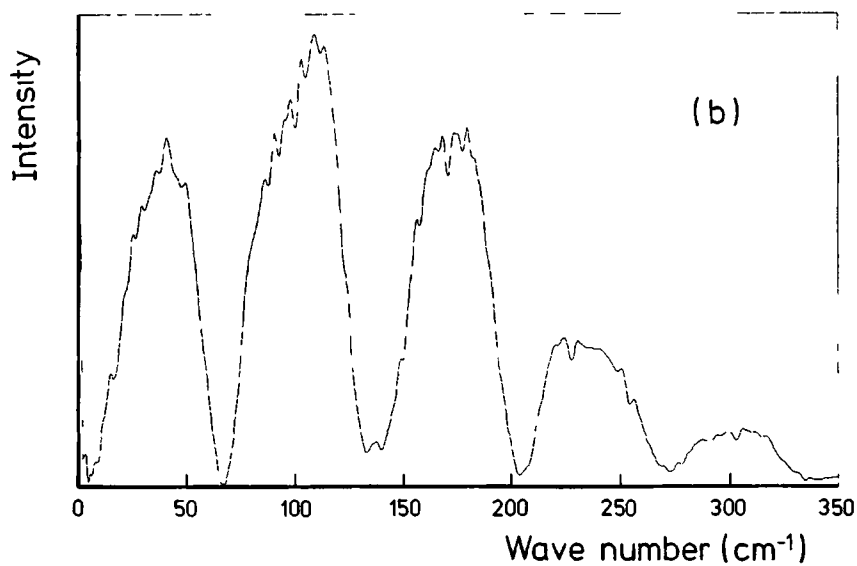
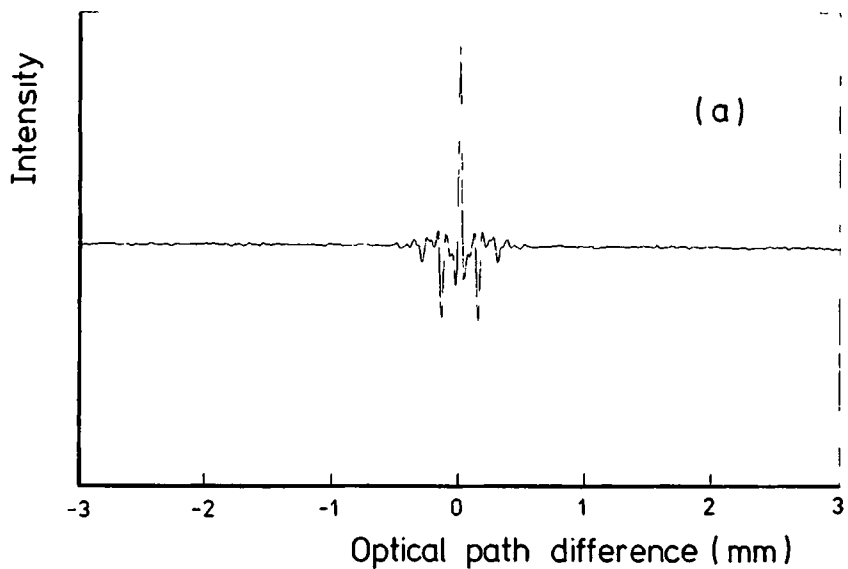


Fig. 3 : Figure (a) and (b) show respectively the interferogram and spectrum of a continuous broad band spectrum. The structure is determined by the efficiency of the 50  $\mu\text{m}$  Mylar beamsplitter.

In the previous section it was demonstrated that the spectral output was determined by the properties of the thin film beam splitter as well. The disadvantage of these thin films is that due to standing wave interference effects, caused by multiple reflections within the thin transparent films, the beam splitter efficiency  $B(\sigma)$  is a periodic archlike function of wave number, reaching zero at multiple values of  $1/(2nd\cos\theta')$ . Here  $d$  is the thickness of the film,  $n$  the refractive index and  $\theta'$  the angle of the refracted beam inside the film. The function  $B(\sigma)$  was first introduced in equation (8); its exact form has been calculated by several authors<sup>(7-10)</sup>. An example of the influence of a beam splitter is shown in figure 3a and 3b, where a Mylar beam splitter of 50  $\mu\text{m}$  thickness was used, with an angle of incidence for the broad band radiation of  $45^\circ$ . Using  $n = 1.72$ <sup>(9)</sup> the wave numbers where  $B(\sigma)$  reaches 0, are found to be multiple values of  $64\text{ cm}^{-1}$ , in accordance with figure 3b.

Interferograms such as in figure 2a and 3a, which are recorded symmetrically around zero path difference, are called double-sided interferograms. In principle, if the interferogram is exactly symmetrically sampled, it is sufficient to record one side only (single-sided interferograms). Actually, often the sampling comb is not positioned so as to record an ordinate at exactly zero path difference. It can easily be shown (Chantry<sup>(1)</sup>, chapter 3) that this does not affect the modulus spectrum of a double-sided interferogram. In case of a single-sided interferogram symmetry must be restored by numerical computation<sup>(11,12)</sup>. Single-sided interferograms (not used in the present investigations) are treated in details in textbooks (see Bell<sup>(13)</sup> or Vanasse and Sakai<sup>(14)</sup>).

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ABSTRACT

In far-infrared Fourier spectroscopy with a Michelson interferometer deviations in the measured transmission spectra of samples with non-negligible reflection can be observed. This effect is investigated both theoretically and experimentally. It is shown that the spectral element at a certain wave number contains contributions of subharmonics of that wave number.

1. Introduction.

One of the most widely used instruments in far-infrared spectroscopy is the Michelson interferometer with thin film beamsplitter. An extensive review of the Fourier spectroscopic technique on which this instrument is based has been given by Chantry<sup>(1)</sup>. In order to determine the transmission spectrum of some sample one usually performs two measurements: 1) A sample interferogram with a sample in the light-path between interferometer and detector, and: 2) A background interferogram without the sample. These interferograms are Fourier transformed, and the ratio of the two obtained spectra then yields the desired transmission spectrum.

However, it should be noted that a non-negligible amount of radiation can be reflected back into the interferometer system by a reflecting sample. This can lead to significant distortions in the spectra. This effect seems to have escaped notice until now. It is the purpose of this paper to present results of a theoretical and experimental investigation into the effect of this reflected radiation on the measured transmission spectrum.

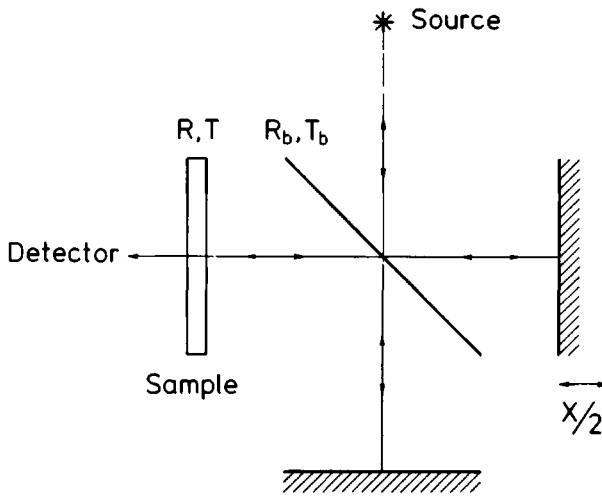


Fig. 1 : Schematic diagram of a Michelson interferometer in a configuration for transmission measurements. The energy reflection and transmission coefficients of the beamsplitter are  $R_b$  and  $T_b$  respectively, those of the sample are  $P$  and  $T$ .

## 2. Theory.

Fig. 1 shows schematically a Michelson interferometer with an incident beam of radiation. There are two emergent beams, one towards the sample and detector, the other towards the source. In the following calculations we neglect the effect of absorption in the beamsplitter, which is a good approximation for the first lobe of the interference pattern of a Mylar beamsplitter<sup>(2)</sup>.

We assume the source to be monochromatic with wave number  $\sigma$  and intensity  $S$ , and the optical path difference between the two arms of the interferometer to be  $x$ . Then the phase difference between the beams recombining at the beamsplitter will be  $\Delta = 2\pi\sigma x$ , the energy  $I_0$  arriving at the sample will be given by

$$I_0(\Delta) = SB(1 + \cos\Delta) , \quad (1)$$

and the energy  $I'_0$  returning to the source by

$$I'_0(\Delta) = S(A - B \cdot \cos \Delta) . \quad (2)$$

(See for instance: Loewenstein and Engelsrath<sup>(3)</sup>.) The beamsplitter properties are represented by the quantities  $A = R_b^2 + T_b^2$  and  $B = 2R_b T_b$ , where  $R_b$  and  $T_b$  denote the energy reflection and transmission coefficients of the beamsplitter respectively. For a non-absorbing beamsplitter we have  $R_b + T_b = 1$  and  $A + B = 1$ . The quantity  $B$  is often called the efficiency of the beamsplitter.

The sample will transmit a fraction  $T$  of the incident radiation towards the detector, and will reflect a fraction  $R$  back into the interferometer ( $R < 1$ ). To simplify the calculations it will be assumed that the reflected radiation has a random phase relation with the original beam, so that no interference will occur and one can simply add the intensities. The sample can then be considered as an extra radiation source with intensity  $RI_0(\Delta)$ . According to Eq. (2) a fraction with energy

$$I_1(\Delta) = RI_0(\Delta)(A - B \cdot \cos \Delta)$$

will return to the sample, and will again be partly reflected, etc. The total energy  $I$  reaching the detector, after transmission through the sample, will therefore be

$$I(\Delta) = TSB(1 + \cos \Delta) \sum_{n=0}^{\infty} R^n (A - B \cdot \cos \Delta)^n . \quad (3)$$

The term in the infinite sum is  $< 1$  (because  $R < 1$  and  $A + B = 1$ ), so the geometrical series can be summed. Introducing the quantity  $C$ , defined as

$$C = \frac{RB}{1 - RA} , \quad (4)$$

the summation of Eq. (3) yields:

$$I(\Delta) = \frac{TS}{R} \left[ 1 - \frac{1 - C}{1 + C \cdot \cos \Delta} \right] . \quad (5)$$



It follows from Eq. (4) that  $0 \leq C < 1$ . In order to perform the Fourier transform it is convenient to write Eq. (5) in the form of a Fourier series. It is shown in the Appendix that the function

$$f(\Delta) = (1 + C \cdot \cos \Delta)^{-1}$$

can be represented by a Fourier series

$$f(\Delta) = \frac{a_0}{2} + \sum_{n=1}^{\infty} a_n \cdot \cos(n\Delta)$$

with coefficients  $a_n$  given by

$$a_n = \frac{2}{\sqrt{1 - C^2}} \left( \frac{\sqrt{1 - C^2} - 1}{C} \right)^n. \quad (6)$$

Defining the coefficients  $B_n$  as:

$$B_0 = \frac{1}{R} \left[ 1 - (1 - C) \frac{a_0}{2} \right] \quad (7.a)$$

$$B_n = \frac{C - 1}{nR} \cdot a_n \quad \text{for } n = 1, 2, 3, \dots, \quad (7.b)$$

the Fourier series representation of  $I(\Delta)$  can be expressed as:

$$I(\Delta) = TS \left[ B_0 + \sum_{n=1}^{\infty} B_n \cdot n \cdot \cos(n\Delta) \right]. \quad (8)$$

Fig. 2 shows the calculated  $B_0$ ,  $B_1$ ,  $B_2$  and  $B_3$  as functions of  $R$  for various values of the beamsplitter efficiency  $B$ . It can easily be shown that for  $R \rightarrow 0$  the values  $B_0 \rightarrow B$ ,  $B_1 \rightarrow B$  and  $B_n \rightarrow 0$  for  $n \geq 2$ , so that in this limit Eq. (8) reduces to the well known form of Eq. (1)

$$\lim_{R \rightarrow 0} I(\Delta) = TSB(1 + \cos \Delta).$$

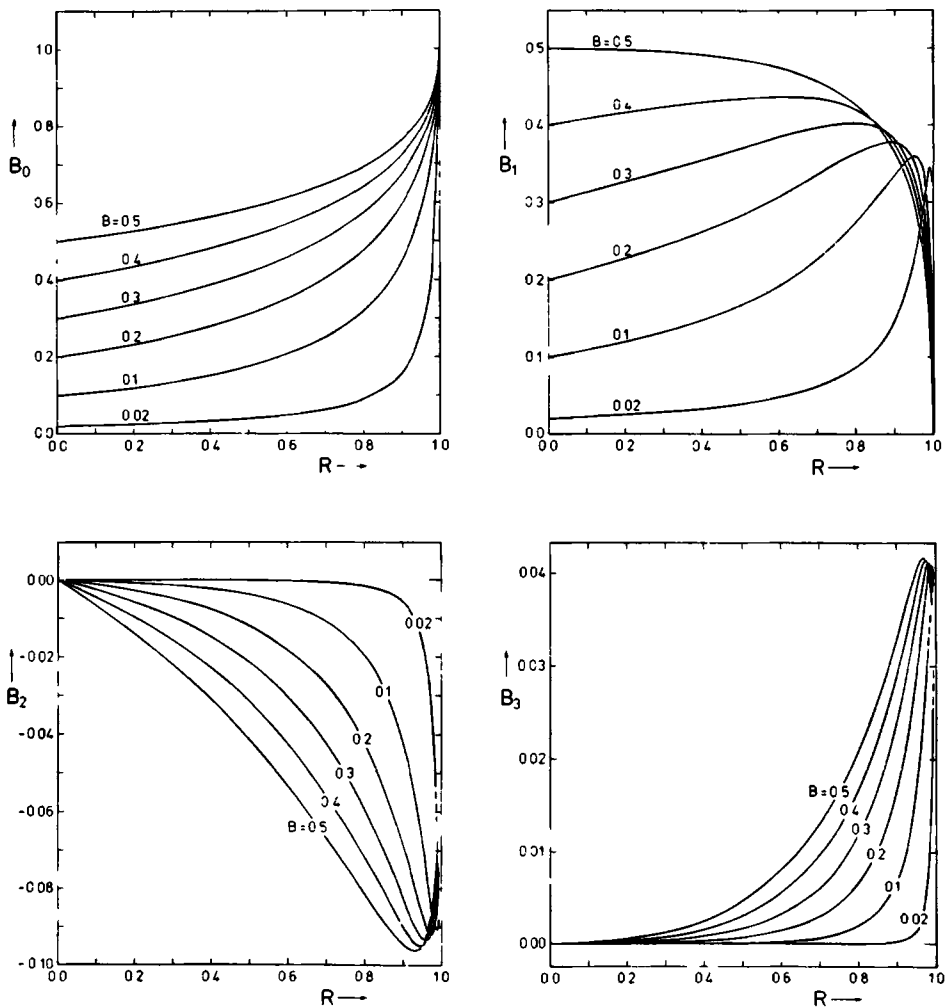


Fig. 2 : The calculated values of the quantities  $B_0$ ,  $B_1$ ,  $B_2$  and  $B_3$  as functions of  $R$  for various values of the beamsplitter efficiency  $B$  (non-absorbing beamsplitter).

It is obvious from Eq. (8) that the interferogram-function  $I(\Delta)$  contains higher harmonics of the basic frequency due to the reflected radiation.

So far the source was assumed to be monochromatic with wavenumber  $\sigma$  and intensity  $S$ . If the source emits a broad spectrum  $S(\sigma)$ , the total energy  $I'(x)$  reaching the detector can be found by integrating  $I(\Delta)$  over all wavenumbers  $\sigma$ . In doing so one has to take into account that the quantities  $A, B, T, R, C$  and  $B_n$  may also be functions of  $\sigma$ . The interferogram-function  $I'(x)$  for a broad spectrum is thus found to be

$$I'(x) = \int_0^{\infty} T(\sigma) \cdot S(\sigma) \cdot B_0(\sigma) \cdot d\sigma \quad (9)$$

$$+ \int_0^{\infty} \sum_{n=1}^{\infty} T(\sigma) \cdot S(\sigma) \cdot B_n(\sigma) \cdot n \cdot \cos(2\pi n \sigma x) \cdot d\sigma .$$

The first integral in Eq. (9) is independent of  $x$ , and corresponds to the average of  $I'(x)$  over all  $x$ . The second integral in Eq. (9) is the proper interferogram  $I(x)$ , which, by changing the variable  $\sigma$  to  $\sigma/n$ , can be written as

$$I(x) = \int_0^{\infty} \sum_{n=1}^{\infty} T\left(\frac{\sigma}{n}\right) \cdot S\left(\frac{\sigma}{n}\right) \cdot B_n\left(\frac{\sigma}{n}\right) \cdot \cos(2\pi \sigma x) \cdot d\sigma . \quad (10)$$

In most experiments the intensity of the source is chopped and the detector provides a signal proportional to  $I'(x)$ . After phase-sensitive detection and subtraction of the average level one obtains an interferogram proportional to  $I(x)$ . The spectrum one is interested in is then calculated as the Fourier transform of this interferogram. One sees immediately from Eq. (10) that this spectrum  $P(\sigma)$ , omitting multiplicative constants, is given by:

$$P(\sigma) = \sum_{n=1}^{\infty} T\left(\frac{\sigma}{n}\right) \cdot S\left(\frac{\sigma}{n}\right) \cdot B_n\left(\frac{\sigma}{n}\right) . \quad (11)$$

This means that the spectral element with wavenumber  $\sigma$  contains not only information about the sample, source and instrument for wavenumber  $\sigma$ , but

also for a series of subharmonics of  $\sigma$ .

For small values of  $R$  Eq. (11) reduces to the well known undistorted spectrum:

$$\lim_{R \rightarrow 0} P(\sigma) = T(\sigma) \cdot S(\sigma) \cdot B(\sigma) .$$

The background spectrum (without the sample) is simply  $S(\sigma) \cdot B(\sigma)$ , so that for  $R \rightarrow 0$  the transmission spectrum, calculated as the ratio of the sample and background spectrum, is indeed equal to  $T(\sigma)$ . However, for non-negligible values of  $R$  one does not obtain the proper transmission spectrum in this simple way, and the corrections discussed above have to be taken into account.

### 3. Experiment.

In order to test the analysis of section 2 we performed some experiments in which the transmission spectra of samples with high reflection were measured and compared with the known transmission. A modular FIR interferometer<sup>(4)</sup>, manufactured by Grubb Parsons Company, was employed with a 12.5  $\mu\text{m}$  Mylar beamsplitter. One of the modifications allowed the use of both the standard broadband source (i.e. a mercury lamp) or an external monochromatic source (i.e. a FIR-laser or a monochromator). A TPX lens focussed the radiation emerging from the interferometer onto the entrance hole of a lightpipe which contained the sample-holder. The reflected radiation re-entered the interferometer along the reversed path, while the transmitted power was detected at the exit of the lightpipe.

#### a. Monochromatic source.

The properties of an interferometer are most clearly studied by employing a monochromatic source. In our experiments this was an optically pumped FIR laser, operating at the 118.8  $\mu\text{m}$  ( $\sigma_0 = 84.2 \text{ cm}^{-1}$ ) transition of methanol<sup>(5)</sup>. The radiation was coupled into the interferometer via a lightpipe and collimated by a TPX lens; the emerging power was measured with a Molelectron P3 pyro-electric detector. A 750 lines/inch nickel mesh (manufactured by Buckbee and Mears Company, Minnesota) was used as reflecting sample; the measured transmission coefficient at  $84.2 \text{ cm}^{-1}$  was  $T = 0.15$ .

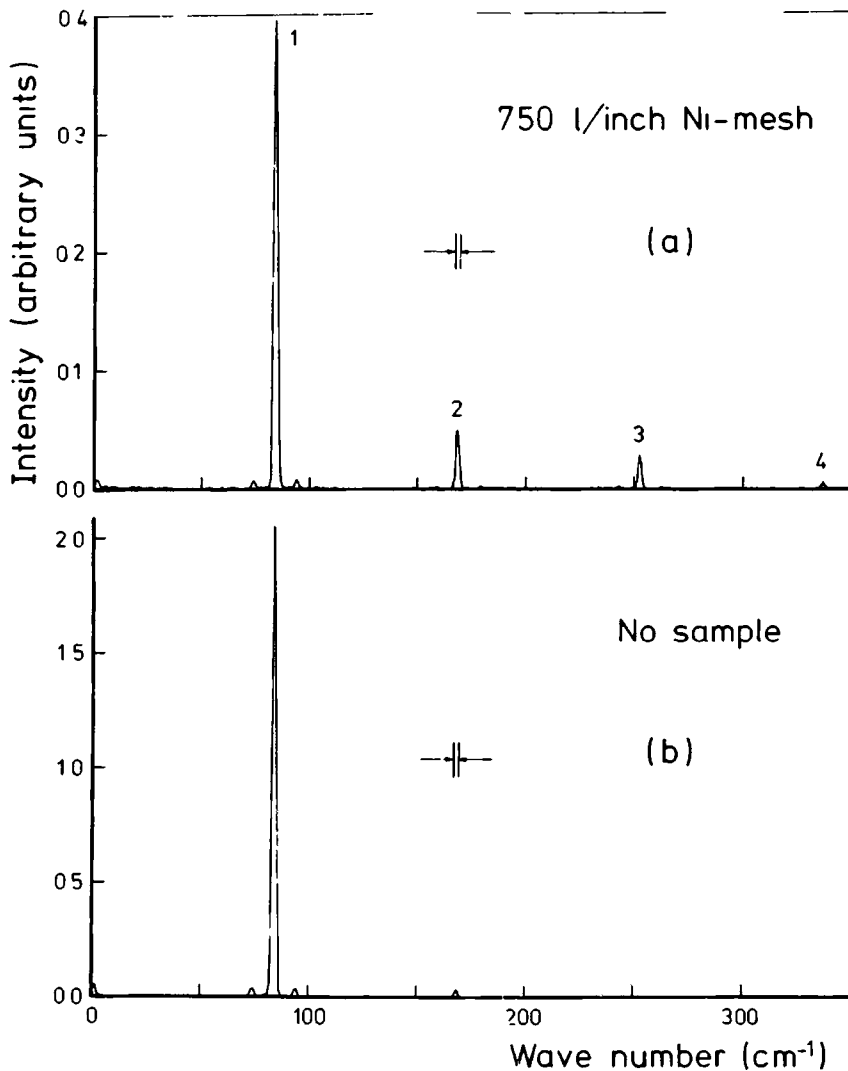


Fig. 3 : The transmission spectrum of a 750 lines/inch nickel mesh, showing the effect on the transformed spectrum of the reflected radiation (1). The source was monochromatic with radiation wave number  $\sigma_0 = 84.2 \text{ cm}^{-1}$ ; the background spectrum is shown as curve (b). The transmission of the sample at  $84.2 \text{ cm}^{-1}$  is  $T = 0.15$ .

The measured background interferogram was an almost perfect cosine of which the Fourier transform yields the monochromatic peak at  $\sigma_0 = 84.2 \text{ cm}^{-1}$  shown in Fig. 3b. A very small harmonic component at  $2\sigma_0$  can be seen which may be due to a reflection at the detector or to a possible non-linearity of the detector sensitivity. The small side-peaks at  $10 \text{ cm}^{-1}$  to the left and right of the main peak are caused by a small periodical pitch error in the micrometer screw of the mirror drive of the interferometer<sup>(6)</sup>.

The measured transmission spectrum (see Fig. 3a) of the sample shows very clearly the peaks at harmonics of  $\sigma_0$  as predicted by the theory of section 2. As our computer program calculated the modulus of the complex transformed spectrum, the negative sign of the peaks at even multiples of  $\sigma_0$  does not show up in Fig. 3a. The intensities of the peaks relative to the background peak are given in Table 1. We also calculated the relative peak intensities with the formula  $T(\sigma_0) \cdot B_n(\sigma_0) / B(\sigma_0)$  given by the theory; the results are also given in Table 1, where the negative signs for the even peaks have been omitted. In the calculations a loss of 5% of the radiation was assumed so that the effective reflection coefficient of the sample was taken as  $R = 0.8$ .

In calculating the efficiency of the beamsplitter the polarization properties have to be taken into account<sup>(2,3)</sup>. The originally polarized radiation of the laser proved to be totally depolarized by entering the interferometer through a lightpipe and collimator lens. The calculations were performed separately for the parallel and perpendicularly polarized components of the incident radiation, and the results were averaged. For a non-absorbing  $12.5 \text{ }\mu\text{m}$  Mylar beamsplitter and  $45^\circ$  angle of incidence, the calculated efficiency of the two components was found to be  $B_{\parallel} = 0.11$  and  $B_{\perp} = 0.46$  respectively.

As can be seen from Table 1 the agreement between theory and experiment is rather good. It is interesting to note that the efficiency of the beamsplitter has a minimum at  $250 \text{ cm}^{-1}$  so that without the reflection effect discussed in this paper there could not be a noticeable intensity in the region where peak no. 3 is located.

#### b. Broadband source.

In order to investigate the influence of the reflection effect on broadband spectra, we measured the transmission of polished brass disks

Table 1 : The experimentally and theoretically determined relative peak intensities in the transmission spectrum of a 750 lines/inch nickel mesh with a monochromatic source of wave number  $84.2 \text{ cm}^{-1}$ .

peak no.	1	2	3	4
Experiment	0.19	0.024	0.012	0.002
Theory	0.19	0.026	0.006	0.002

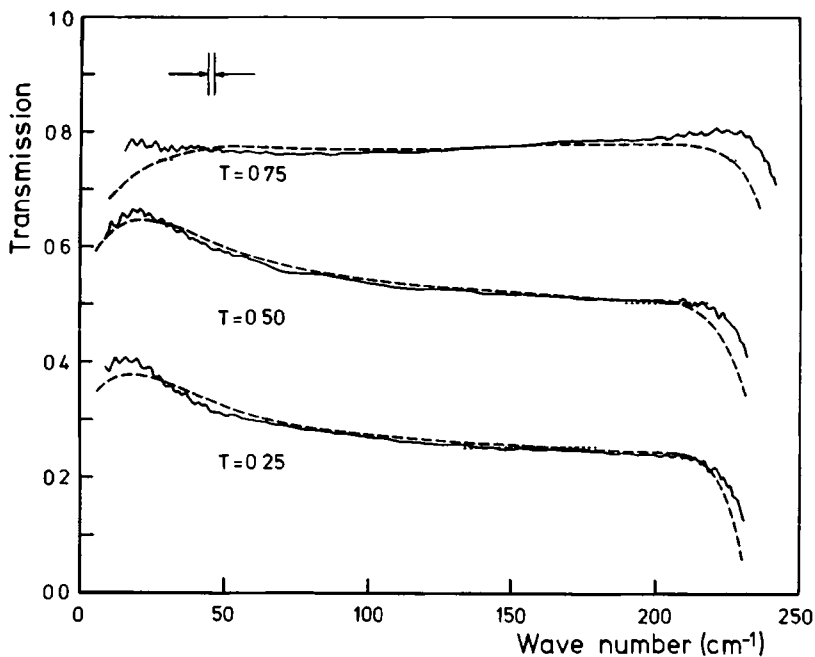


Fig. 4 : The measured (solid curves) and calculated (dashed curves) transmission spectra of polished brass disks from which sectors of  $90^\circ$ ,  $180^\circ$  and  $270^\circ$  were removed. The deviation due to the reflection effect from the ideal constant transmission coefficients  $T = 0.25$ ,  $0.50$  and  $0.75$  respectively is clearly demonstrated.

from which sectors of  $90^\circ$ ,  $180^\circ$ , and  $270^\circ$  were removed. The broadband source was a medium pressure mercury lamp (Philips HPK 125 W). An IR50 Golay cell was used as detector. Without the reflection effect under study, the expected transmission spectra should have a constant level at  $T = 0.25$ ,  $0.50$ , and  $0.75$  respectively. But, as can be seen in Fig. 4, the experimental results exhibit a distinct deviation from this ideal behaviour. Only the spectrum with the lowest reflection ( $T = 0.75$ ) approximates the real transmission spectrum almost up to the regions near  $0 \text{ cm}^{-1}$  and  $250 \text{ cm}^{-1}$ , where the ratio of sample and background spectrum becomes inaccurate because of the minima in the beamsplitter efficiency. However, the measured transmission spectra conform very well to the calculated spectra (dashed curves in Fig. 4) if the reflection effect is included. In these calculations effective reflection coefficients were chosen by assuming a loss of 5% of the radiation, yielding  $R = 0.70$ ,  $0.45$ , and  $0.20$  respectively. The beamsplitter efficiency  $B(\sigma)$  was calculated as in Ref. 2, ignoring absorption, and averaging over all polarization directions. The combination of source-spectrum and transmission of the present filters and windows was approximated by an analytical function  $S(\sigma)$ , which fitted the experimental curve within 10%.

As the spectral level at wavenumber  $\sigma$  is mainly determined by the first two terms of the infinite series of Eq. (11), i.e.  $T(\sigma) \cdot S(\sigma) \cdot B_1(\sigma)$  and  $T(\frac{1}{2}\sigma) \cdot S(\frac{1}{2}\sigma) \cdot B_2(\frac{1}{2}\sigma)$ , the main result of the reflection effect will be an increase of the level for low wavenumbers, and a decrease near  $250 \text{ cm}^{-1}$ . This is due to the maximum in the beamsplitter efficiency at  $125 \text{ cm}^{-1}$ , and the minima at  $0 \text{ cm}^{-1}$  and  $250 \text{ cm}^{-1}$ .

#### 4. Conclusions.

We have shown theoretically and experimentally that due to reflections from the sample the measured transmission spectrum can exhibit a substantial deviation from the real spectrum if a Michelson interferometer is employed. The best way to avoid this effect is to arrange the optical set-up in such a way that the reflected radiation is prevented to re-enter the interferometer.



## Appendix.

In this appendix the calculation will be given of the coefficients of the Fourier series representation of the function:

$$f(\Delta) = (1 + C \cdot \cos \Delta)^{-1}, \quad (\text{A1})$$

where  $0 \leq C < 1$ . This function  $f(\Delta)$  is a bounded periodic function of period  $2\pi$  (i.e.  $f(\Delta + 2\pi) = f(\Delta)$ ), and satisfies the Dirichlet conditions:

a) In any period  $f(\Delta)$  is continuous;

b) In any period  $f(\Delta)$  has only a finite number of maxima and minima.

In addition,  $f(\Delta)$  is even (i.e.  $f(-\Delta) = f(\Delta)$ ). Then  $f(\Delta)$  may be represented by a Fourier series

$$f(\Delta) = \frac{a_0}{2} + \sum_{n=1}^{\infty} a_n \cdot \cos(n\Delta) \quad (\text{A2})$$

where

$$a_n = \frac{1}{\pi} \int_{-\pi}^{\pi} f(\Delta) \cdot \cos(n\Delta) \cdot d\Delta. \quad (\text{A3})$$

Writing the Euler definition for the cosine, and using the fact that  $f(\Delta)$  is even one obtains the complex form of  $a_n$ :

$$a_n = \frac{2}{\pi} \int_{-\pi}^{\pi} \frac{e^{in\Delta}}{2 + C(e^{i\Delta} + e^{-i\Delta})} d\Delta.$$

Substituting  $z = e^{i\Delta}$  and integrating along the unity circle yields:

$$a_n = \frac{2}{\pi} \oint \frac{z^n}{2 + C(z + z^{-1})} \cdot \frac{dz}{iz}. \quad (\text{A4})$$

Introducing the quantities E and F, defined as:

$$E = \left(\frac{-1}{C}\right)(1 + \sqrt{1 - C^2}) \quad (\text{A5a})$$

$$F = \left(\frac{-1}{C}\right)(1 - \sqrt{1 - C^2}) \quad (\text{A5b})$$

Eq. (A4) can be written as:

$$a_n = \frac{4}{C} \cdot \frac{1}{2\pi i} \oint \frac{z^n}{(z - E)(z - F)} dz . \quad (\text{A6})$$

Since  $0 \leq C < 1$  we have  $E < -1$  and  $-1 < F \leq 0$ , so that  $F$  lies inside the unity circle and  $E$  does not. The residu theorem is stated as

$$\frac{1}{2\pi i} \oint \frac{g(z)}{z - a} dz = g(a)$$

if  $a$  lies inside the closed integration path, and  $g(z)$  does not have any singularities inside this path. This can be applied to Eq. (A6) by substituting  $a = F$  and

$$g(z) = \frac{z^n}{z - E} ,$$

yielding:

$$a_n = \frac{4}{C} \cdot \frac{F^n}{F - E}$$

or, using the definitions for  $E$  and  $F$ :

$$a_n = \frac{2}{\sqrt{1 - C^2}} \left( \frac{\sqrt{1 - C^2} - 1}{C} \right)^n . \quad (\text{A7})$$

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#### IV.1 INTRODUCTION TO PHOTOTHERMAL IONIZATION SPECTROSCOPY (PTIS)

##### 1. General introduction.

In the last decade extensive investigations into the ultrapurification of germanium have taken place, mainly with the purpose of fabrication of nuclear radiation detectors<sup>(1,2,3)</sup>.

As criterion for the purity of germanium is currently used the total concentration of shallow donors and acceptors (i.e. elements like As, Sb, P, Li, Al, B, etc.). One of the difficulties which were met for many years was that while the net concentration of the residual donors and acceptors in the Ge crystals could be measured, one had very little information about the chemical identities of these impurities. However, a knowledge of the chemical identities is obligatory for the improvement in the technology of preparation of ultrapure germanium.

Here a new technique, developed by Russian scientists<sup>(4,5,6)</sup>, came to the rescue, which could be used to detect and identify impurities in germanium in concentrations down to the  $10^9 \text{ cm}^{-3}$  range and possibly even lower. This technique of photothermal ionization spectroscopy (PTIS) is based on two experimentally established facts: 1) Under very definite conditions the spectrum of the extrinsic photoconductivity is a line spectrum which reflects the energy level structure of the impurities; 2) The magnitude of the photoresponse is independent of the concentrations of the impurities down to very low concentrations. A recent review of the PTIS technique was given by Kogan and Lifshits<sup>(7)</sup>. Since 1970 many experimentalists have used PTIS for the investigation of ultra-pure Ge. (See for instance Refs. 3, 8-13.)

In our laboratory we have performed measurements on ultra-pure germanium samples obtained from Dr. H.J.A. van Dijk of Philips Research Laboratories in Eindhoven.

Before explaining the PTIS technique a brief review will be given of the essential details of the theory of shallow donors and acceptors in germanium.

## 2. Theoretical considerations.

Kittel and Mitchell<sup>(14)</sup> and Kohn and Luttinger<sup>(15)</sup> presented the first comprehensive theoretical model for shallow donors in germanium; the so-called "effective mass approximation". The model is based on the following qualitative picture: The group V donor has five electrons outside of closed shells. Four of these complete the bonds with neighbouring Ge atoms. The fifth finds itself in the Coulomb field of the singly-charged impurity ion.

A series of quantum states can be assigned to this additional electron similar to those of the single electron in the hydrogen atom, but with some important modifications. First, the "hydrogenic" impurity is not in a vacuum but in a lattice with dielectric constant,  $\epsilon$ , so that all binding-energies emerging from the hydrogen atom are reduced by a factor  $\epsilon^2$  and the linear dimensions of the wave functions are increased by a factor  $\epsilon$ . Furthermore, the electric charge moving in the field of the impurity ion does not have the mass  $m_0$  of the free electron but has an effective mass  $m^*$ , which is anisotropic. The dielectric constant and the reduced mass together result in approximately 1000 times smaller energies than in the hydrogen atom. The wave functions of the excited states extend over thousands of crystal cells, but are zero in the central cell occupied by the impurity. The corresponding energy levels are therefore almost insensitive to the chemical nature of the impurity atoms, and the "effective mass theory" provides a very accurate description of the excited levels.

The differences that make the identification of the individual impurities possible are due to departures of the lowest levels from the effective mass approximation. The ground state wave function has a maximum in the crystal central cell. Therefore the ground state energy is sensitive to the form of the crystal field in the central cell and depends on the chemical nature of the impurity. This "central cell correction" to the effective mass approximation causes a "chemical shift" of the ground state level, characteristic for the impurity in question. Theoretical calculations were performed by Philips<sup>(16)</sup>, while Reuszer and Fisher<sup>(17)</sup> deduced the chemical shifts for various donors from measured absorption spectra.

Most of the considerations above apply mutatis mutandis also to acceptors in Ge. On the other hand, acceptor states represent a rather different situation, because of the special structure of the top of the valence band. The acceptor problem was first formulated by Kohn<sup>(18)</sup>, while numerical calculations were performed by Schechter<sup>(19)</sup>, Mendelson and James<sup>(20)</sup> and

more recently by Baldereschi and Lipari<sup>(21,22)</sup>. A very extensive review of the theory of acceptors and donors in Ge was given by Bassani et al.<sup>(23)</sup>.

The chemical shifts of the ground states of several shallow acceptors in germanium were determined by Jones and Fisher<sup>(24)</sup> from transmission measurements on samples with impurity concentrations of  $10^{15} \text{ cm}^{-3}$ . The absorption peaks they observed were interpreted as optical transitions from the ground state of the impurities to bound excited states. From a combination of theory and experiment they deduced the energy level schemes shown in Fig. 1. As can be seen the higher excited states of all acceptors are virtually the same, as was predicted by the effective mass theory. Only the

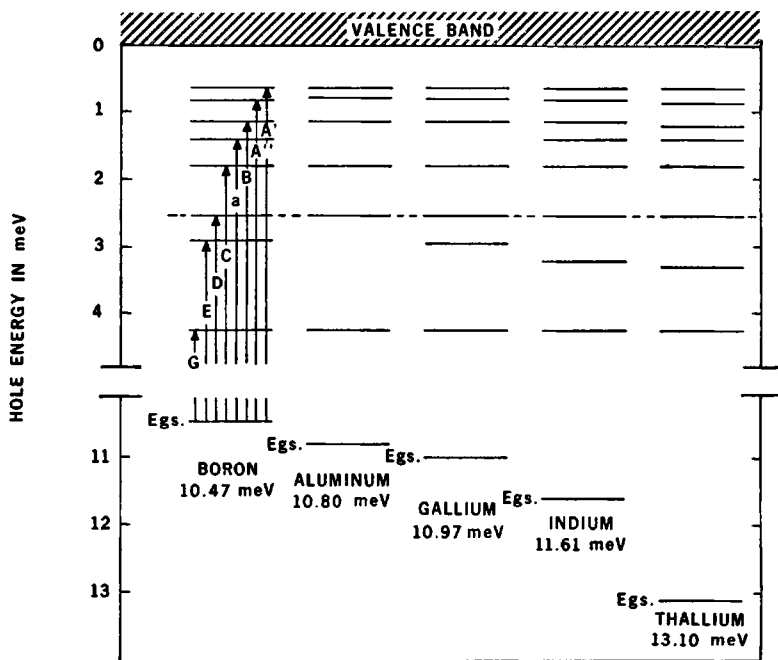


Fig. 1 : The bound states of some shallow acceptors in germanium, as determined by Jones and Fisher<sup>(24)</sup>. The numbers in the figure give the ground state energies ( $E_{g.s.}$ ) of the impurities relative to the valence band.

most strongly bound states, and especially the ground state  $E_{g.s.}$ , differ among different impurities. The resulting absorption spectra therefore have the same structure but are shifted versus each other, the shifts depending on the differences in ground state energies.

The disadvantage of using transmission spectroscopy for the study of impurities is that this technique asks for samples with a high impurity concentration. However, the size of the wave functions then produces a severe overlapping between the impurity atoms, resulting in a relatively low resolution in the spectra. In contrast with this PTIS on very pure samples, where the impurity atoms are far apart, produces spectra with very narrow lines.

### 3. Photothermal ionization spectroscopy (PTIS).

In PTIS one measures the change in conductivity of a sample under illumination with radiation in the same region as used in transmission spectroscopy. For shallow donors and acceptors in germanium this is the far infrared region (wave number  $10 - 200 \text{ cm}^{-1}$ ).

As an example we show in Fig. 2 the photoconductivity spectra at several temperatures of a polycrystalline sample, containing residual impurities in a concentration of  $\sim 10^{10} - 10^{11} \text{ cm}^{-3}$ . It is seen that above a temperature of 5 K a photoconductivity spectrum with a distinct line structure arises for photon energies below the impurity ionization energy (i.e. the onset of the continuum at  $\sim 85 \text{ cm}^{-1}$ ). The positions of the peaks on the energy scale coincide with the peaks in the optical absorption spectrum of the boron and aluminum acceptors known from the literature<sup>(24)</sup>. This proves that the lines are due to optical transitions of the charge carriers from the ground state to bound excited states of the impurity. However, in order to contribute to the conductivity the charge carriers must be released into the free band. We see in Fig. 2 that the heights of the peaks increase with increase in temperature. This suggests that the ionization of the impurity is a two-step process: 1) charge carriers in the ground state are optically excited to some intermediate state; 2) Subsequently a fraction of these excited charge carriers may be promoted thermally into the free band, producing an increase in conductivity of the sample. This is why the name of photothermal ionization was introduced for this two-step process.

In order to obtain enough phonons of sufficient energy one would like to keep the Ge sample at high temperatures, while to keep most of the impurities in their ground state low temperatures would be preferred. This leads to an experimentally-determined optimum temperature of around 8 K for observation of shallow impurities in germanium.

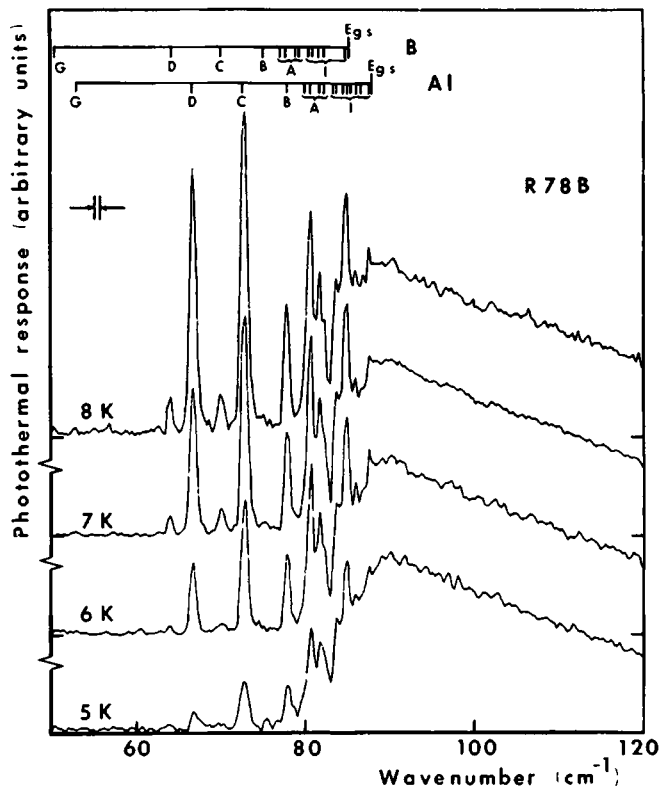
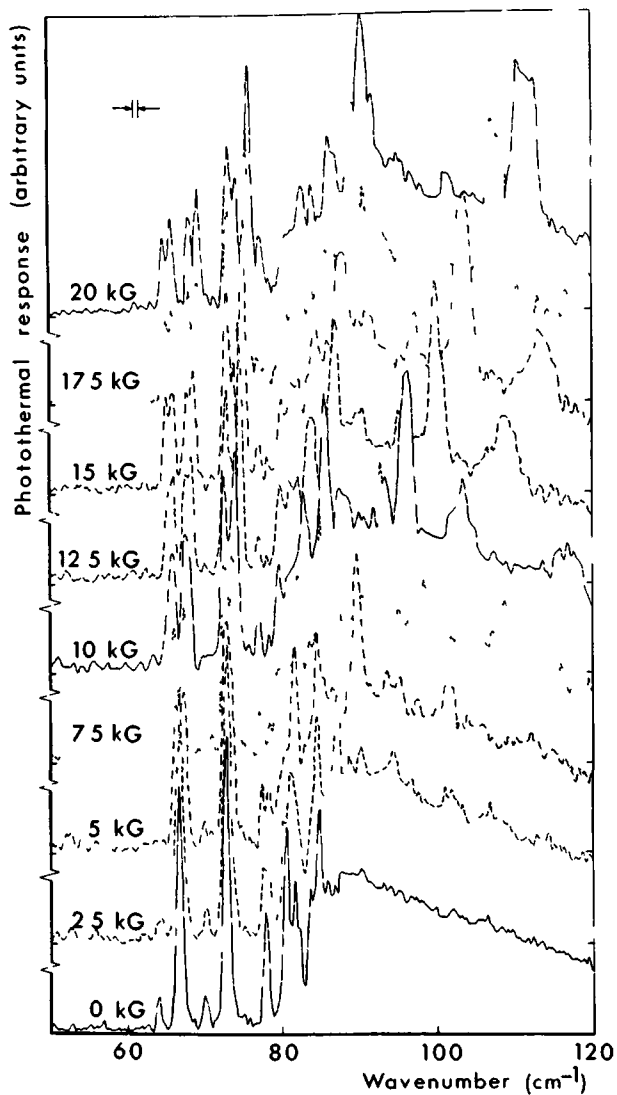


Fig. 2 : Photoconductivity spectra at several temperatures of a polycrystalline Ge sample, containing aluminum and boron in a concentration of  $\sim 10^{10} - 10^{11} \text{ cm}^{-3}$ .

The most important and essential difference between PTIS and other analysis methods is that the recorded signal does not decrease with decrease of impurity concentration down to a very low concentration<sup>(7)</sup>. This is a consequence of the fact that both the average number of current carriers,  $n$ , and the change  $\Delta n$  under illumination depend equally on the concentration of the impurities. The voltage sensitivity of the photo-response is proportional to the relative change of the carrier concentration  $\Delta n/n$ . This enables one to detect the impurities even if their concentration is extremely low and undetectable by all other analysis methods.

As the positions of the spectral lines reflect the energy spectrum of the impurity atoms PTIS enables the determination of the chemical nature of





*Fig. 3 : Photoconductivity spectra at 7.5 K of shallow acceptors in germanium for several values of the applied magnetic field (same sample as in Fig. 2).*

the impurities. Accurate spectra for a number of donors and acceptors can be found in the literature (for instance in Refs. 8,9 and 10). The identification of the impurities in the spectra of Fig. 2 was done by comparison with the B and Al spectra given by Haller and Hansen<sup>(9)</sup>; these spectra are shown in the top of Fig. 2.

#### 4. Applications of PTIS.

Originally the PTIS technique was only used as an analytical tool in the production of ultra-pure germanium. However, the inherent advantages over other analysis methods, such as the much better spectral resolution, have predicted PTIS to become one of the most powerful tools for the study of the basic physical aspects of germanium and a variety of other semiconductors.

The research subjects in which we have been engaged are:

1. An accurate determination of the ground state energies of some impurities from the temperature dependence of the peak intensities in the photoconductivity spectra;
2. The influence of a magnetic field on the energy level scheme of acceptors in germanium.

Figure 2 showed the effect of the temperature on the intensity of the peaks in a spectrum, whereas Fig. 3 shows the effect of a magnetic field on the same spectrum. As can be seen from Fig. 3 the spectrum not only exhibits a splitting or shift of the peaks below the continuum onset, but also develops a set of peaks in the continuum.

The results of our investigations will be presented in the form of a set of papers. Preliminary results are presented in Section IV.2, while in Sections IV.3 and IV.5 the temperature effect is explained in more detail. Section IV.4 is a supplement to Section IV.3. In Section IV.6 we present measurements in a magnetic field on a monocrystalline sample, where the orientation was known. The experimental results will be discussed in terms of existing theories about Zeeman effect and Landau levels in germanium.

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# MAGNETO-OPTICAL DETERMINATION OF THE GROUND STATE LEVELS OF SOME SHALLOW IMPURITIES IN HIGH PURITY GERMANIUM

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We determine the impurity ground state levels of aluminium boron and phosphor in germanium from very accurate measurements of the photo thermal conductivity as a function of the frequency of the incident radiation the temperature and the magnetic field strength From our magnetic data information is obtained on the field dependence of even the highest levels observed up to now

## 1. Introduction and experimental method

During the last decade there has been considerable interest in the study of the energy levels of the impurity states in germanium and silicon The experimental techniques used fall mainly into two groups far infrared transmission measurements using spectrometers or interferometers on samples containing impurities at concentrations of  $\sim 10^{15}$ – $10^{16}$  at/cm<sup>3</sup>, and photo-thermal conductivity measurements which enable investigation of high-purity samples with  $10^{10}$ – $10^{11}$  impurities/cm<sup>3</sup> Here also far infrared grating spectrometers or interferometers and sometimes lasers are used The experimental results of these two methods are in good agreement both with each other and with theoretical calculations For a review see Bassani et al [1]

However, up to now the precise determination of the ground state levels of the impurities has met difficulties Seccombe and Korn [2] assume that the two step photo thermal conductivity process has a relative ionization probability which varies as  $P_{ion}(T) \propto \exp \{-(E_B - E_i)/kT\}$ , where  $E_i$  is the energy of the intermediate level and  $E_B$  is the energy of the banded higher energy levels From the measured temperature dependence they find activation energies  $E_B - E_i$  which lie about 3 cm<sup>-1</sup> below the ionization energy of the isolated impurity which is attributed to impurity banding The same temperature dependence has been assumed by Simmonds et al [3] and previously by Nagasaka [4] whereby the latter remarks that the measured temperature dependence deviates from an exponential behaviour This is tentatively attributed to the effect of clustering of the impurities

Haller and Hansen [5] on the other hand determine the energy difference of the ground state energy and the band edge from the shape of the onset of the continuum observed in their photo thermal conductivity data They assume that this shape is due to the Fermi distribution of the electron energy at the top of the (valence) band Their obtained values for the ground state energies are about 1 cm<sup>-1</sup> higher than the theoretical values from the literature However, since the shape of the onset of the continuum, according to our observations, does not change with the temperature between 2 and 8 K, this shape cannot be due to the Fermi distribution but may be caused for instance by an averaging over  $k$  space by the effect of two phonon processes in the thermal process Although the frequency of the onset of the continuum gives information over the ground state energy its value cannot be obtained from this with high accuracy

It is our aim to derive the right temperature dependence of the photo thermal conductivity response and to obtain accurate values for the ground state energies for some shallow acceptors and donors in high purity germanium from detailed measurements of the photo-thermal conductivity as a function of the temperature For the determination of the ground state levels the degeneracy of the intermediate levels has to be known This degeneracy is obtained from measured level splittings in a magnetic field up to 20 kG The field dependencies of the energy levels are compared with current theories of both the low and high field region

The measurements were performed both on polycrystalline and on monocrystalline samples obtained from the same ingot of high purity

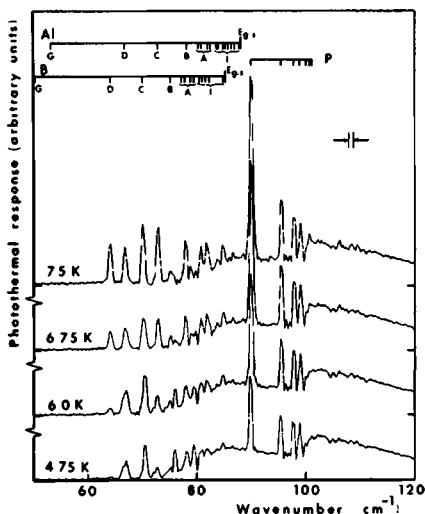


Fig. 1 Photoconductivity spectra at several temperatures of a Ge sample containing both donors (P) and acceptors (B and Al). The Al and B spectra at the top are after Haller and Hansen [5], the P spectrum after Seccombe [2]

germanium containing aluminium, boron and phosphor in concentrations ranging from  $10^{10}$  to  $10^{11}$  at./cm<sup>3</sup>. The concentrations of the impurities of the two monocrystalline samples J78A and J78B were ( $2 \times 10^{11}$  at./cm<sup>3</sup> B,  $4.7 \times 10^{10}$  at./cm<sup>3</sup> Al,  $1.9 \times 10^{10}$  at./cm<sup>3</sup> P) and ( $1.5 \times 10^{11}$  at./cm<sup>3</sup> P,  $4.7 \times 10^{10}$  at./cm<sup>3</sup> Al), respectively. These concentrations were determined with the method outlined in ref. 6 where also the experimental details will be given.

As is shown in fig. 1 with this method both donor and acceptor impurities present in the same sample can be investigated. Using a superconducting coil magnetic fields up to 20 kG could be applied.

## 2. Determination of the ground state energy

Generalizing the calculations of Spenke [7] it can be shown that the thermal ionization mechanism associated with the two-step photo-thermal conductivity process leads to a line intensity  $I$  given by:

$$I = A \left[ 1 + \left( \frac{1}{g g'} \right) \exp(\Delta E/kT) \right]^{-1} \quad (1)$$

for both donors and acceptors. Here  $\Delta E$  is the

energy-difference between the intermediate level and the band,  $g$  is the degeneracy of the excited level and  $g'$  is the degeneracy of the band at  $k=0$ . The factor  $A$  is proportional to the concentration of the impurity and to the optical transition probability of the spectral line.

We measured the temperature dependence of the line intensities at temperatures ranging from 2 to 10 K. As an example some of these spectra are shown in fig. 1. All spectra were corrected for the spectral background of the far infrared interferometer and were normalized to the same continuum height. This continuum is caused by direct optical excitations to the band edge. The relative line intensities were plotted on a logarithmic scale versus  $10/T$  as is shown in fig. 2 for boron in the monocrystalline sample J78A. In this figure the solid curves represent computer fits to the experimental data to eq. (1) using  $\Delta E$  and  $A$  as parameters. The degeneracy factors  $g$  were deduced from the measured level splittings in a magnetic field (see the next section). The degeneracy of the conduction band of germanium is  $g' = 1$  and that of its valence band is  $g' = 2$ . As is shown in fig. 2 the measured temperature dependencies of the line intensities are fully in agreement with eq. (1) within the experimental error.

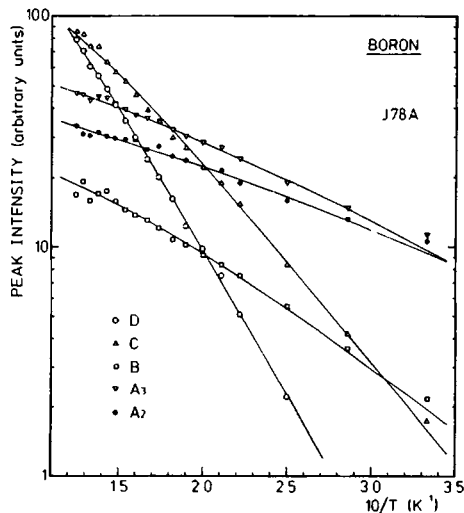


Fig. 2 Logarithmic plot of the relative line intensities versus  $10/T$  for several boron lines in sample J78A. The solid curves represent computer fits to eq. (1)

For each transition observed the sum of  $\Delta E$  and the energy of the transition  $E$  as obtained from the measured spectra yields the energy of the ground state  $E_g$  of the impurity involved. In this way we obtain for the ground state energies of aluminium, boron and phosphor the values presented in table I. As is shown in this table these values are in excellent agreement with the theoretically deduced values for Al and B of Jones and Fisher [8] and the value for P of Reuszer and Fisher [9].

Table I  
The ground state energies of Al, B and P in germanium in  $\text{cm}^{-1}$

	Al	B	P
$E_g$ (this work)	$87.1 \pm 0.4$	$84.5 \pm 0.1$	$102.9 \pm 0.5$
$E_g$ (refs. 8 and 9)	$87.1 \pm 0.2$	$84.5 \pm 0.2$	$102.9 \pm 0.3$

### 3. The effect of a magnetic field

In a low magnetic field, in the low field limit, the Zeeman term connected with the magnetic field produces a linear splitting of the  $l \neq 0$  levels while the diamagnetic term gives a quadratic correction. On the other hand, in the high field limit the impurity potential produces a set of sublevels associated with each Landau level. Again we refer to the review of Bassani et al. [1].

Experimentally Boyle et al. [10] investigated As donors in germanium by measuring the far infrared transmission of samples containing  $\sim 1.5 \times 10^{15}$  arsenic atoms/ $\text{cm}^3$  in magnetic fields up to 30 kG. Horii et al. [11] did the same for arsenic and antimony up to 46 kG, while Soepangkat et al. [12] studied boron and thallium acceptors in germanium in magnetic fields up to 20 kG using the same method. Nsida et al. [13] investigated germanium containing antimony and arsenic impurities by measuring the photoconductivity in a varying magnetic field up to 40 kG under the influence of intraband radiation of a laser source.

In our experiment we measured the photo-thermal conductivity response of our high-purity germanium samples as a function of the frequency of the far infrared radiation at several values of the magnetic field up to 20 kG. Fig. 3

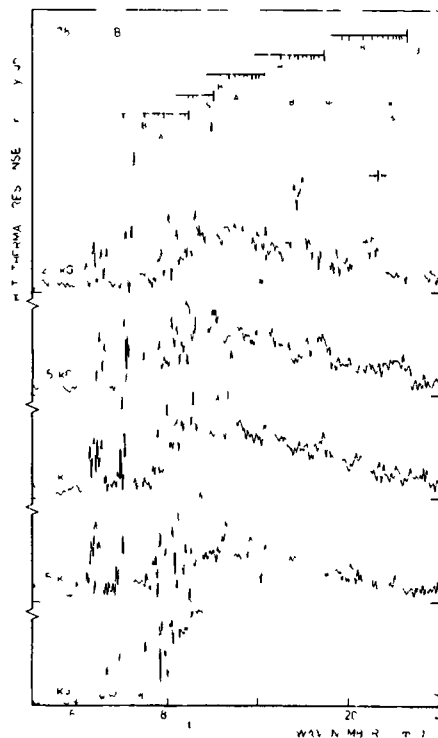


Fig. 3 Photoconductivity spectra of sample J78A at 7.5 K for several values of the magnetic field with  $B//[100]$ . Impurity concentrations:  $2 \times 10^{17}$  at/ $\text{cm}^3$  B,  $4.7 \times 10^{17}$  at/ $\text{cm}^3$  B,  $1.9 \times 10^{17}$  at/ $\text{cm}^3$  Al, P. The labelling at the top shows the boron spectrum with the Zeeman splitting and the first four Landau levels with their associated sublevels at 15 kG.

shows that the general behaviour sketched above manifests itself in these spectra in a very clear and illustrative way. After a shift and splitting of the hydrogen-like levels lying below the band edge to higher energies in low fields, these levels are forced up to energies higher than the band edge energy in higher fields. At these fields the continuum above the band edge changes more and more in a number of peaks: the Landau levels with their associated sublevels. By plotting the resonance frequencies as a function of the magnetic field each line could easily be followed and identified. In this way the identification indicated in fig. 3 has been performed. The notation is that of Jones and Fisher

[8] From the energy differences of the first four Landau levels which can be identified we obtain an effective mass of  $m^*/m = 0.08 \pm 0.05$  which has to be compared with the effective mass  $m^*/m = 0.04$  of the light hole, cf. Sze [14]. The inaccuracy is caused by the non linearity of the shift of the Landau levels with the magnetic field in this region since the high field limit is not entirely reached. For the lower levels, our results are in accordance with those of the authors mentioned above. However, in contrast to these authors, we obtained also information about the higher levels. More quantitative results about these magnetic data will be published separately [15].

We thank Dr. H. J. A. van Dijk, Philips Research Laboratories, for providing the Ge samples and performing the impurity concentration analysis. This work was performed as part of the research program of the Stichting voor Fundamenteel Onderzoek der Materie (FOM) with financial support from the Nederlandse Organisatie voor Zuiver Wetenschappelijk Onderzoek (ZWO).

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### IV.3 Temperature dependence of the photothermal conductivity of high-purity germanium containing very low concentrations of Al, B, and P

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The temperature dependence of the photothermal conductivity in the far-infrared region (10 to 200  $\text{cm}^{-1}$ ) is studied in pure germanium containing  $\sim 10^{10}$  atoms/ $\text{cm}^3$  of Al, B, and P. A theoretical expression for the temperature dependence of the intensities is given which is in excellent agreement with the experimental results. From these very detailed spectra, reliable values for the ground-state energies of the impurities are obtained and a quantitative chemical analysis of the impurity concentration as a function of the position in a Czochralski-grown single crystal is achieved.

## I INTRODUCTION

The photothermal conductivity in the far-infrared region of very pure germanium containing residual impurities of very low concentrations ( $\approx 10^{10}$  atoms/ $\text{cm}^3$ ) has been the subject of several studies in the past.<sup>1-6</sup> A recent review of this subject was given by Kogan and Lifshits.<sup>7</sup> However, surprisingly enough, the temperature dependence of this effect has drawn little attention so far.

It is generally believed that, in order to get a photothermal response, the donors or acceptors are ionized through a two-step process. First, the electrons (or holes) are excited from their ground state to a higher bound state by irradiation with light in the far-infrared region (10–200  $\text{cm}^{-1}$ ). Then, a subsequent excitation to the free conduction (or valence) band occurs by interaction with thermal phonons. The photoconductivity is measured as a function of the radiation frequency, whereas the sample temperature is held constant, typically between 2 and 10 K. At higher temperatures the photoconductivity signal decreases due to the increase of the carrier density in the free band, while lower temperatures provide too few thermal phonons necessary for the second step of the process. Although the influence of the temperature on the effect is thus qualitatively well understood, the precise temperature dependence has never been studied in a quantitative way. Most authors assume a temperature dependence typically of the form  $\exp(-\Delta E/kT)$ , where  $\Delta E$  is the energy difference between the excited level and the conduction (or valence) band edge.<sup>2,3,6,8</sup> However, the agreement with experimental results is rather poor, and estimates for the binding energy of the impurities yield values which are about 3  $\text{cm}^{-1}$  too low.<sup>3</sup>

In this paper, we will concentrate upon this aspect. In Sec. II the experimental details will be described, while in Sec. III the results will be presented and

analyzed. It will be shown that the measurements are in very good agreement with our calculations of the temperature dependence of the photothermal response. Finally, a qualitative and a quantitative chemical analysis of the samples will be given, based on the photoconductivity data. This analysis will be compared with results based on Hall-effect measurements on the same ingot from which the samples were obtained. Preliminary results of these investigations have been published before.<sup>9</sup>

## II SAMPLE PREPARATION AND EXPERIMENTAL DETAILS

All our germanium samples were cut from the same boule. These ultrapure samples were kindly supplied by Dr. H. J. A. van Dijk of the Philips Research Laboratories, who also performed Hall-effect measurements to determine the residual impurity concentration. The polycrystalline samples, R78A and R78B, were obtained from both of the ends of the high-purity zone-refined ingot. Then, from the remaining ingot, a single crystal was grown with the Czochralski method. Again from both ends of this crystal two monocrystalline samples J78A and J78B were cut. Figure 1 shows the variation of the impurity concentration along this single crystal as calculated from Hall-effect measurements, assuming aluminum, boron, and phosphorus as the main impurities and using the known segregation properties of these elements.

The samples used for the photoconductivity measurements were cubes with dimensions of  $\sim 1 \times 1 \times 1 \text{ cm}^3$ . The cubes were polished and etched, and then electrical contacts were produced by wetting with a mixture of In-Hg (50/50). Although these contacts were not entirely ohmic, we found no evidence that the contacts had any effect on the data.



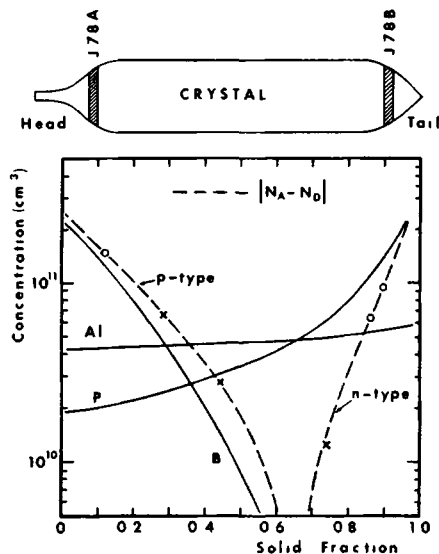


FIG 1 Distribution of phosphorus, boron, and aluminum impurities along a germanium single crystal, calculated from Hall-effect measurements (Segregation coefficients used  $P: k_P=0.1$ ,  $B: k_B=5.7$ ,  $Al: k_{Al}=0.9$ ).  $N_A$  and  $N_D$  are the concentrations of the acceptors and donors, respectively

The samples were positioned in a semispherical cavity wherein the chopped far-infrared radiation entered through a conical light pipe. Thermal contact was made by clamping the sample onto the base of the cavity. The clamp and the sample could be heated above the temperature of the surrounding helium bath with an electronically controlled heater system. The temperature of the sample was stabilized within 0.1 K. An Allen-Bradley carbon resistor was used as a thermometer.

Using standard lock-in techniques, the photoconductivity was measured as a function of the wave number of the incident radiation produced by a Grubb-Parson Michelson interferometer operating with phase modulation at 90 Hz. The data were punched on tape, and subsequently handled and analyzed on a PDP-12 laboratory computer linked to a plotter. All spectra were corrected for the varying spectral background of the interferometer by dividing through a spectrum obtained by replacing the Ge sample with a Si bolometer.

### III. RESULTS AND DISCUSSION

Figure 2 shows the photothermal response as a function of the wave number of the incident radiation for the polycrystalline samples R78A and R78B.

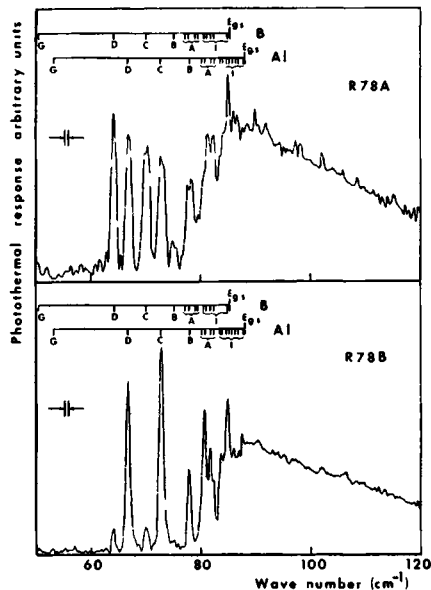


FIG 2 Photoconductivity spectra of the polycrystalline germanium samples R78A and R78B containing boron and aluminum impurities. The spectra at the top are from Refs 4 and 5.

at a temperature of 7.5 K. Indicated are also the expected transitions for boron and aluminum known from the literature.<sup>4,5</sup> From these spectra one can conclude that R78A contains B and Al in about equal amounts, while in sample R78B the Al concentration exceeds the B concentration as is expected in a zone-refined ingot. Figure 3 shows similar data for the monocrystalline samples, J78A contains both B and Al, with B having the highest concentration, while in sample J78B, apart from phosphorus, some lithium is present as well.<sup>3,5</sup> In addition, indicated by an asterisk in the figure, in J78B some other transitions, with an origin not entirely known, are found, these transitions may be due to a LiO complex.<sup>10</sup> The wave numbers where the transitions are observed agree within 0.2  $\text{cm}^{-1}$  with the expected ones from the literature<sup>3-5</sup>, however, due to the limited resolution, the A and I transitions of aluminum and boron are not fully resolved in these spectra. The continuum in the spectra is caused by direct optical transitions from the ground state into the valence (or conduction) band. In this context, it is interesting to note that with the method of photothermal ionization both donor and acceptor impurities can be seen simultaneously in one sample.

The strength of a spectral line is proportional to the thermal ionization probability of the excited states

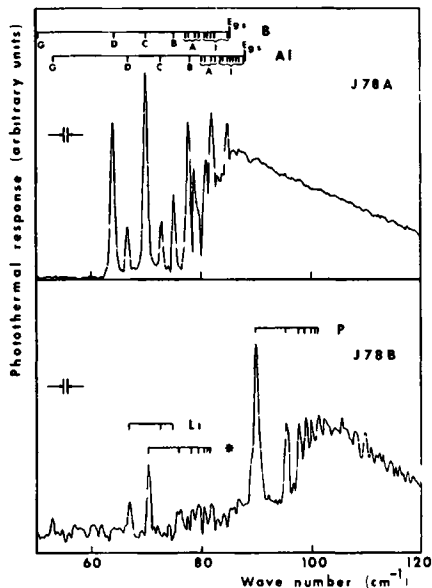


FIG 3 Photoconductivity spectra at  $T = 7.5$  K of the monocrystalline germanium samples J78A, containing boron and aluminum, and J78B, containing phosphorus, lithium, and probably a LiO complex. The spectra at the top are from Refs 3, 4, 5, and 10.

and this can be studied by comparing the line intensities at various temperatures with the height of the continuum.<sup>2</sup> Figure 4 shows the spectrum of sample J78A, containing aluminum and boron, at different temperatures. All these spectra are corrected for the spectral background of the interferometer and are normalized to the same continuum height. Similar measurements were also performed on a sample containing phosphorus.

From the voltage drop across the sample, the current through the sample and the measured change in voltage in a typical experiment, it is possible to get a rough estimate of the number of carriers involved. Typical data yield for the number  $n$  of the free carriers in the band  $n \sim 10^8$  ( $\text{cm}^{-3}$ ) and the change  $\Delta n$  in carriers due to photothermal ionization  $\Delta n \sim 10^5$  ( $\text{cm}^{-3}$ ). This means that  $n$  and  $\Delta n$  are small compared with the total number of impurity states. Therefore the average number of optical excitations to one of the excited levels may be considered as independent of temperature at very low temperature. Since the excited levels are very localized in  $k$  space,<sup>11</sup> thermal ionization from this level will at low temperatures only take place to the nearest band extremum. The average number of ionizations in

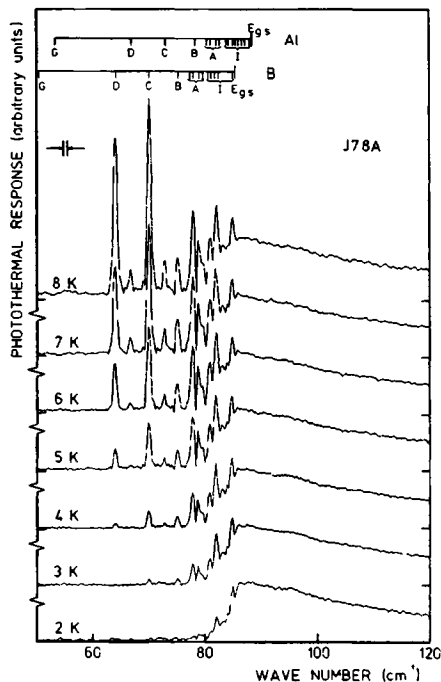


FIG 4 Photoconductivity spectra at several temperatures of the monocrystalline germanium sample J78A. Impurity concentration B:  $2.1 \times 10^{11}$  atoms/ $\text{cm}^3$ , Al:  $4.3 \times 10^{10}$  atoms/ $\text{cm}^3$ .

equilibrium is found in the usual way by optimizing the number of possibilities for ionization at constant total energy. Hereby the strong Coulomb interaction of electrons in an excited state has to be taken into account which can be done in a similar way as in considerations regarding donor and acceptor ionization probabilities.<sup>12</sup> In this manner the line intensity  $I$ , which is proportional to the average number of ionizations is found to be given by

$$I = A \left[ 1 + \frac{1}{gg'} \exp \left( \frac{\Delta E}{kT} \right) \right]^{-1} \quad (1)$$

for both donors and acceptors. Here,  $\Delta E$  is the energy difference between the intermediate level and the band edge,  $g$  is the degeneracy of the excited level, and  $g'$  is the degeneracy of the band. The factor  $A$  is proportional to the concentration of the impurity and to the optical transition probability for the spectral line under consideration. For donor states, where spin-orbit interaction is small, thermal processes with spin flip may be neglected so that  $g'$  is the band de-

generacy apart from spin, thus  $g' = 1$  for germanium. Although for acceptor states the situation is less obvious, the degeneracy of the valence band in germanium will be taken only as the effective-mass degeneracy at  $\bar{k} = 0$ , yielding  $g' = 2$ .

Figure 5 shows the relative line intensities for the phosphorus spectrum, plotted on a logarithmic scale, as a function of the inverse temperature. The solid curves in this figure represent computer fits of Eq. (1) to the experimental data, using  $\Delta E$  and  $A$  as adjustable parameters. The degeneracy factors  $g$  were deduced from measured level splittings due to magnetic fields. Here our own observations<sup>9,13</sup> confirmed the splittings found in the literature for the lower levels of the donors<sup>11</sup> and acceptors,<sup>14,15</sup> and extended the results to higher levels. As is shown in Fig. 5, the measured temperature dependence of the line intensities is within experimental accuracy in agreement with Eq. (1). Similar results were obtained for the Al and B spectra of sample J78A. Table I collects the numerical results of the computer fits for the parameters  $A$  and  $\Delta E$ . In the fitting procedure also the product  $gg'$  was varied by integer values. It was not possible however to obtain a reasonable fit for other values of  $gg'$  than those quoted which supports the assumption of the neglect of spin in the band degeneracy made above.

The energy of the ground state  $E_{gs}$  of the impurity involved is now simply the sum of  $\Delta E$  and the energy of the optical transition  $E_{trans}$  (as obtained from the measured spectra). In this way it is possible to obtain the ground-state energies of boron, aluminum and phosphorus as presented in Table II. As can be seen

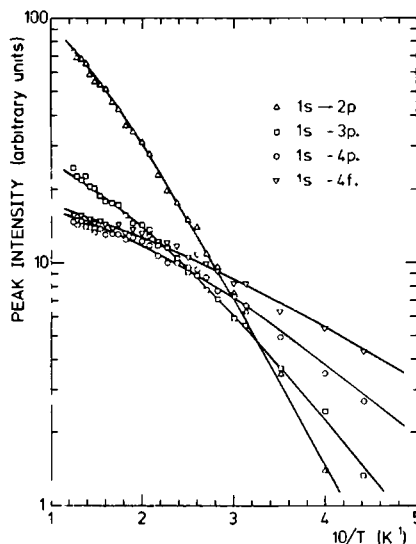


FIG. 5. Logarithmic plot of the relative line intensities vs  $10/T$  for several phosphorus lines. The solid curves represent computer fits to Eq. (1). (See also Table I.)

from this table, these experimentally found values are in excellent agreement with the theoretically deduced values of Jones and Fisher<sup>16</sup> for B and Al and of Reuszer and Fisher<sup>17</sup> for P.

As a factor  $A$  is proportional to the impurity con-

TABLE I. Values of  $A$  and  $\Delta E$  as obtained from a computer fit from the temperature dependence of various lines in the B, Al, and P spectra [see Eq. (1)].

Transition		$E_{trans}(\text{cm}^{-1})$	$g$	$A$ (arb. units)	$\Delta E(\text{cm}^{-1})$
Boron (J78A)	D	$64.0 \pm 0.2$	4	$450 \pm 15$	$20.5 \pm 0.2$
	C	70.1	3	$265 \pm 8$	$14.4 \pm 0.2$
	B	75.2	2	$44 \pm 1$	$9.3 \pm 0.2$
	$A_3$	77.9	1	$123 \pm 3$	$6.6 \pm 0.2$
	$A_2$	78.9	1	$78 \pm 2$	$5.6 \pm 0.2$
Aluminum (J78A)	D	$66.7 \pm 0.2$	4	$70 \pm 7$	$20.5 \pm 0.6$
	C	72.8	3	$41 \pm 3$	$14.1 \pm 0.5$
	$I_3$	85.0	1	$36 \pm 2$	$2.3 \pm 0.5$
Phosphorus	$1s \rightarrow 2p_{\pm}$	$90.0 \pm 0.2$	4	$285 \pm 10$	$12.9 \pm 0.2$
	$1s \rightarrow 3p_{\pm}$	95.5	4	$44 \pm 3$	$7.5 \pm 0.2$
	$1s \rightarrow 4p_{\pm}$	97.9	4	$26 \pm 2$	$5.3 \pm 0.3$
	$1s \rightarrow 4f_{\pm}$	99.0	2	$33 \pm 3$	$4.0 \pm 0.3$

TABLE II Energies of the ground state of B, Al, and P in germanium, measured in  $\text{cm}^{-1}$ 

	B	Al	P
$E_{gs}$ (experimental, this work)	$84.5 \pm 0.2$	$87.1 \pm 0.3$	$103.0 \pm 0.2$
$E_{gs}$ (theoretical, Refs 16 and 17)	$84.5 \pm 0.2$	$87.1 \pm 0.2$	$102.9 \pm 0.3$

centration, it is in principle possible to get some information on the relative impurity content from a determination of  $A$ . Assuming the same transition probability for transitions between similar states of different donors or acceptors, the ratio of the  $A$  factors for corresponding lines will therefore indicate the ratio of the concentrations of the impurities. For sample J78A one finds for the ratio of the  $A$  factors for the C and D transitions of boron and of aluminum a value of  $6.4 \pm 0.5$ . This value can be compared with the concentrations obtained from Hall-effect measurements as presented in Fig. 1. From this figure one concludes that for sample J78A the ratio of the concentrations of B and Al is

$$(2.1 \times 10^{11} \text{ cm}^{-3}) / (4.3 \times 10^{10} \text{ cm}^{-3}) = 5 \pm 1,$$

which is in good agreement with the value obtained from the photoconductivity measurements.

In conclusion we can remark that we have given a theoretical expression for the temperature depen-

dence of the photoconductive response which is in excellent agreement with our experimental results. The values found from these data for the ground-state energies of various impurities confirm theoretical calculations very well. In addition, we have shown that the concentration ratios of the impurities as obtained from photoconductive measurements are in reasonable agreement with values obtained from Hall-effect measurements.

#### ACKNOWLEDGMENTS

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<sup>16</sup>R. L. Jones and P. Fisher, *J. Phys. Chem. Solids* **26**, 1125 (1965).

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#### IV.4 DETERMINATION OF THE IMPURITY CONCENTRATION PROFILE IN A Ge SINGLE CRYSTAL GROWN WITH THE CZOCHRALSKI METHOD

Figure 1 of the preceding paper (Section IV.3) showed the distribution of the phosphorus, boron and aluminum impurities along the germanium crystal, from which our samples J78A and J78B were cut. In this section it will be shown how this figure was obtained from Hall effect measurements, performed by Dr. H.J.A. van Dijk of Philips Research Laboratories.

The variation of the impurity concentration along a crystal which is grown by the Czochralski method, is caused by the segregation of dissolved impurities during non-equilibrium solidification<sup>(1,2)</sup>. For low impurity concentrations the ratio of the concentration  $C_S$  in the just solidified material and the concentration  $C_L$  in the remaining liquid will be a constant  $K$ , the so-called segregation coefficient.

If the location in the crystal is denoted by the weight fraction  $g$  of the original melt that was solidified at that point, the distribution of every impurity along the crystal can be described by<sup>(3)</sup>:

$$C_S(g) = C_S(0)(1-g)^{K-1}, \quad (1)$$

where  $C_S(0)$  is the concentration at the head of the crystal (i.e.  $g=0$ ).

The number of uncompensated holes  $N_A - N_D$  is determined by the net impurity concentration

$$[N_A - N_D](g) = \sum_A C_S^A(g) - \sum_D C_S^D(g), \quad (2)$$

where A and D denote acceptors and donors respectively. The value of  $N_A - N_D$  at various positions in the crystal can be deduced from Hall effect measurements<sup>(4)</sup>.

The photothermal conductivity spectra of samples J78A and J78B (Fig. 3, Section IV.3) reveal that the Al and B acceptors and P donor are the most important residual impurities in our crystal. From the measured values of  $N_A - N_D$  for three different locations  $g$ , indicated in Fig. 1 by small circles, equations (1) and (2) can be solved for  $C_S^{Al}(0)$ ,  $C_S^B(0)$  and  $C_S^P(0)$ . The concentration profile for Al, B and P can then be calculated with Eq. (1). These

solutions are represented in Fig. 1 (Section IV.3) by the solid curves, whereas the dashed curve shows the calculated profile of  $|N_A - N_D|$ .

As a check the values of  $N_A - N_D$  were also measured for three intermediate positions, indicated in Fig. 1 (Section IV.3) by crosses; they fit very well to the calculated  $|N_A - N_D|$  curve.

The segregation coefficient used for these calculations were:  $K_P=0.1$ ,  $K_B=5.7$  and  $K_{Al}=0.9$ . Aluminum has normally a segregation coefficient  $K_0 = C_S(0)/C_L(0)=0.1$ , but when a crystal is grown with the Czochralski method Al shows only a small segregation<sup>(5)</sup>, so that in Eq. (1) we have used  $K_{Al}=0.9$ .

However, in the zone-refining process<sup>(6)</sup> used for the purification of the original Ge ingot, one has  $K_{Al}=0.1$ . We find boron and aluminum in about equal amounts in the region where the zone is started (sample R78A was cut from this region). Near the opposite end the boron concentration is reduced but aluminum is increased (sample R78B).

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#### IV.5 TEMPERATURE DEPENDENCE OF THE PHOTOTHERMAL CONDUCTIVITY OF SEMICONDUCTORS AT LOW TEMPERATURES

##### ABSTRACT

A careful analysis has been made of the temperature dependence of the signal strength in a photoconductivity experiment in a semiconductor. The temperature dependence of the line intensity in a photothermal conductivity spectrum has been calculated. The results are in excellent agreement with measurements on high purity germanium containing boron, aluminum and phosphorus as residual impurities ( $\sim 10^{11}$  atoms/cm<sup>3</sup>).

## 1. Introduction.

Photothermal conductivity spectroscopy is of considerable interest for the study of impurity states in semiconductors (Lifshits and Nad' 1965). It offers the most sensitive method of analysis of impurities at very low concentrations ( $10^{10}$ - $10^{11}$  atoms/cm<sup>3</sup>) in semiconductors. The method is sensitive enough to investigate subtle features such as the magnetic field dependence of central cell corrections (Stoelinga et al 1978). A lot of work has been done related to this subject. A recent review has been given by Kogan and Lifshits (1977, further referred to as Ref. I).

However, the well known strong temperature dependence of the photothermal conductivity signal has met relatively little attention so far. Frequently a simple exponential behaviour,  $\exp(-\Delta E/kT)$ , has been assumed for the relative ionization probability of the two-step photothermal conductivity process (Lifshits et al 1968, Seccombe and Korn 1972, Stillman et al 1972, Simmonds et al 1974, Bykova et al 1976, Ref. I). Here  $\Delta E$  is the energy difference between the band edge and the intermediate level to which optical excitation from the ground state takes place followed by thermal ionization from this level to the band edge. The experimental results do not agree very well with this assumption; this has been tentatively attributed to the effect of clustering of the impurities (Nagasaka and Narita 1969).

In this paper we will analyze the temperature dependence of the photothermal conductivity in semiconductors in detail, and derive a theoretical expression for this temperature dependence which according to our measurements on shallow impurities in high purity germanium is in excellent agreement with the experimental results. This



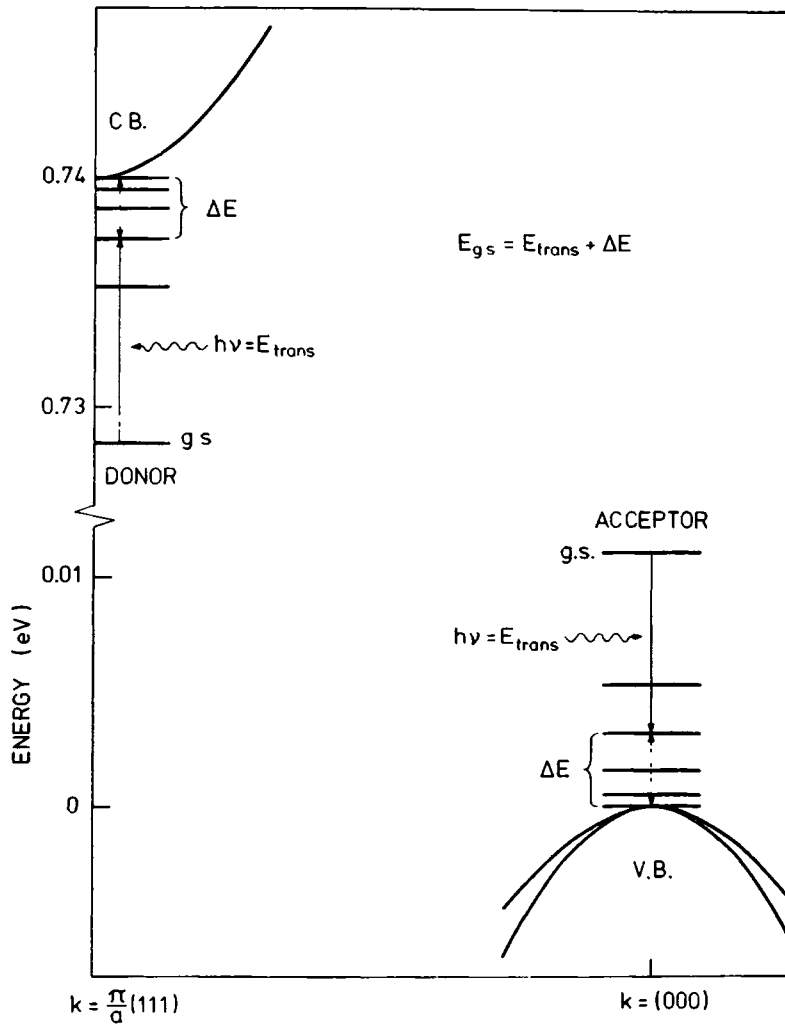


Fig. 1 : Illustration of the mechanism for photothermal ionization. Shown is part of the band structure of germanium along the (111) direction in  $\vec{k}$  space, with the conduction band (C.B.), the valence band (V.B.) and some shallow donor and acceptor levels, i.e. the impurity ground state (g.s.) and the excited levels.

expression enables one to determine the ground state levels of the impurities very accurately from photothermal conductivity experiments (Jongbloets et al 1977 and 1979, further referred to as Ref. II and III). In analyzing the experimental data it will become clear why photothermal conductivity spectroscopy is applicable only in a relatively small temperature region (2-10 K).

## 2. Signal formation in a photoconductivity experiment.

In photothermal conductivity spectroscopy, electrons (or holes) are excited from their ground state with binding energy  $E_{g.s.}$  to an excited state by radiation of energy  $h\nu$  equal to the transition energy  $E_{trans}$ , as illustrated in figure 1. The radiation frequency  $\nu$  is usually in the far infrared region ( $10\text{-}200\text{ cm}^{-1}$ ). While being in such an excited state, the extra energy  $\Delta E$  needed to overcome the remaining part of the binding energy is gained by interaction with thermal phonons. This ionization can be observed as an increase in the electric conductivity of the sample. The photoconductivity is usually measured as a function of the radiation frequency at constant temperature, typically between 2 and 10 K, yielding characteristic "hydrogen-like" excitation spectra for the various shallow impurities (Refs. I, II and III). In the measured spectra one can distinguish between a continuum, caused by direct photo-ionization from the bound ground state into the free band, and sharp peaks due to the two-step photothermal ionization process via bound intermediate states.

In order to analyze the temperature dependence of the photothermal ionization process, we will treat a typical experiment as an example, the results are of general validity however. In such an

experiment the source for the chopped excitation radiation is usually a Michelson interferometer. The sample is connected to an electric circuit so that a bias current is provided by a battery ( $V_B$ ) via a load resistor ( $R_L$ ) in series with the sample. The change in voltage drop across the sample due to the radiation is measured at various temperatures, using standard lock-in techniques.

The electric conductance of a sample of area  $A$  and length  $\ell$  is given by  $Ae\mu n/\ell$ , where  $n$  is the free carrier concentration,  $\mu$  is the mobility and  $e$  is the charge of the carrier. Introducing the quantity  $\alpha = R_L Ae/\ell$  the d.c. voltage drop  $V$  across the sample may be written as

$$V = V_B(1 + \alpha\mu n)^{-1} . \quad (1)$$

Without the chopped excitation radiation the number of ionizations,  $n$ , will be determined by the background radiation and by the few phonons of sufficient energy available at low temperature. The chopped radiation causes a change  $\Delta n$  of the free carrier concentration, and thus a change in the conductivity of the sample. Assuming  $\Delta n \ll n$  the photoconductive response will be an a.c. voltage with amplitude

$$\Delta V = V_B \alpha \mu \cdot \Delta n (1 + \alpha \mu n)^{-2} \quad (2)$$

so that  $\Delta V$  is directly proportional to  $\Delta n$ .

One now has to distinguish between the change in free carrier density  $\Delta n_p$  due to photothermal ionization via an intermediate level and the change  $\Delta n_c$  due to direct photo-ionization. The "continuum" carrier density variation  $\Delta n_c$  is expected to be independent of temperature in the temperature region of present interest. However, it

will be shown below that the "peak" carrier density variation  $\Delta n_p$  is strongly temperature dependent, and that this temperature dependence can be investigated by looking at the ratio  $\Delta V_p / \Delta V_c$ .

### 3. Derivation of the temperature dependence of the photothermal ionization process.

As will be shown in section 4 in a typical experiment the free carrier densities  $n$ ,  $\Delta n_c$  and  $\Delta n_p$  are small compared to the total number of impurity states, so that most of the impurities remain in their ground state. The average number  $N$  of optical transitions to one of the excited states may therefore be treated as independent of temperature at the low temperatures at which these experiments are usually carried out, so that only the subsequent thermal ionization needs to be considered.  $N$  is directly proportional to the impurity concentration.

In electron-phonon interaction the longitudinal acoustic phonons are most important. At the low thermal energies ( $< 10$  K) where the experiments under consideration take place these phonons have a small wave vector  $\vec{k}$ . Since the excited states are very localized in  $\vec{k}$ -space (Bassani et al 1974), ionization by phonons from these levels will only take place to the nearest band extremum. Generally the excited state is  $g$ -fold degenerated, so the strong Coulomb interaction of electrons (holes) in such a state has to be taken into account. This can be done by stating that an excited level can be occupied by only one electron (hole) but in  $g$  different ways (Spence 1958). The number of ways,  $g'$ , in which a particle can be transferred from one of the  $g$  degenerate excited states to a band state depends on the degeneracy of the band and on the selection rules for such transitions via a

phonon. The number of ways  $W$  in which  $\Delta n_p$  particles can be transferred from the  $N$  optical excited levels to a band state is then given by:

$$W = (gg')^{\Delta n_p} \cdot N! [\Delta n_p! (N - \Delta n_p)!]^{-1} . \quad (3)$$

The equilibrium state is now easily obtained in the well known manner by demanding that  $W$  is maximal under the condition that the energy:

$$U = N \cdot E_{\text{trans}} + \Delta n_p \cdot \Delta E$$

is a constant. Maximizing (with respect to  $\Delta n_p$ ) of:

$$\ln W + \beta(U - N \cdot E_{\text{trans}} - \Delta n_p \cdot \Delta E)$$

using Stirling's formula yields:

$$\Delta n_p = N [1 + (\frac{1}{gg'}) \exp(\beta \cdot \Delta E)]^{-1} , \quad (4)$$

where  $\beta = 1/kT$ . The intensity  $I$  of a line in the photothermal conductivity spectrum relative to that of the continuum is now given by:

$$\begin{aligned} I &= \Delta V_p / \Delta V_c = \Delta n_p / \Delta n_c \\ &= A [1 + (\frac{1}{gg'}) \exp(\Delta E/kT)]^{-1} , \end{aligned} \quad (5)$$

where equation (2) has been used. Here  $A$  contains the optical transition probability of the line under consideration.

Since both  $\Delta n_p$  and  $\Delta n_c$  are proportional to the impurity concentration the intensity  $I$  is independent of the latter. If a sample contains several impurities the various  $\Delta n_c$ 's add up in forming one continuum band in the spectrum. The ratios of the intensities of corresponding peaks, all compared with the intensity of the combined continuum band, yield a measure of the ratios of the concentrations

of these impurities in the sample, cf. the experimental work of Ref. III.

As argued above, in equation (5)  $g$  is the degeneracy of the excited bound state while  $g'$  is partly determined by the selection rules for transitions via phonons from this state to the nearest band extremum. For donor states in germanium and silicon, where the spin-orbit interaction is small, thermal processes in which spin flips occur may be neglected and the selection rule simply converts to that of spin conservation. In this case  $g'$  is the band degeneracy apart from spin, thus  $g' = 1$  for germanium and silicon. Although for acceptor states, where spin-orbit interaction is not negligible, the situation is less obvious the degeneracy of the valence band at  $\vec{k} = 0$  will be taken in analogy only as the effective mass degeneracy, yielding  $g' = 2$  for germanium and silicon.

#### 4. Comparison with experimental results.

In this section the analysis of the previous sections will be compared with experimental results obtained on high purity germanium samples. These samples contained aluminum, boron and phosphorus as residual impurities in concentrations of  $10^{10}$ - $10^{11}$  atoms/cm<sup>3</sup> (Refs. II and III). The dimensions of the cube-shaped samples were  $1 \times 1 \times 1$  cm<sup>3</sup>. The bias voltage  $V_B$  was provided by a 12.6 Volt mercury battery over a load resistor  $R_L$  of 5 M $\Omega$  in series with the sample. The quantity  $\alpha$  introduced in section 2 has thus a value of  $8 \times 10^{-13}$  [V $\cdot$ cm $\cdot$ s].

The photoconductive signal strength  $\Delta V_p$  and  $\Delta V_c$ , corresponding with the change in carrier density  $\Delta n_p$  and  $\Delta n_c$  respectively, can be deduced from the measured spectra by taking the area of a peak and of

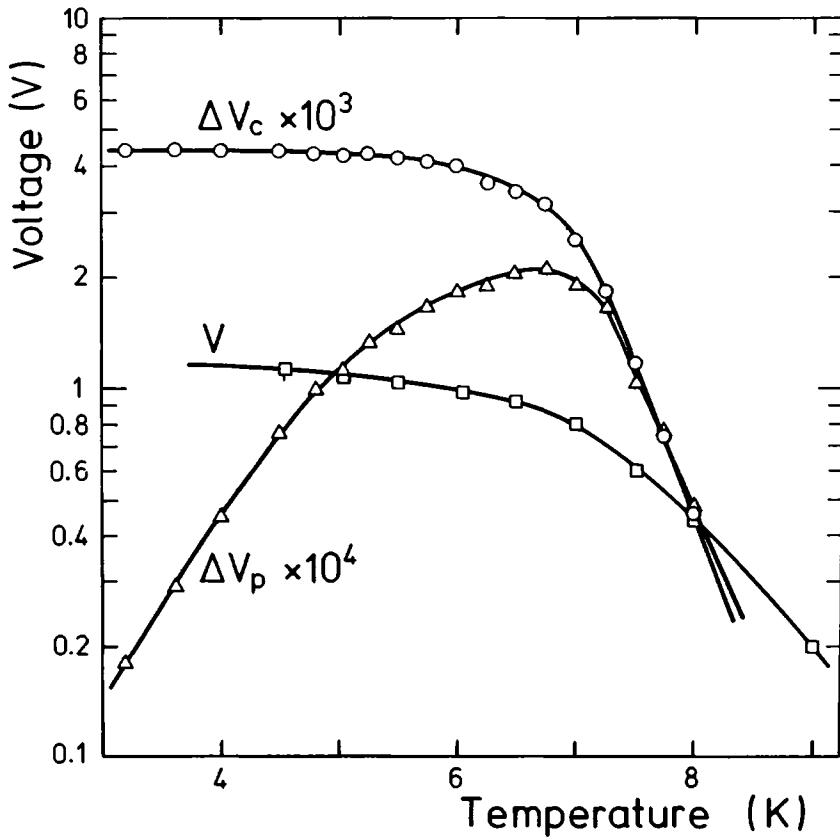


Fig. 2 : For a germanium sample with impurity concentration  $2.1 \times 10^{11}$  atoms/cm<sup>3</sup> are shown as a function of temperature: the voltage drop  $V$  across the sample, the contribution  $\Delta V_c$  of the continuum to the photoconductive signal, and the contribution  $\Delta V_p$  of the peak belonging to the C-transition of the boron spectrum.

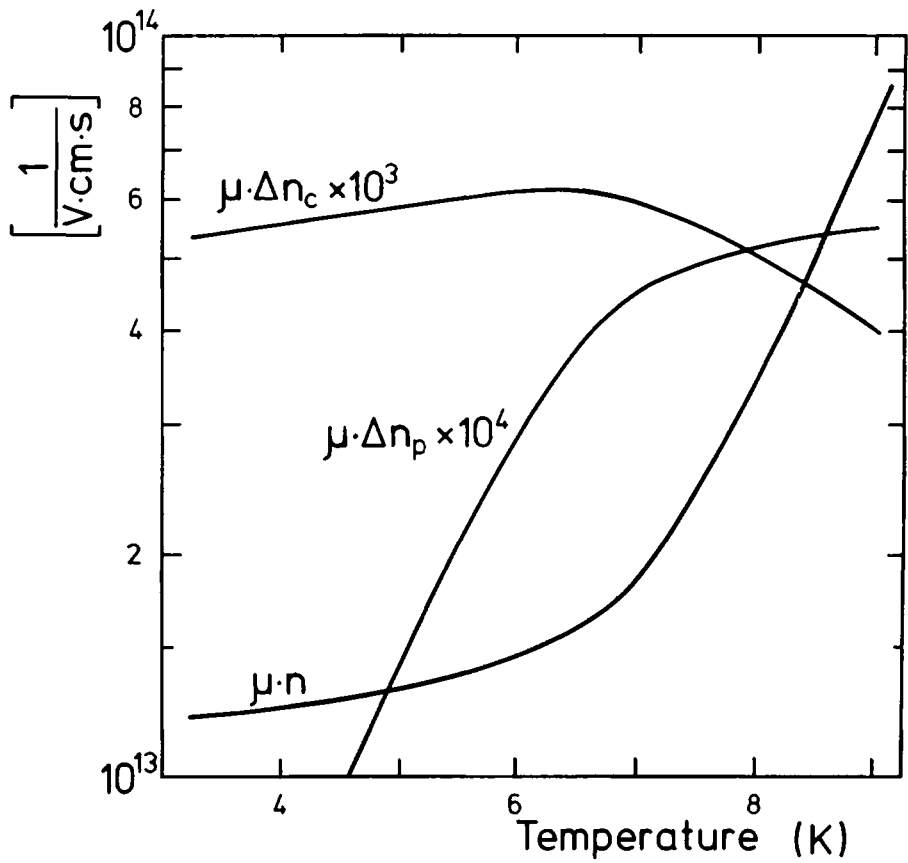


Fig. 3 : The values for  $\mu n$ ,  $\mu \cdot \Delta n_c$  and  $\mu \cdot \Delta n_p$  as a function of temperature calculated from figure 2 with equations (1) and (2).

Here  $\mu$  is the carrier mobility,  $n$  the carrier concentration,  $\Delta n_p$  the variation of the latter by photothermal ionization via an excited state and  $\Delta n_c$  its variation by direct photo-ionization.



the continuum. In figure 2 this is shown as a function of temperature for the C-transition peak and continuum of the boron spectrum from a sample with impurity concentration  $2.1 \times 10^{11}$  atoms/cm<sup>3</sup>. Also shown in figure 2 is the d.c. voltage-drop  $V$  across the sample. It is seen that  $\Delta V_c$  decreases very rapidly above 7 K whereas  $V$  decreases more slowly. The peak strength  $\Delta V_p$  increases with temperature as expected, but above 7 K  $\Delta V_p$  decreases along with  $\Delta V_c$ .

The values for  $\mu n$ ,  $\mu \cdot \Delta n_c$  and  $\mu \cdot \Delta n_p$  can now be calculated from the measured  $V$ ,  $\Delta V_c$  and  $\Delta V_p$  with equations (1) and (2); the results are presented in figure 3. The precise temperature dependence of  $\mu$  is not known but according to Debye and Conwell (1954) typical values for  $\mu$  lie between  $10^5$ - $10^6$  [cm<sup>2</sup>/V.s] in this temperature range and for this impurity concentration. As the "continuum" carrier density variation  $\Delta n_c$  is expected to be independent of temperature, the shape of the  $\mu \cdot \Delta n_c$  curve in figure 3 will reflect the temperature dependence of  $\mu$ .

We thus conclude from figure 3 that  $\Delta n_c \approx 10^5$  [cm<sup>-3</sup>] and that the strongly temperature dependent value of  $\Delta n_p$  is of the order of  $10^4$  [cm<sup>-3</sup>]. The mean free carrier concentration  $n$  approaches an asymptotic value for low temperatures, probably due to ionization by background radiation. For temperatures above 7 K  $n$  increases because of thermal ionization. In the considered temperature region the value of  $n$  is of the order of  $10^8$  [cm<sup>-3</sup>]. We may conclude from all this that although  $\Delta n_c$  does not change with temperature, the photoconductive signal (i.e.  $\Delta V_c$  and  $\Delta V_p$ ) decreases very rapidly above 7 K due to the increase of  $n$ , according to equation (2). Furthermore we see that  $n$ ,  $\Delta n_c$  and  $\Delta n_p$  are very small compared with the impurity concentration ( $2.1 \times 10^{11}$  atoms/cm<sup>3</sup>), so that practically

most impurities remain in their ground states. The temperature dependence of the photothermal ionization process can now be studied by comparing the measured peak strength  $\Delta V_p$  at various temperatures with the area of the continuum  $\Delta V_c$ , so by studying:

$$I = \Delta V_p / \Delta V_c = \Delta n_p / \Delta n_c .$$

In references II and III where equation (5) was already propounded, it was shown that experimental results on boron, aluminum and phosphorus dopes in high purity germanium were in excellent agreement with this equation. This is illustrated in figure 4 for two lines of the phosphorus spectrum. These results justify the assumption made about  $g'$  in section 3. The line intensities were taken from spectra which were normalized to the same continuum area. The solid curves in this figure represent least square computer fits of equation (5) to the experimental data, using  $\Delta E$  and  $A$  as adjustable parameters. The degeneracy  $g$  of the intermediate levels was deduced from measured level splittings due to a magnetic field (Bassani et al 1974). The numerical results of these fits for  $\Delta E$  are presented in table 1, together with the ground state energies ( $E_{g.s.}$ ) calculated as:

$$E_{g.s.} = E_{trans} + \Delta E . \quad (6)$$

The values for  $E_{g.s.}$  found in this way from the various spectral lines are consistent with each other within experimental accuracy and, as is shown in reference III, agree with data obtained by other means. As an illustration that the simple exponential behaviour

$$I \propto \exp(-\Delta E/kT) , \quad (7)$$

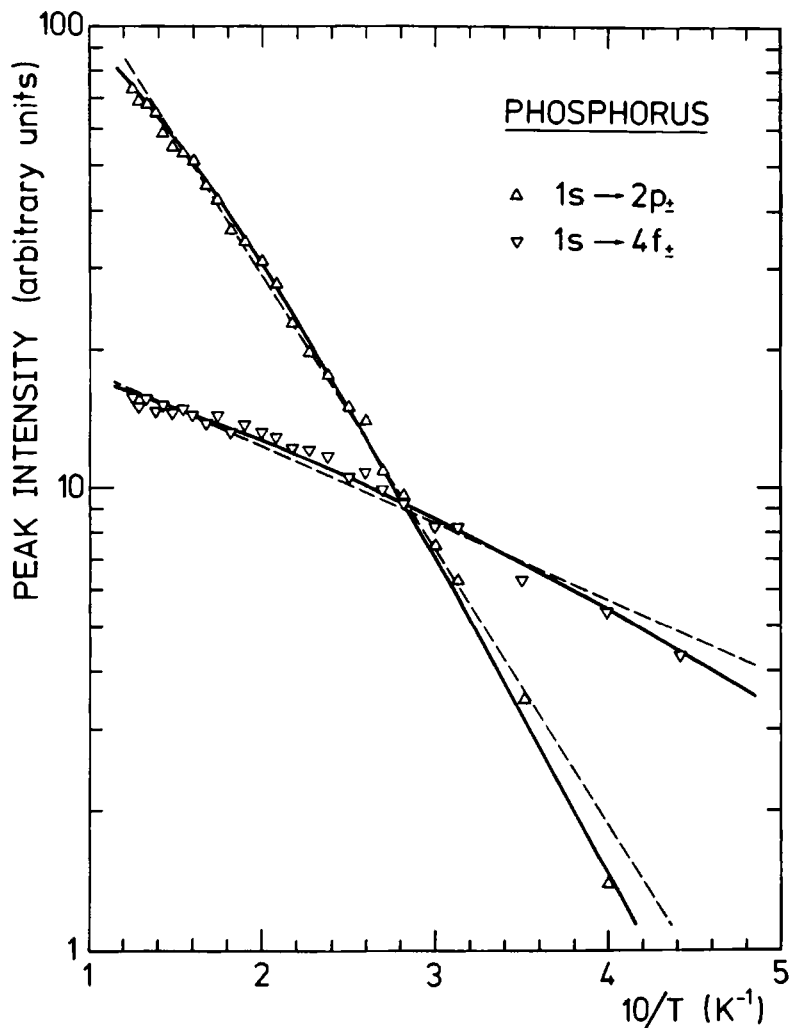


Fig. 4 : Logarithmic plot of the relative line intensities versus  $10/T$  for two phosphorus lines. The solid curves represent least square computer fits to equation (5), the dotted lines those to equation (7). (See also table 1.)

Table 1 : Values of  $\Delta E$  and  $E_{g.s.}$  as obtained from computer fits of equations (5) and (7) from the temperature dependence of two lines of the phosphorus spectrum.

transition	$E_{trans} (cm^{-1})$	$g$	With Eq. (5)		With Eq. (7)	
			$\Delta E (cm^{-1})$	$E_{g.s.} (cm^{-1})$	$\Delta E (cm^{-1})$	$E_{g.s.} (cm^{-1})$
$1s \rightarrow 2p_{+}$	$90.0 \pm 0.2$	4	$12.9 \pm 0.2$	$102.9 \pm 0.2$	$9.5 \pm 0.4$	$99.5 \pm 0.4$
$1s \rightarrow 4f_{+}$	$99.0 \pm 0.2$	2	$4.0 \pm 0.3$	$103.0 \pm 0.3$	$2.6 \pm 0.5$	$101.6 \pm 0.5$

frequently assumed ad hoc in the literature, does not yield very reliable results equation (7) was also fitted to the experimental data. The results are shown in figure 4 (dotted lines) and in table 1. As can be seen from figure 4 the fit is not as good as with equation (5) and the calculated values for  $E_{g.s.}$  are not consistent with each other.

Summarizing we may conclude that the theory about the line intensity in a photoconductivity spectrum, as outlined above, is in excellent agreement with the experimental results.

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ABSTRACT

Photothermal conductivity spectra in the far infrared ( $10 - 200 \text{ cm}^{-1}$ ) of boron doped germanium (dopant concentration  $\sim 10^{10} \text{ atoms/cm}^3$ ) have been obtained in magnetic fields  $B$  up to  $2 \text{ T}$  ( $\vec{B} // [001]$ ) at a temperature of  $T = 7.5 \text{ K}$ . The coefficients of the linear and quadratic field dependence of the Zeeman splittings and the resulting  $g$  values of the associated levels have been determined for the well known D and C transitions, and for the B and several A and I transitions as well. Moreover the field dependence has been studied of transitions to final states which could be identified as  $I_6$  like states associated with light-hole Landau levels. The results are discussed in terms of existing theories.

## 1. Introduction.

From a theoretical point of view, transitions between impurity states of semiconductors like GaAs and InSb are easier to discuss than those in Ge and Si because of the less complicated band structure of the former materials. However, in Ge and Si interactions between the impurity states are easier to avoid since impurity concentrations of  $\sim 10^{10}$  atoms/cm<sup>3</sup> can be achieved, while the lowest impurity concentration in a material like GaAs is of the order of  $10^{14}$  atoms/cm<sup>3</sup>, where interactions cannot be excluded (Larsen 1976, Stoelinga et al. 1978).

Impurity states have been studied mostly by measuring the transmission in the far infrared region (see the review article by Bassani et al. 1974). Soepangkat and Fisher (1973, further referred to as SF) employed this technique to investigate the Zeeman effect of lower levels of the boron and thallium impurities in germanium. For this type of measurements, relatively high impurity concentrations are needed.

By measuring the photothermal conductivity, samples of much higher purity can be used, yielding very narrow line widths of the spectra. This method was used by Lifshits and Nad (1965) for the first time and a recent review of this subject was given by Kogan and Lifshits (1977).

In a set of previous papers (Jongbloets et al. 1977, 1979 and 1980, further referred to as I, II and III) it has been shown how very accurate values for the ground state energies of the impurities can be deduced from the temperature dependence of the peaks in the spectra, obtained with the technique of photothermal ionization spectroscopy. Broeckx et al. (1979, further referred to as BCSV)



used this technique to investigate the Zeeman effect of the D and C transitions of the aluminum acceptor in germanium. Here, experimental results will be presented, not only for the D and C transitions, but also for the B and several A and I transitions of the boron acceptor in germanium. The labeling of the transitions is the same as used by Haller and Hansen (1974).

The Zeeman term connected with the magnetic field produces a linear splitting of the levels, whereas the diamagnetic term gives a quadratic correction. In addition, the impurity potential produces a set of sublevels associated with each Landau level (for detailed references, see the review article of Bassani et al. 1974).

In this paper, the splitting of one of the zero-field levels into sublevels associated with Landau levels is studied. Taking into account the magnetic field dependence of this zero-field level, values of the magnetic field dependence of the light-hole Landau levels are obtained. This dependence is in excellent agreement with theoretical considerations.

## 2. Experimental details.

The measurements were performed on a germanium sample containing  $2.1 \times 10^{11}$  atoms/cm<sup>3</sup> boron and  $4.3 \times 10^{10}$  atoms/cm<sup>3</sup> aluminum. The photothermal conductivity spectrum of the sample was measured as function of the wave number of the far infrared excitation radiation, using a Grubb Parsons Cube interferometer. The position of the peaks in the spectra were determined with an accuracy of  $0.2 \text{ cm}^{-1}$ . A magnetic field of up to 2 Tesla could be applied with a superconducting solenoid (Oxford Instruments). Most of the experimental details were given in reference II.

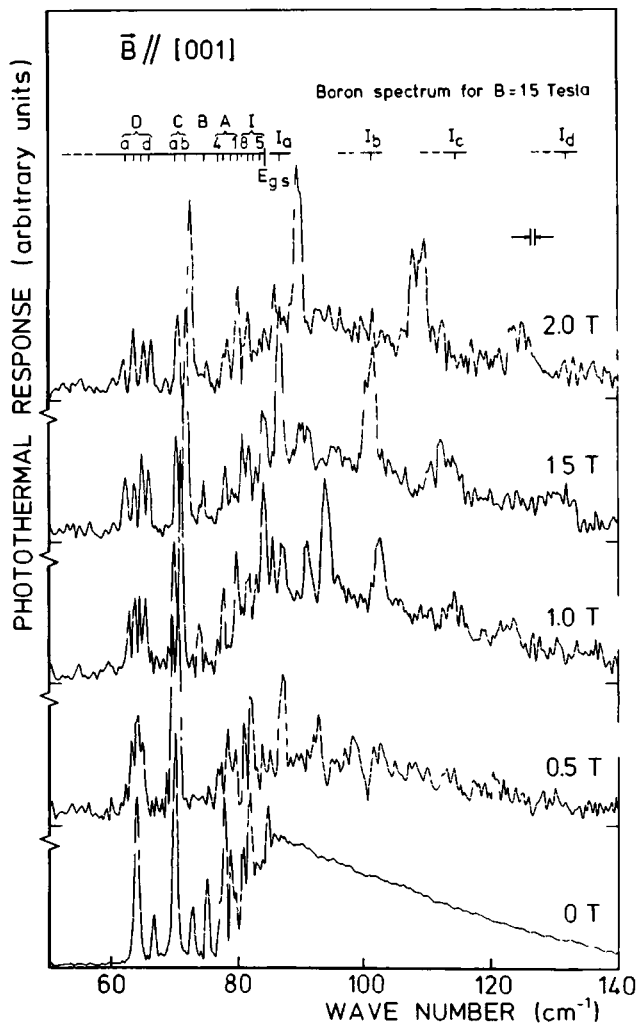


Fig. 1 : Photothermal conductivity spectra at 7.5 K of shallow acceptors in germanium for several values of the magnetic field  $\vec{B} // [001]$ . Impurity concentrations:  $2.1 \times 10^{11}$  at./cm<sup>3</sup> B,  $4.3 \times 10^{10}$  at./cm<sup>3</sup> Al. The labelling at the top shows the boron spectrum at 1.5 Tesla including the Zeeman splitting and four satellite peaks of the  $I_c$  transition associated with Landau levels.

Figure 1 shows some of the measured spectra for several values of the magnetic field with  $\vec{B} // [001]$ . The sample was held at a temperature of 7.5 K. The labeling of the lines of the boron spectrum at 1.5 Tesla is shown in the upper part of figure 1. The zero-field band edge corresponds with the ground state binding energy  $E_{g.s.}$ , as determined by the method outlined in references II and III. The continuum above  $E_{g.s.}$  in the zero-field spectrum is due to direct photo-ionization from the impurity ground state into the free band, whereas the peaks below  $E_{g.s.}$  are caused by the two-step photothermal ionization process. It is clearly visible in figure 1 that on applying a magnetic field the spectra not only exhibit a splitting or shift of the peaks below  $E_{g.s.}$ , but also develop a set of peaks in the continuum. These two types of peaks will be discussed separately in the following sections.

### 3. Field dependence of the peaks below $E_{g.s.}$

In figure 2 the transition energies of the peaks with wave numbers below the continuum onset  $E_{g.s.}$  are plotted as a function of the magnetic field. As a linear and a quadratic Zeeman term is expected for each transition, a quadratic polynomial

$$E(B) = E(0) + aB + bB^2 \quad (1)$$

was fitted to the experimental data. The solid curves in figure 2 calculated in this way, show an excellent fit to the data. The parameters  $a$  and  $b$  for the various spectral lines are summarized in table 1.

A similar procedure was applied by SF to the D and C lines of boron, and by BCSV to the D and C lines of aluminum. BCSV used

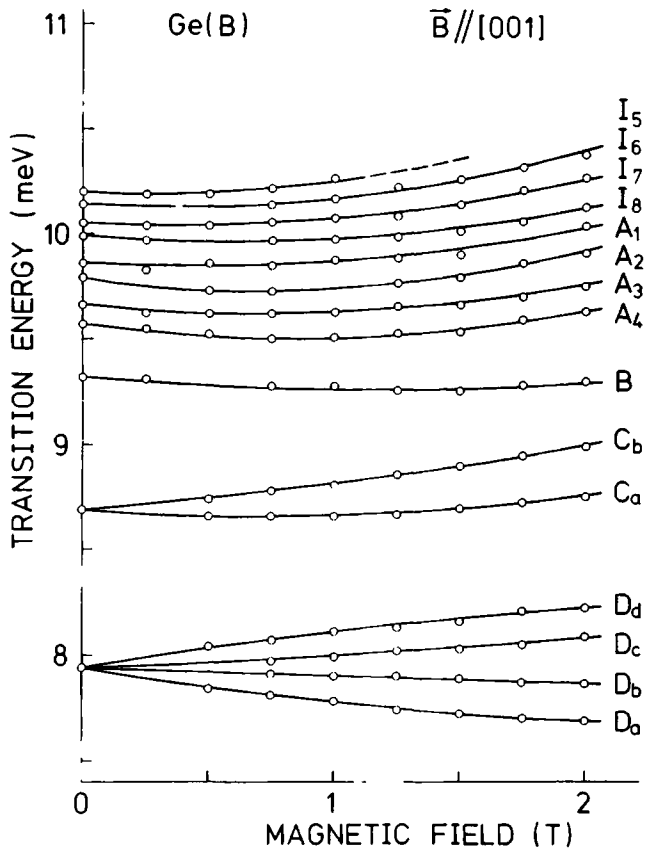


Fig. 2 : Zeeman effect in the photothermal conductivity spectrum of the boron acceptor in germanium with  $\vec{B} // [001]$ . The curves are least square fits of a quadratic polynomial to the experimental data. (See also table 1.)

Table 1 : Linear and quadratic Zeeman parameters for the spectral lines of the boron impurity in germanium with  $\vec{B} // [001]$  as determined from figure 2.

Transition	a (meV/T)	b (meV/T <sup>2</sup> )
D <sub>a</sub>	-0.205 $\pm$ 0.005	0.031 $\pm$ 0.004
D <sub>b</sub>	-0.041 $\pm$ 0.005	0.003 $\pm$ 0.002
D <sub>c</sub>	0.050 $\pm$ 0.005	0.020 $\pm$ 0.004
D <sub>d</sub>	0.198 $\pm$ 0.009	-0.027 $\pm$ 0.006
C <sub>a</sub>	-0.081 $\pm$ 0.005	0.061 $\pm$ 0.004
C <sub>b</sub>	0.094 $\pm$ 0.003	0.032 $\pm$ 0.001
B	-0.094 $\pm$ 0.009	0.042 $\pm$ 0.006
A <sub>4</sub>	-0.17 $\pm$ 0.01	0.103 $\pm$ 0.006
A <sub>3</sub>	-0.117 $\pm$ 0.007	0.086 $\pm$ 0.006
A <sub>2</sub>	-0.175 $\pm$ 0.005	0.124 $\pm$ 0.004
A <sub>1</sub>	-0.07 $\pm$ 0.01	0.08 $\pm$ 0.01
I <sub>8</sub>	-0.09 $\pm$ 0.01	0.082 $\pm$ 0.007
I <sub>7</sub>	-0.057 $\pm$ 0.007	0.084 $\pm$ 0.005
I <sub>6</sub>	-0.079 $\pm$ 0.007	0.108 $\pm$ 0.006
I <sub>5</sub>	-0.07 $\pm$ 0.01	0.12 $\pm$ 0.01

polarized radiation in the Faraday configuration with the polarization vector  $\vec{E} \perp \vec{B}$ , SF used the Voigt configuration with  $\vec{E} \perp \vec{B}$  and  $\vec{E} // \vec{B}$ . So the relative intensities of the peaks within one multiplet could be used for the identification of the transitions (see Bhattacharjee and Rodriguez, 1972). The experimental results of SF and BCSV for  $\vec{B} // [001]$  are collected in table 2. The corresponding notations for the transitions, used by BCSV and SF, and in the present paper are also given there.

In this experiment the sample was mounted inside a semispherical integrating cavity; therefore it was not possible to use polarized radiation. This was not a serious problem, however, because the only lines seen to be splitting into multiplets with  $\vec{B} // [001]$  are the D and C transitions, for which the identification was already given by BCSV. Comparing table 1 and 2 a good agreement can be seen between the present values for the linear and quadratic Zeeman terms for the D and C transitions, and those of SF and BCSV.

For the higher transitions (i.e. B, A and I lines) no comparable experimental values are available from the literature. Comparison with theory is also not possible because, as can be seen from figure 2, the theoretical assumption (Bhattacharjee and Rodriguez, 1972) that the unperturbed states at  $B = 0$  are sufficiently well separated to treat the splitting of each state separately, does not hold for the higher lines any longer. The violence of this assumption for the A and I-lines can be seen from the nearly similar values for  $b$ , which is a consequence of the mutual repulsion of these lines because of their narrow spacing. A common feature of the B, A and I lines is the negative sign for the values of  $a$ .

Table 2 : Linear and quadratic Zeeman parameters  $a$  and  $b$  with  $\vec{B} // [001]$  for the D and C transition of the Al acceptor in Ge, as determined by BCSV (Faraday configuration,  $\vec{E} \perp \vec{B}$ ), together with the values for the B acceptor in Ge, as determined by SF (Voigt configuration,  $\vec{E} \perp \vec{B}$  and  $\vec{E} // \vec{B}$ ).

This work Trans.	BCSV values for Ge(Al)			SF values for Ge(B)		
	Trans.	$a$ (meV/T)	$b$ (meV/T <sup>2</sup> )	Trans.	$a$ (meV/T)	$b$ (meV/T <sup>2</sup> )
$D_a$	$D_4$	-0.208	0.028	$D_1$	-0.176	0.015
				$D_2$	-0.163	0.014
$D_b$	$D_3$	-0.054	0.009	$D_3$	-0.045	0.005
				$D_4$	-0.030	0.002
$D_c$	$D_2$	0.055	0.017	$D_6$	0.057	0.019
				$D_5$	0.046	0.021
$D_d$	$D_1$	0.219	-0.030	$D_8$	0.185	-0.017
				$D_7$	0.187	-0.020
$C_a$	$C_{4,3}$	-0.066	0.062	$C_2$	-0.048	0.052
				$C_1$	-0.064	0.055
$C_b$	$C_{2,1}$	0.120	0.033	$C_3$	0.095	0.038
				$C_4$	0.107	0.037

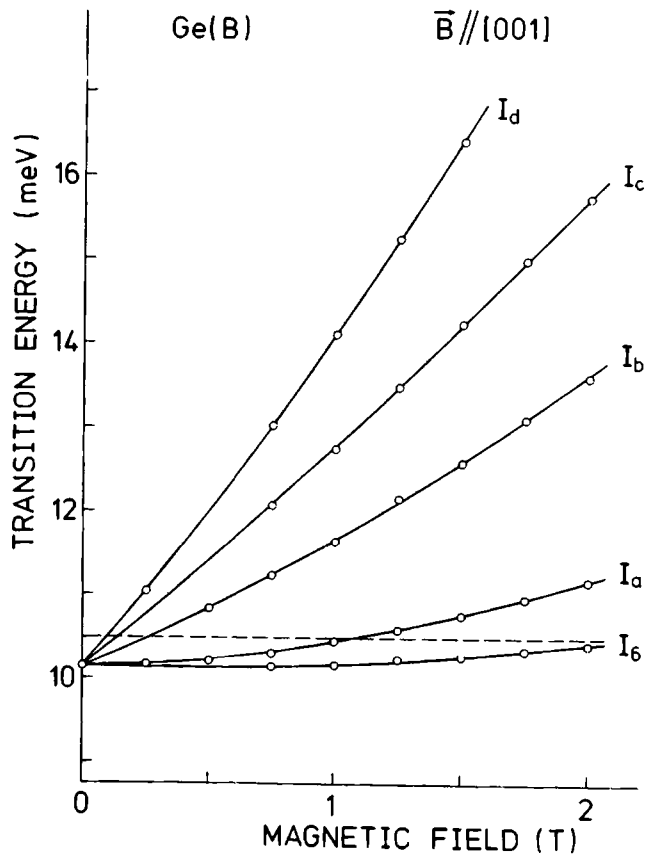


Fig. 3 : Magnetic field dependence in the photoconductivity spectra of the  $I_{6\bullet}$  transition and four satellite transitions to bound states associated with Landau levels. The dashed line indicates the zero-field band edge. The curves are least square fits of a quadratic polynomial to the experimental data. (See table 3.)



#### 4. Field dependence of the peaks in the continuum.

Apart from the peaks below the band edge, the photoconductivity spectra (figure 1) show a great number of peaks in the continuum that are almost impossible to identify unambiguously. However, some very pronounced peaks appear (labelled  $I_a$ ,  $I_b$ ,  $I_c$  and  $I_d$ ) that can easily be recognized in every spectrum. The magnetic field dependence of these peaks is shown in figure 3, together with the  $I_6$  transition. Again a quadratic polynomial was fitted to the experimental points, i.e.

$$E_n(B) = E_n(0) + p_n B + q_n B^2, \quad n = a, b, c, d. \quad (2)$$

A very good fit can be obtained if the zero-field energy intercepts  $E_n(0)$  of the curves coincide with the zero-field  $I_6$  transition energy; this has been plotted in figure 3. The parameters  $p_n$  and  $q_n$  of these fits are summarized in table 3.

This allows the conclusion that the  $I_a$ ,  $I_b$ ,  $I_c$  and  $I_d$  transitions are of the same type as  $I_6$ , i.e. transitions from the impurity ground state to impurity excited states rather than valence band levels. Kaplan (1968) observed similar features in transmission spectra of boron in germanium and attributed the excited states tentatively to light-hole Landau states. The present very accurate experimental data support this conclusion (section 6).

#### 5. Zeeman effect.

Soepangkat and Fisher (1973) interpreted their absorption spectra of the boron impurity in germanium under the assumption that the Zeeman splitting of the ground level is of the same order of magnitude as that of the final state of the D transition. However,

Table 3 : Linear and quadratic fit parameters for the magnetic field dependence of the  $I_a$ ,  $I_b$ ,  $I_c$  and  $I_d$  transitions of the boron impurity in germanium with  $\vec{B} // [001]$  (see figure 3). Also listed are the differences between these parameters and those of the  $I_6$  transition.

Transition	$p_n$ (meV/T)	$q_n$ (meV/T <sup>2</sup> )	$p_n - a(I_6)$ (meV/T)	$q_n - b(I_6)$ (meV/T <sup>2</sup> )
$I_a$	$0.11 \pm 0.01$	$0.20 \pm 0.01$	$0.19 \pm 0.02$	$0.09 \pm 0.01$
$I_b$	$1.28 \pm 0.03$	$0.24 \pm 0.03$	$1.35 \pm 0.03$	$0.13 \pm 0.03$
$I_c$	$2.46 \pm 0.04$	$0.16 \pm 0.03$	$2.53 \pm 0.04$	$0.05 \pm 0.03$
$I_d$	$3.47 \pm 0.04$	$0.47 \pm 0.04$	$3.55 \pm 0.04$	$0.36 \pm 0.04$

Table 4 : Calculated parameters for the linear Zeeman effect of the final state of the D and C transitions with  $\vec{B} // [001]$  for Ge(B). Also shown are the values of BCSV for Ge(Al).

		$g'_{1/2}$	$g'_{3/2}$	$r$	$g'_1$
D {	this work	$7.0 \pm 0.2$	$-0.52 \pm 0.04$	$-0.120 \pm 0.005$	$7.9 \pm 0.2$
	BCSV	$7.4 \pm 0.2$	$-0.63 \pm 0.07$	$-0.119 \pm 0.006$	$8.4 \pm 0.2$
C {	this work	$3.0 \pm 0.1$	$1.01 \pm 0.04$	$-0.077 \pm 0.004$	$3.3 \pm 0.1$
	BCSV	$3.2 \pm 0.2$	$1.07 \pm 0.07$	$-1/13$	$3.5 \pm 0.2$

Tokumoto and Ishiguro (1977) concluded from magnetoacoustic resonance attenuation measurements on Ga-doped Ge that the splitting of the ground state of shallow acceptors in Ge is much smaller than previously assumed. BCSV pointed out that the ground state splitting is in fact so small that the observed line-splittings in photothermal conductivity spectra would reflect the splittings in the final state of the transitions only. This simplifies greatly the interpretation of the spectra, BCSV proposed the assignment of both the D and C lines to  $\Gamma_8^+ \rightarrow \Gamma_8^-$  transition. We will proceed along the same line and compare our results for the Zeeman effect of the D and C transitions of boron in germanium with  $\vec{B}/[001]$  with the results of BCSV for aluminum.

Using symmetry considerations Bhattacharjee and Rodriguez (1972) have shown that with  $\vec{B}/[001]$  the energy shifts of the Zeeman levels of a  $\Gamma_8$  state are given by

$$\Delta E_{\mu} = \mu_B (g_1' \mu + g_2' \mu^3) B + [q_1 + (q_2 + q_3) \mu^2] B^2, \quad (3)$$

where  $\mu = 3/2, 1/2, -1/2, -3/2$ , and  $\mu_B$  is the Bohr magneton. Introducing the parameter  $r = g_2'/4g_1'$  which measures the cubic anisotropy, and the principal g-factors  $g_{1/2}'$  and  $g_{3/2}'$  defined by

$$\begin{aligned} g_{1/2}' &= g_1' (1 + r), \\ g_{3/2}' &= g_1' (1 + 9r), \end{aligned} \quad (4)$$

the linear part of equation (3) can be written as:

$$\begin{aligned} \Delta E_{\pm 1/2} &= \pm 1/2 g_{1/2}' \mu_B B \\ \Delta E_{\pm 3/2} &= \pm 3/2 g_{3/2}' \mu_B B. \end{aligned} \quad (5)$$

So the splitting of a  $\Gamma_8$  level will be symmetrical relative to the zero-field energy for  $\mu = -1/2, +1/2$  and  $\mu = -3/2, +3/2$ .

BCSV assign to the D multiplet with  $\vec{B} // [001]$  the transitions:  $D_a = \Gamma_8 \rightarrow -1/2$ ,  $D_b = \Gamma_8 \rightarrow +3/2$ ,  $D_c = \Gamma_8 \rightarrow -3/2$  and  $D_d = \Gamma_8 \rightarrow +1/2$ . With this assignment the principal g-factors  $g'_{1/2}$  and  $g'_{3/2}$  can be deduced from the measured parameters  $a$  with equation (5), and from this  $g'_1$  and  $r$  can be calculated with equation (4). The results are shown in table 4, together with the results of BCSV. For the C multiplet BCSV give the assignment  $C_a = \Gamma_8 \rightarrow -3/2, -1/2$  and  $C_b = \Gamma_8 \rightarrow +3/2, +1/2$ . This leads to the g-factors for the C-final state which are also summarized in table 4. As can be seen from table 4 there is a good agreement between our values and those of BCSV.

## 6. Landau levels.

In a uniform magnetic field  $\vec{B}$  the continuum of states in the energy band coalesces into a set of discrete quantum states, the Landau levels. Usually these Landau levels form a "ladder" of equidistant energies with spacing  $\hbar\omega_c$  defined by the classical cyclotron frequency  $\omega_c = eB/m^*$ ,  $m^*$  being the effective mass of the charge carrier. However, for degenerate bands such as the valence band edge in Ge the situation is more complex. Using the quantum-mechanical effective mass formalism Luttinger (1956) predicted that the lowest Landau levels would have irregular spacings, while the higher levels tend to uniform intervals. Numerical values for the Landau energy levels in Ge were derived by Evtuhov (1962) from second order perturbation theory.

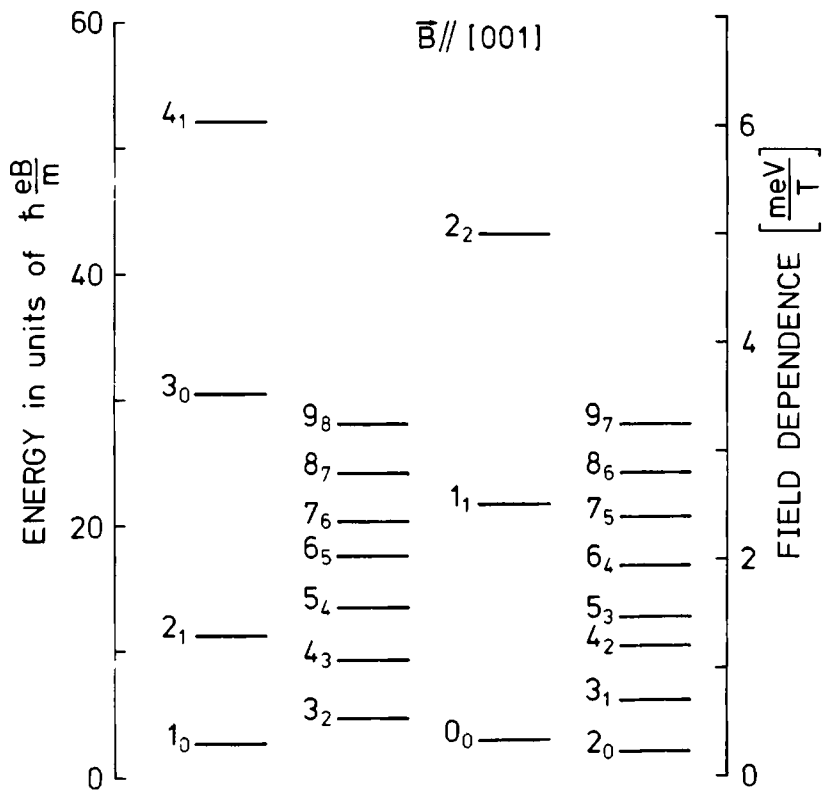


Fig. 4 : Valence-band Landau levels in germanium with  $\vec{B} // [001]$ , as calculated by Hensel and Suzuki (1974).

Very extensive calculations have been carried out by Hensel and Suzuki (1974) who applied uniaxial stress to the Ge crystal as a means to uncouple the valence bands in order to decipher the complicated quantum spectra. They found a separate Landau ladder for every value of the magnetic quantum number  $M_J$ , i.e. two heavy hole ladders ( $M_J = \pm 1/2$ ) and two light hole ladders ( $M_J = \pm 3/2$ ). Their results for  $\vec{B}/[001]$  are plotted in figure 4.

Boyle and Howard (1961) and Horii and Nisida (1971) have shown that in Zeeman transmission spectra of Ge donors every bound state with energy  $E(B)$  has a set of satellite states associated with conduction band Landau levels. The energy  $E_n(B)$  of these satellite states is given by

$$E_n(B) = E(B) + L_n(B) ,$$

where  $L_n(B)$  is the energy of the  $n$ -th Landau level relative to the zero-field band edge. Applying the same idea to acceptor spectra, the valence band Landau levels can be deduced from the experimental data of figure 3 and table 2 by subtracting the linear and quadratic fit parameters of the  $I_6$  transition from those of the  $I_a$ ,  $I_b$ ,  $I_c$  and  $I_d$  transition i.e.

$$L_n(B) = [p_n - a(I_6)] \cdot B + [q_n - b(I_6)] \cdot B^2 \quad n = a, b, c, d .$$

The results are listed in table 3. A comparison of the values of the linear field dependence with the calculations of Hensel and Suzuki (figure 4), yields that the value of  $(p_a - a(I_6))$  for  $L_a$  has a fair agreement with Landau levels  $l_0$  and  $0_0$ , whereas the corresponding values for  $L_b$ ,  $L_c$  and  $L_d$  agree excellently with the Landau levels

$2_1$ ,  $1_1$  and  $3_0$  respectively. So apparently the spectral lines  $I_a$ ,  $I_b$ ,  $I_c$  and  $I_d$  can be assigned to transitions from the boron ground state to  $I_6$ -like final states associated with light-hole Landau levels.

In conclusion it should be emphasized that the technique of photothermal ionization spectroscopy not only provides an excellent method for the study of the Zeeman effect of shallow impurities in germanium, but also for the analysis of the Landau level structure of the energy bands. It is shown that the results obtained with this technique are in excellent agreement with existing theories.

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This appendix shows the computer program, that has been developed for the processing of the data from the interferometer equipment, described in Chapter II. It is intended for handling phase-modulated interferograms, but only minor modifications are necessary for amplitude-modulation. The program is written in assembler language for a PDP-12 (DEC) laboratory computer. But it is also suited for a PDP-8 system; only the instructions for reading and writing of magnetic tapes have to be adapted.

The program is divided into two parts:

1. The main program "PHASE2", the structure of which is shown in Fig. A.1.
2. "FIRUTIL", a series of utility subroutines, such as:
  - a software floating point system,
  - a subroutine to type text strings,
  - an incremental plotter subroutine,
  - a subroutine to plot alphanumeric characters,
  - a subroutine to read papertapes and convert codes,
  - a subroutine to read or write magnetic tapes.

The program "PHASE2" reads interferograms, which are punched on paper-tape. "PHASE2" can handle a single interferogram, or a combination of a sample- and background-interferogram. These sample- and background-interferograms can be registered as separate measurements, but also simultaneously with a two-channel system. "PHASE2" has also the possibility to transform a series of interferograms and average the resulting spectra.

Each interferogram is preceded by control data, such as step length and number of channels used, and a series of 30 BCD characters defined by the operator with thumbwheel switches on the scanner-control unit. These 30 characters are used to provide the program with an identification number for the spectrum, and information about the sensitivity of the detection system. This last information is necessary to calculate the correct average and ratio of spectra.

The maximum number of interferometric data points that can be transformed is  $N = 1024$ . The interferogram should ideally be double sided with approximately an equal number of data points on either side of zero path. However, there is no need to have exactly the correct amount as the program is arranged to either truncate or pad the data as necessary.

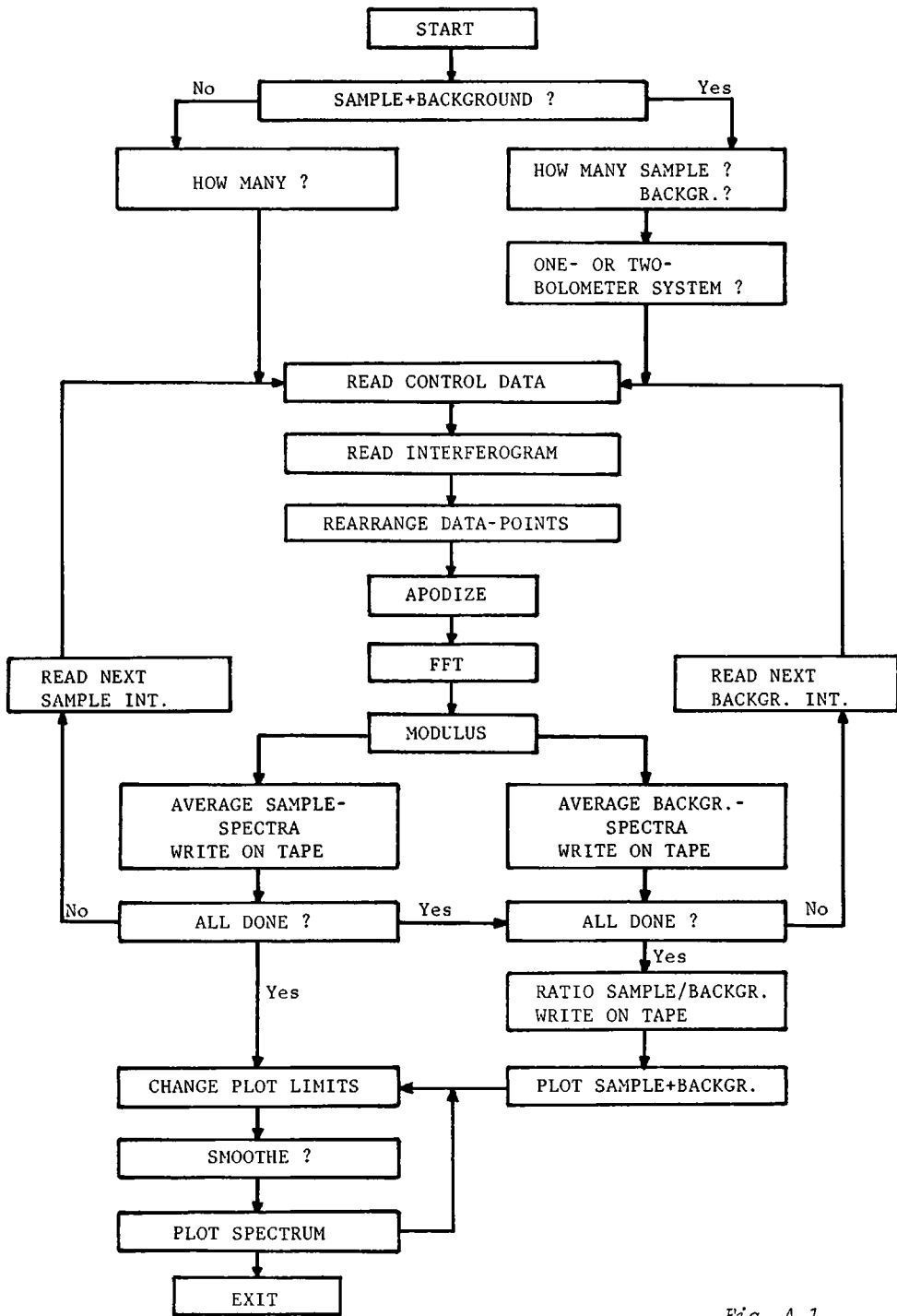


Fig. A.1

After reading the entire interferogram the maximum positive and negative values of the data are located, and the data point in between is attributed to zero path difference. The average value of the entire data array is subtracted from every data point to remove a possible D.C. offset present in the interferogram. The data array is then padded out to the maximum size of  $N = 1024$  (if necessary) with zero's and then re-located such that the order is undisturbed but zero path is in the first location. The data array is then apodized from both ends towards the middle with a "cosine squared" function, after which the program calls for the "Fast Fourier Transform" (FFT). The FFT subroutine is based on a program written by P.L. Walton<sup>(1)</sup>, which employs a modification of the Cooley-Tukey fast Fourier transform algorithm<sup>(2)</sup>. The FFT program calculates the complex spectrum

$$P(n \cdot \Delta\sigma) = \frac{1}{N} \sum_{m=0}^{N-1} I(m \cdot \Delta x) \cdot \exp\left(-\frac{2\pi i}{N} \cdot n \cdot m\right), \quad n = 0, 1, \dots, N/2.$$

The spectrum points are equidistant with a frequency-interval  $\Delta\sigma = (N \cdot \Delta x)^{-1}$ , where  $\Delta x$  is the sampling interval of the interferogram points. Then the modulus-spectrum is calculated as

$$|P(n \cdot \Delta\sigma)| = \sqrt{[\operatorname{Re}(P)]^2 + [\operatorname{Im}(P)]^2},$$

and stored on magnetic tape.

If a number of spectra must be averaged, the relevant interferograms are now read one by one, transformed and averaged with the already stored spectrum. The same procedure is then repeated for possible background-spectra, and the ratio of the (averaged) sample- and background-spectrum is calculated. Finally the resultant spectrum is plotted on an incremental plotter, within wave number limits chosen by the operator. Noisy spectra can be improved somewhat (at the cost of resolution) by a mathematical smoothing procedure, and then plotted again.

### References.

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[illegible]

0365	5174	JMS SAMPLE1	0.46	3074	DIA MAXLOH	/LOCATION OF MAX NUM
0372	0607	CHCKN1 CHECK	0547	4407	FENTER	
0373	4772	/READ SAMPLE INTERFEROGRAM.	0.20	5120	FGET 1	
0374	1036	SAMPLE, JMS 1 (CHECK)	0551	510	FSUB 1 AJL	/5 A(J)
0375	3531	TAD TFRST	0552	0000	FEXT	
0376	4404	DCA 1 BKT5	0553	1160	TAD HORD	
0377	3533	WRTTE	0554	2720	SFA SNA CLA	/ 0 /
0378	4404	MS69	0555	5364	JMF +7	/N1
0400	4605	JMS 1 LOCK	0556	4407	FENTER	
0401	4210	JMS SKE9	0557	5110	FGET 1 AJL	
0402	4406	READ	0560	6170	FPUT 5	/YES 5 A(J)
0403	4406	READ	0561	0000	FEXT	
0404	7213	JMS DATA	0.61	1106	TAD 1	
0405	0700	LOCN1 DOKIN	0.63	3074	DCA MINLOH	/LOCATION OF MINIMUM
0406	1003	(DOKIN) CONTN	0.64	7240	SFA	
0407	7621	NUL 7612	0.65	1013	TAD NR	/IF NR 1 1 AJL
0410	0000	SKE9 0	0.66	7450	SNA	
0411	1217	TAD M9	0.67	1173	JMF +14	
0412	3156	DCA COUNT	0.70	1011	TAD KUN	
0413	4406	READ	0.71	2640	SFA CLA	
0414	1156	ISZ COUNT	0.72	4130	JMS SKE9	
0415	5213	JMF 2	0.73	7136	ISZ 1	
0416	5610	JMF 5 SKE9	0.74	713	CLA (1) MA 15	
0417	7767	M9 7767	0.75	106	TAD 1	
0420	0607	CHCKN1 CHECK	0.76	2640	SFA CLA	/ 10 5
		/READ BACK FOUND INTERFEROGRAM.	0.77	5267	JMS DATA	/NO LON INT
0421	7344	BKGA SFA (1) KAI	0600	4406	READ	/YES KAI KAI
0422	1023	TAD NR	0601	1157	AD END	
0423	7640	SFA CLA	0602	7640	SFA CLA	
0424	5244	JMF BACKGA	0603	00	JMS 1	
0425	4544	JMS 1 BACKGD	0604	7606	M 1 (ALCI	
0426	4406	READ	0605	7620	END 160	
0427	1055	TAD STORI	0606	1013	LOCN1 LOCN1	
0428	3156	DCA STORI			/READ AND CHECK BACKGA STEE	
0431	4406	READ	0607	0000	CHCKN1	
0432	3055	DCA 51 KAI	0610	4406	KAI	/READ WITH TAD
0433	1073	TAD 51 KAI	0611	1057	TAD MML	/IF 1 "NEW TAD"
0434	1207	TAD 51 KAI	0612	7240	SFA CLA	
0435	7640	SFA 4	0.61	1173	JMS 1	
0436	5777	JMS KAI	0614	4406	KAI	
0437	1156	TAD STORI	0615	7241	CLA	
0440	1057	TAD MML	0616	1023	TAD NR	/READ AND CHECK KAI
0441	7640	SFA CLA	0617	7240	SFA CLA	
0442	5727	JMF KAI	0618	0.43	JMS 1 KAI	
0443	4745	JMS 1 FORWARD	0619	4407	SNA	
0444	4620	BACKGA JMS 1 CHECK	0620	4407	INT 1	
0445	4710	JMS SKE9	0621	0011	INT 1	/READ AND CHECK TAD
0446	4406	READ	0622	0002	LOCN1	
0447	4406	READ	0623	45	FSUB TAD	
0450	4404	WRTTE	0624	0000	FEXT	
0451	1160	M9 1160	0625	1160	TAD HORD	
0452	4605	JMS 1 LOCK	0630	7640	SFA CLA	
		/READ INTERFEROGRAM.	0631	4535	JMS 1 BKG	
0453	7332	DATA1 CLA 51 KAI	0632	4406	READ	
0454	5123	DCA 1	0633	4107	FENTER	
0455	7333	CLA TAD STL RTR	0634	0011	FEXT	/KAI 1
0456	3124	DCA TAD	0635	0005	FEXT	
0457	3312	CLA STL RTR	0636	6115	FPUT 1	
0460	5120	DCA 5	0637	0011	INT 1	/SKE 6 BACK
0461	7337	CLA STL RTR	0640	0000	FEXT	
0462	3121	DCA 51	0641	0000	READ	
0463	3126	DCA 51	0642	4406	CLA (1)	
0464	3127	DCA 51	0643	1000	JMF 1 CHECK	
0465	3130	DCA 51	0644	607	STEE 5	
0466	3106	DCA 1	0645	0000	0	
0467	4606	DATA1 JMS 1 CONTRL	0646	0000	0	
0470	1106	TAD 1	0647	0000	0	
0471	4503	JMS 1 INDEX			/LOCK IN AMPLIFIED SETTINGS.	
0472	3110	DCA AJL	0648	0000	LOCKN1	
0473	1023	TAD NR	0649	4404	WRTTE	
0474	1011	TAD KUN	0650	4407	MS69	/N1
0475	7740	SFA SFA CLA	0651	4407	FENTER	
0476	4210	JMS SKE9	0652	5113	FPUT 1	
0477	4407	FENTER	0653	0012	INT 1	/KAI NR
0500	0011	INPUT	0654	0050	0	
0501	3005	5	0655	0000	TEXT	
0502	0000	FEXT	0660	4404	WRTTE	
0503	4406	READ	0661	3171	MS69	/SKE 1
0504	0041	CLA	0662	1027	AD 1 FEXT	/SKE 1
0505	0001	IAI	0663	4543	JMS 1 FLOAT	
0506	7640	SFA CLA	0664	4407	FENTER	
0507	5114	JMF +5	0665	6167	FPUT FFAI	
0510	4407	FENTER	0666	0011	INT 1	/SENSITIVITY
0511	0057	FPUT 10000	0667	0003	0	
0512	1052	FPUT 10000	0670	6120	FPUT 1	
0513	0000	FEXT	0671	0012	INT 1	
0514	4406	READ	0672	0050	0	
0515	704	IAI	0673	1470	FPUT 5	
0516	704	IAI	0674	4162	FPUT FFAI	/SENS 5 SKE 5
0517	7640	SFA CLA	0675	6044	FEXT	
0520	5774	JMF +4	0676	0000	TEXT	
0521	4407	FENTER	0677	4406	READ	
0522	0010	NEGATE	0678	0000	JMS 1 CONTRL	
0523	4406	READ	0701	1055	DCA STORI	
0524	4406	READ	0702	7344	SFA (1) KAI	/ 1
0525	4407	FENTER	0703	1055	TAD STORI	
0526	4077	FPUT 10000	0704	7200	SFA	/SCALE 1
0527	3044	JMS SKE9	0705	5113	JMF +4	
0530	6510	FPUT 1 AJL	0706	4404	WRTTE	/YES
0531	1156	FPUT 5	0707	3623	MS69	/NANO
0532	6126	FPUT 1	0710	5130	JMF +4.0	
0533	5510	FPUT 1 AJL	0711	7440	SFA	/CLA 2
0534	7123	FSUB 1	0712	316	JMF +4	
0535	0000	FEXT	0713	4404	WRTTE	/YES
0536	1160	TAD HORD	0714	7626	MS69	/MICRO
0537	7250	SFA SFA CLA	0715	5720	JMF +4	
0540	2347	JMF +7	0716	4404	WRTTE	/SCALE 1
0541	4407	FENTER	0717	7412	MS69	/MICRO
0542	5510	FPUT 1 AJL	0720	4404	WRTTE	
0543	6123	FPUT 1	0721	3636	MS69	/VOL 1
0744	0000	FEXT	0722	1055	TAD STORI	
0745	1106	TAD 1	0723	7041	IAI	
			0724	3156	DCA COUNT	

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0725 44C/ FENTER
0726 5044 FGE SEN
0727 3341 FMFY L1000
0730 6044 FMFUT SENS /SEN5 SEN5X(1000 SEN5SL)
0731 0000 FEX*
0732 2156 ISZ COUNT
0733 5325 JMF -4
0734 1346 TAD N7
0735 3156 DCA COUNT
0736 4406 REAL
0737 2 56 .SZ COUNT
0740 5336 JMP -2
0741 7300 CLA CLL
0742 5650 JMF I LOCKIN C1000, 00.2 /1000
0744 3720 1720
0745 0000 0000
0746 7771 N7, /PATC1 OR AFODISATION,
0747 0000 0
0750 1074 TAD MAXLOC
0751 7104 CLL RAL /-RM
0752 7041 CIA /O ?
0753 1104 TAD J /J-2RM
0754 7510 54A /NO ?
0755 7200 (LA /NO STORE M
0756 1074 TAD MAXLOC /YES, STORE M
0757 7041 CIA /YES, STORE M
0760 1105 TAD P5 J /M-512 OK (J M 512)
0761 7041 CIA /O ?
0762 7500 54A /NO, STORE M
0763 7200 (LA /YES, STORE M OK M
0764 1105 TAD PS12
0765 3055 DCA STORE
0766 1055 TAD STORE
0767 7104 CLL RAL /2STORE
0770 4543 JMS I FLOAT
0771 4407 FENTER
0772 4100 FDIU P1
0773 6123 1101 F /T 285*ORE/1
0774 0000 FEXT
0775 5747 JMF I 100
1000 7746 M/FRO, 0032 /-RM O
1001 7652 MFNR, 3 0160 /-RM 1
1002 0002 MLC, 160 0156 /RM, LOWER CASE
/CONTROL OF DATA, (MODEL DATA) END, LOW TAFE *)
1003 0000 CONTR, 0
1004 4406 READ
1005 1700 TAD MZERO /ZERO ?
1006 7450 SNA /YES, EXT
1007 7603 JMF I CONTR /END OF DATA ?
1010 1701 TAD MND
1011 7450 SNA /YES, START (ALI.
1012 5225 JMF CALCU /LOWER CASE )
1013 1702 TAD MLC /NO ERROR
1014 7440 SZA CLA /YES, SKIP
1015 4535 JMS I BUG /FILL NEXT ZERO
1016 4406 READ
1017 1700 TAD MZERO
1018 7450 SNA CLA
1021 5603 JMF I CONTR
1027 5216 JMF
/START OF CALCULATIONS.
1028 4406 CALCU, REAR
1029 7300 CLA CLL
1030 1074 TAD MAXLOC /CALC. LOR. OF CENTRE
1031 1075 TAD MINLOC
1032 7110 CLL RAL /MAXLOC=
1033 3074 DCA MAXLOC /MINLIN+MAXLOC)/2
1034 4404 WRITE
1035 1545 MSG11
1036 1106 TAD J /NO. OF DATA POINTS*
1037 4543 JMS I FLOAT
1038 4407 FENTER
1039 6115 FFUT C /C-1/DAT(J)
1040 0040 OUTPUT /FAINT
1041 5126 FGEI V
1042 4115 FDIU C /AVERAGE=
1043 6126 FFUT V /V U/1-LOAT(J)
1044 0000 FEXT
1045 1074 TAD MI
1046 1022 TAD KUN
1047 7440 SZA CLA
1050 5262 JMF CALL I
1051 1026 TAD MSG11A
1052 1027 TAD MSG11B
1053 7440 SZA CLA
1054 5262 JMF CALL I /NO
1055 4407 FENTER /YES, CALCULATE
1056 7047 FGET RES /RESOLUTION
1057 4115 FDIU C /RES=10000/5451E5#J
1060 6047 FFUT RES
1061 0000 FEXT
1062 3107 DCA R /INITIALIZE R
1063 1107 TAD R
1064 4503 DIA ANL
1065 1112 DIA ANL
1066 4407 FENTER
1067 5512 FGET I ANL /SURFACT AVERAGE
1070 2126 V FMFUT V /A(A)-A(A) V
1071 6512 FFUT I ANL
1072 0000 FEXT
1073 2107 ISZ R
1074 1107 TAD R
1075 7041 (IA
1076 1106 TAD J
1077 7440 SZA CLA /A J ?
1080 5263 JMF CALC /NO-CONTINUE
1081 1106 TAD J /YES
1082 4503 JMS I INDEXL
1083 7055 DCA STORE
1104 6711 (DF 0
1105 3405 DIA I 1000 /L AN 100
1106 2025 ISZ 5100 /J 1 7777
1107 5305 JMP -2
1110 6301 100 00
1111 4736 JMS I AF DI /PREFAKE FUK
1112 7200 CLA /AFODISATION
/K A-RANGE DATA 31 F RAT /CENTRE IS IN LOCATION 0
1113 05 TAD P512
1114 7041 CIA
1115 074 DIA
1116 7450 SNA MAXLOC 5 2
1117 5364 JMS I INDEXL /YES, JUNE
1118 3114 DLA M /H MAXLIN *12
1119 1007 DCA M
112 114 TAD 4
1123 7210 SZA -4 /H 0 ?
1124 5337 JMS I INDEXL /NO
1125 1107 TAD N /H 5-SHIP LEFT H PLACES
1126 4503 JMS I INDEXL /A 5-A(A)
1127 5115 DCA ANL /A 5-A(A)
1128 7210 SZA -4 /H 0 ?
1129 5337 JMS I INDEXL /NO
1130 1107 TAD N /H 5-SHIP LEFT H PLACES
1131 4503 JMS I INDEXL /A 5-A(A)
1132 7210 SZA -4 /H 0 ?
1133 5337 JMS I INDEXL /NO
1134 1107 TAD N /H 5-SHIP LEFT H PLACES
1135 4503 JMS I INDEXL /A 5-A(A)
1136 7210 SZA -4 /H 0 ?
1137 5337 JMS I INDEXL /NO
1138 1107 TAD N /H 5-SHIP LEFT H PLACES
1139 4503 JMS I INDEXL /A 5-A(A)
1140 7210 SZA -4 /H 0 ?
1141 5337 JMS I INDEXL /NO
1142 1107 TAD N /H 5-SHIP LEFT H PLACES
1143 4503 JMS I INDEXL /A 5-A(A)
1144 7210 SZA -4 /H 0 ?
1145 5337 JMS I INDEXL /NO
1146 1107 TAD N /H 5-SHIP LEFT H PLACES
1147 7001 IAC
1148 7041 CIA
1149 4503 JMS I INDEXL /A 5-A(A)
1150 5112 DCA ANL /A 5-A(A)
1151 4407 KRG2, FENTER
1152 7210 SZA -4 /H 0 ?
1153 5337 JMS I INDEXL /NO
1154 1107 TAD N /H 5-SHIP LEFT H PLACES
1155 4503 JMS I INDEXL /A 5-A(A)
1156 7210 SZA -4 /H 0 ?
1157 5337 JMS I INDEXL /NO
1158 1107 TAD N /H 5-SHIP LEFT H PLACES
1159 4503 JMS I INDEXL /A 5-A(A)
1160 7210 SZA -4 /H 0 ?
1161 5337 JMS I INDEXL /NO
1162 1107 TAD N /H 5-SHIP LEFT H PLACES
1163 4503 JMS I INDEXL /A 5-A(A)
1164 7210 SZA -4 /H 0 ?
1165 5337 JMS I INDEXL /NO
1166 1107 TAD N /H 5-SHIP LEFT H PLACES
1167 4503 JMS I INDEXL /A 5-A(A)
1168 7210 SZA -4 /H 0 ?
1169 5337 JMS I INDEXL /NO
1170 1107 TAD N /H 5-SHIP LEFT H PLACES
1171 4503 JMS I INDEXL /A 5-A(A)
1172 7210 SZA -4 /H 0 ?
1173 5337 JMS I INDEXL /NO
1174 1107 TAD N /H 5-SHIP LEFT H PLACES
1175 4503 JMS I INDEXL /A 5-A(A)
1176 7210 SZA -4 /H 0 ?
1177 5337 JMS I INDEXL /NO
1178 1107 TAD N /H 5-SHIP LEFT H PLACES
1179 4503 JMS I INDEXL /A 5-A(A)
1180 7210 SZA -4 /H 0 ?
1181 5337 JMS I INDEXL /NO
1182 1107 TAD N /H 5-SHIP LEFT H PLACES
1183 4503 JMS I INDEXL /A 5-A(A)
1184 7210 SZA -4 /H 0 ?
1185 5337 JMS I INDEXL /NO
1186 1107 TAD N /H 5-SHIP LEFT H PLACES
1187 4503 JMS I INDEXL /A 5-A(A)
1188 7210 SZA -4 /H 0 ?
1189 5337 JMS I INDEXL /NO
1190 1107 TAD N /H 5-SHIP LEFT H PLACES
1191 4503 JMS I INDEXL /A 5-A(A)
1192 7210 SZA -4 /H 0 ?
1193 5337 JMS I INDEXL /NO
1194 1107 TAD N /H 5-SHIP LEFT H PLACES
1195 4503 JMS I INDEXL /A 5-A(A)
1196 7210 SZA -4 /H 0 ?
1197 5337 JMS I INDEXL /NO
1198 1107 TAD N /H 5-SHIP LEFT H PLACES
1199 4503 JMS I INDEXL /A 5-A(A)
1200 7210 SZA -4 /H 0 ?
1201 5337 JMS I INDEXL /NO
1202 1107 TAD N /H 5-SHIP LEFT H PLACES
1203 4503 JMS I INDEXL /A 5-A(A)
1204 7210 SZA -4 /H 0 ?
1205 5337 JMS I INDEXL /NO
1206 1107 TAD N /H 5-SHIP LEFT H PLACES
1207 4503 JMS I INDEXL /A 5-A(A)
1208 7210 SZA -4 /H 0 ?
1209 5337 JMS I INDEXL /NO
1210 1107 TAD N /H 5-SHIP LEFT H PLACES
1211 4503 JMS I INDEXL /A 5-A(A)
1212 7210 SZA -4 /H 0 ?
1213 5337 JMS I INDEXL /NO
1214 1107 TAD N /H 5-SHIP LEFT H PLACES
1215 4503 JMS I INDEXL /A 5-A(A)
1216 7210 SZA -4 /H 0 ?
1217 5337 JMS I INDEXL /NO
1218 1107 TAD N /H 5-SHIP LEFT H PLACES
1219 4503 JMS I INDEXL /A 5-A(A)
1220 7210 SZA -4 /H 0 ?
1221 5337 JMS I INDEXL /NO
1222 1107 TAD N /H 5-SHIP LEFT H PLACES
1223 4503 JMS I INDEXL /A 5-A(A)
1224 7210 SZA -4 /H 0 ?
1225 5337 JMS I INDEXL /NO
1226 1107 TAD N /H 5-SHIP LEFT H PLACES
1227 4503 JMS I INDEXL /A 5-A(A)
1228 7210 SZA -4 /H 0 ?
1229 5337 JMS I INDEXL /NO
1230 1107 TAD N /H 5-SHIP LEFT H PLACES
1231 4503 JMS I INDEXL /A 5-A(A)
1232 7210 SZA -4 /H 0 ?
1233 5337 JMS I INDEXL /NO
1234 1107 TAD N /H 5-SHIP LEFT H PLACES
1235 4503 JMS I INDEXL /A 5-A(A)
1236 7210 SZA -4 /H 0 ?
1237 5337 JMS I INDEXL /NO
1238 1107 TAD N /H 5-SHIP LEFT H PLACES
1239 4503 JMS I INDEXL /A 5-A(A)
1240 7210 SZA -4 /H 0 ?
1241 5337 JMS I INDEXL /NO
1242 1107 TAD N /H 5-SHIP LEFT H PLACES
1243 4503 JMS I INDEXL /A 5-A(A)
1244 7210 SZA -4 /H 0 ?
1245 5337 JMS I INDEXL /NO
1246 1107 TAD N /H 5-SHIP LEFT H PLACES
1247 4503 JMS I INDEXL /A 5-A(A)
1248 7210 SZA -4 /H 0 ?
1249 5337 JMS I INDEXL /NO
1250 1107 TAD N /H 5-SHIP LEFT H PLACES
1251 4503 JMS I INDEXL /A 5-A(A)
1252 7210 SZA -4 /H 0 ?
1253 5337 JMS I INDEXL /NO
1254 1107 TAD N /H 5-SHIP LEFT H PLACES
1255 4503 JMS I INDEXL /A 5-A(A)
1256 7210 SZA -4 /H 0 ?
1257 5337 JMS I INDEXL /NO
1258 1107 TAD N /H 5-SHIP LEFT H PLACES
1259 4503 JMS I INDEXL /A 5-A(A)
1260 7210 SZA -4 /H 0 ?
1261 5337 JMS I INDEXL /NO

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1262	412)	DCA S	/A=0 OR 512, SET S 0	1443	4407	FENTER	
1263	5121	DCA 511		1444	5047	FGET AFS	
1264	5122	DCA 512		1445	1047	FADD RES	/CALCULATE
1265	4407	FENTER		1446	1047	FAIR RES	/RESOLUTION
1266	5125	FGET I		1447	6047	FFUT RES	/FREQ-3KELS
1267	0001	SOME		1450	6030	FPUT FLOW	/FLOW-RES
1270	6131	FPUT F		1451	0000	FEXT	
1271	5130	FGET S		1452	1105	TAD PS12	
1272	0001	SOME		1453	4543	JMS I FLOAT	
1273	1133	FADD T		1454	4407	FENTER	
1274	0007	SOKK I	/MODULUS	1455	3041	FMFY DELF	
1275	6512	FPUT I ANI	/A(K)=50000(I-1453)	1456	6033	FPUT FHIGH	/FHIGH 512*INFL
1276	0000	FEXT		1457	0000	FEXT	
1277	4477	JMS I LAST	/A=512	1460	7344	STA CLL RAL	
1278	5236	JMF FMSR	/NO, CONTINUE	1461	1024	TAD NI	
1301	1037	TAD ISFLTA	/FIRST SPECTRUM ?	1462	7640	SZA CLA	/NI=2 ?
1302	1036	TAD NSFCTR		1463	5320	JMF DIRECT	
1303	7640	SZA CLA		1464	4404	WRITE	/YES
1404	5333	JMF NEXT	/YES, WRITE ON L TAPE	1465	3333	MESG23	/PLOT
1405	1107	DCA I	/NO, MOVE DATA	1466	4404	WRITE	/SAMPLE AND
1306	4476	MOVE		1467	3504	MESG7	/BACKGROUND
1307	4407	FENTER		1470	4537	JMS I ANSWER	
1310	5132	FGET I ANI		1471	1140	TAD ANSWR	
1311	6510	FPUT I AR	/A(HFSTOKE) A(K)	1472	1060	TAD MINYES	
1312	0000	FEXT		1473	7640	SZA CLA	
1313	4477	JMS I LAST	/A=512 ?	1474	5312	JMF DIRECT	/NO-PLOT RATIO
1314	5306	JMF MOVE	/NO, CONTINUE	1475	1056	TAD IFIRST	/YES
1315	1061	TAD MINB	/YES, READ PREVIOUS	1476	3533	DCA I BLATP	
1316	4732	JMS I KDTAPE	/SUM FROM L TAPE	1477	1061	TAD MINB	/READ SAMPLE
1317	1061	TAD MINB		1500	4532	JMS I KDTAPE	/FROM L-TAPE
1318	1533	TAD I BLATP		1501	4472	JMS I NORMI	/NORMALIZE
1319	3333	DCA I BLATP	/IF BLK FBK-L-B	1502	4466	JMS I SFFLT	/FLOT SCALES
1320	3107	DCA I		1503	4467	JMS I SFFLT	/FLOT SAMPLE
1321	4476	JMF I INIIL		1504	1061	TAD MINB	/IF BLK BACKGROUND
1322	5131	FGET I ANI		1505	4531	JMS I KDTAPE	/NORMALIZE
1323	1510	FADD I AR		1506	4472	JMS I NORMI	/NORMALIZE
1324	6511	FPUT I ANI	/A(K) A(K) SUM(K)	1507	4467	JMS I SFFLT	/FLOT
1325	0000	FEXT		1510	1061	TAD MINB	/READ RATIO SPECTRUM
1326	4477	JMS I LAST	/A=512 ?	1511	4531	JMS I KDTAPE	/FROM L-TAPE
1327	5323	JMF AND	/NO, CONTINUE	1512	4471	JMS I MAXIM	/FIND MAXIMUM
1328	201	CLA LAC		1513	5321	JMF DIRECT+1	
1329	1036	TAD NSFLTA	/NSFLTA 1 ?	1514	0007	HNRD0	/100
1330	7640	SZA CLA		1515	3100	3100	
1331	5343	JMF AND	/NO	1516	0000	0000	
1332	7344	STA CLL RAL	/YES	1517	1401	FPLOT52	
1333	1036	TAD NI	/NI=2 ?	1518	4472	DIRL1	/IF NI=1, NORMALIZE
1334	7640	SZA CLA		1521	4404	WRITE	
1335	1061	JMF I RATIO	/NO, CALCULATE FREQ.	1522	3730	MESG28	/FLOW
1336	4531	JMS I KDTAPE	/YES, WRITE	1523	4407	FENTER	
1337	7344	STA CLL RAL	/ON L-TAPE	1524	5030	FGET FLOW	
1338	1036	TAD NI		1525	0012	OUTPUT	
1339	7640	SZA CLA		1526	0040	40	/PRINT FLOW
1340	1036	TAD NI	/NR ? ?	1527	0000	FEXT	
1341	7640	SZA CLA		1530	4404	WRITE	
1342	1036	TAD NI	/NO	1531	1736	MESG29	/FHIGH
1343	7640	SZA CLA	/YES, READ AND	1532	4407	FENTER	
1344	1036	TAD NI	/TRANSFORM BACKGR	1533	5033	FPGET FHIGH	
1345	7640	SZA CLA		1534	0017	OUTPUT	
1346	1036	TAD NI	/RUN ?	1535	0040	40	/PRINT FHIGH
1347	7640	SZA CLA	/LAST SPECTRUM ?	1536	0000	FEXT	
1348	1036	TAD NI	/NO, READ NEXT SAMPLE	1537	4404	WRITE	
1349	7640	SZA CLA	/CALCULATE RATIO	1540	3652	MESG18	/MAXIMUM(IN X):*
1350	1036	TAD NI		1541	4407	FENTER	
1351	7640	SZA CLA		1542	5036	FGET MAX	
1352	1036	TAD NI	/RUN NI ?	1543	3314	FMFY HNRD0	
1353	7640	SZA CLA		1544	0012	OUTPUT	
1354	1036	TAD NI		1545	0030	30	/PRINT MAX
1355	7640	SZA CLA		1546	0000	FEXT	
1356	1036	TAD NI	/YES, LAST SAMPLE ?	1547	4546	JMS I PDRIF	
1357	7640	SZA CLA	/NO, CONTINUE	1548	7081	CLA TAC	
1358	1036	TAD NI	/YES	1551	3022	DCA RUN	
1359	7640	SZA CLA	/SFCTR-BNK	1552	4404	WRITE	/NEW FREQ LIMITS ?
1360	1036	TAD NI		1553	3715	MESG27	
1361	7640	SZA CLA	/NSFCTR -PSFCTR	1554	4533	JMS I ANSWER	
1362	1036	TAD NI	/READ FIRST BANK	1555	1140	TAD ANSWR	
1363	7640	SZA CLA		1556	1060	TAD MINYES	
1364	1036	TAD NI		1557	7640	SZA CLA	
1365	7640	SZA CLA		1560	5717	JMF I PLOT+1	/NO
1366	1036	TAD NI	/IF BLK=FBK-L-B	1561	4404	WRITE	/YES
1367	7640	SZA CLA	/LAST BACKGR ?	1562	3732	MESG28	/FLOW
1368	1036	TAD NI	/NO, CONTINUE	1563	4541	JMS I SICNV	/INPUT
1369	7640	SZA CLA		1564	4543	JMS I FLOAT	
1370	1036	TAD NI		1565	4407	FENTER	
1371	7640	SZA CLA		1566	6030	FPUT FLOW	
1372	1036	TAD NI		1567	0000	FEXT	
1373	7640	SZA CLA		1570	4404	WRITE	
1374	1036	TAD NI	/READ SAMPLE AND BACKGR	1571	3736	MESG29	/FHIGH
1375	7640	SZA CLA		1572	4541	JMS I SICNV	/INPUT
1376	1036	TAD NI		1573	4543	FENTER	
1377	7640	SZA CLA	/IF NI=1, CALCULATE	1574	4407	FPUT FHIGH	
1378	1036	TAD NI	/FREQUENCY	1575	0000	FEXT	
1379	7640	SZA CLA		1576	0000	FEXT	
1380	1036	TAD NI		1577	4546	JMS I PDRIF	
1381	7640	SZA CLA		1600	2022	ISZ RUN	
1382	1036	TAD NI		1601	4404	WRITE	/NEW MAXIMUM **
1383	7640	SZA CLA		1602	3641	MESG17	
1384	1036	TAD NI	/CALCULATE RATIO	1603	4537	JMS I ANSWER	
1385	7640	SZA CLA	/A(K)/A(HFSTOKE)	1604	1140	TAD ANSWR	
1386	1036	TAD NI		1605	1060	TAD MINYES	
1387	7640	SZA CLA	/CALCULATE FREQUENCY	1606	7640	SZA CLA	
1388	1036	TAD NI		1607	5222	JMF SH	/NO
1389	7640	SZA CLA		1610	4404	WRITE	/YES
1390	1036	TAD NI	/A(K)+FSTORE=DELF*FLOAT(K)	1611	3652	MESG18	/MAXIMUM(IN X):*
1391	7640	SZA CLA		1612	4541	JMS I SICNV	/INPUT
1392	1036	TAD NI	/A=512 ?	1613	4543	JMS I FLOAT	
1393	7640	SZA CLA	/NO, CONTINUE	1614	4407	FENTER	
1394	1036	TAD NI		1615	3243	FMFY HNRD0	
1395	7640	SZA CLA	/IF NI=2, WRITE RATIO	1616	2036	FPUT MAX	
1396	1036	TAD NI	/ON L-TAPE	1617	0000	FEXT	
1397	7640	SZA CLA		1620	4546	JMS I FCRIF	
1398	1036	TAD NI		1621	2031	ISZ RUN	
1399	7640	SZA CLA		1622	4404	WRITE	/SMOOTHE **
1400	1036	TAD NI		1623	1743	MESG30	



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1674 4517 JMS I ANSWER
1625 1140 TAD ANWK
1626 1060 TAD MINYS
1677 7640 SFA CLA
1630 5231 JMS I J3 /NO START FLOT
1631 4473 JMS I SMTH1 /YES SMOOTH
1632 2022 ISZ RUN
1633 1022 TAD RUN
1634 7650 SMA LI A
1635 5646 JMS I HGN
1636 4466 JMS I SCFLT /FLOT SCALES
1637 4465 JMS I RESNR /FLOT NR AND RES
1640 4467 JMS I SHFLT /FLOT SFLTRUM
1641 5642 JMF I J1
1642 5551 FLOPS
1643 7777 HNDK 7772 /O.OI
1644 7436 2436
1645 5607 5607
1646 5361 HGNW HGNW
1647 5551 FLOPS
1648 5361 HGNW HGNW
1649 5551 FLOPS
1650 0000 FFI 0
1651 4317 JMS FFIH1 /PERMUTE TO COMPLEX
1652 4662 JMS I HATH /DO FFT
1657 4661 JMS I UNSK1 /UNSCRAMBLE
1660 5654 JMF I FFI /EXIT
1661 7000 UNSK1 7000
1662 7000 HATH 7000
1663 7000 /SUBROUTINE TO FIND FLOATING ADDRESSES OF
/((C(A(512+J)+A(K)), AND A(512+K))
1664 0000 SIHOF 0
1664 1106 TAD J
1665 4303 JMS INDEX
1666 5110 DCA A1 /ADDRESS OF A(J)
1667 1106 TAD J512
1670 1106 TAD J
1671 4303 JMS INDEX
1672 5111 DCA ANP1 /ADDRESS OF A(512+J)
1673 1106 TAD K
1674 4303 JMS INDEX
1675 5111 DCA AK /ADDRESS OF A(K)
1676 1106 TAD J512
1677 1107 TAD K
1700 4303 JMS INDEX
1701 5111 DCA ANP1 /ADDRESS OF A(512+J)
1702 5665 JMF I SETOF /EXIT
1703 0000 /SUBROUTINE TO FIND FLOATING ADDRESS OF A(C(A(512+J)
INDEX 0
1704 5111 DCA INXS1R
1705 5111 TAD INXS1R
1706 7004 LIT KAI
1707 5111 TAD INXS1R
1710 5703 JMS I INMX1
1711 0000 INXS1R 0
1712 0000 /SUBROUTINE TO HIT REVERSE PERMUT 1024
/EDDATING DATA
FITH1 0
1713 0000 FITH1 0
1714 7151 CIA SFL IAC KTR
1715 7001 TAD
1716 5172 DCA (TR1 /1072 DATA
1717 7001 CIA IAI /NO.01 DATA
1718 5106 DCA I
1721 1106 FTHNXT TAD J
1722 5170 DCA FLF1
1723 5177 TAD F10
1724 7041 CIA
1725 5175 DCA (TR2 /10 BITS TO BE REV.
1726 5171 FIFLOF DCA FLF1
1727 5170 TAD FIP1
1730 7110 CIL KAR /READ J FROM K TO I
1731 5170 DCA FLF1
1732 5171 TAD FLK1
1733 7004 KAL /MULTIPLY FROM L TO R
1734 7171 LIT (TR1 /HAVE ALL BITS BEEN REV.
1735 5176 JMF FLFLOF /NO-NEXT
1736 7107 DCA K /K BIT REVERSED J
1737 1107 TAD K
1740 7041 CIA
1741 1106 TAD I
1742 7710 CIA CLA /IS K J ?
1743 4375 JMS FLTHOV /YES SWITCH DATA J,K
1744 7106 LIT J /NEXT J
1745 5172 ISZ CTR1 /HAVE ALL DATA BEEN DONE
1746 5321 JMF FTHNXT /NO-NEXT
1747 5712 JMF I FLTHNT /EXIT
1750 0000 FLFJ 0
1751 0000 FLFN 0
1752 0000 /SUBROUTINE TO SWITCH DATA A(J) AND A(K)
FLTHOV 0
1753 1106 TAD J
1754 4503 JMS I INMX1
1755 7001 DCA AJL /ADDRESS OF A(J)
1756 1107 TAD K
1757 4503 JMS I INMX1L /ADDRESS OF A(K)
1760 3117 DCA AKL
1761 4407 FENTER
1762 5512 FGFI T AI
1763 6374 FFUT T3 /T3-A(K)
1764 5510 FGFI T AJL /A(K)-A(J)
1765 6512 FFUT T AI
1766 5174 FLOF T5 /A(I) T3
1767 6510 FFUT T AI
1770 0000 FFI
1771 5717 JMF I FLTHOV /EXIT
1772 0000 CTR1 0
1773 0000 CTR2 0
1774 0000 I3 0
1775 0000 0
1776 0000 0
1777 0012 F10 12
20000
/SUBROUTINE TO TRANSFORM 512 FLOATING-POINT
2000 0000 /COMPLEX DATA
MAIN 0
2001 7201 CLA IAC
2002 3366 DCA DJ
2003 1275 TAD PY /START WITH DJ=1
2004 7041 CIA
2005 3367 DCA MINF /9 INTERMEDIATE TRANS.
2006 4304 JMS LOOP1 /SET UP FOR INT. TRANS.
2007 3370 DCA L /L=0
2010 1114 TAD H
2011 7041 CIA
2012 3371 DCA MINL /H PART.-INT.
2013 1370 TAD L
2014 3106 DCA J /J=L
2015 1104 TAD J
2016 1114 TAD H
2017 3107 DCA K /K=J+H
2018 4674 JMS I SETOF /GET ADDRESSES
2021 4407 FENTER
2022 5513 FGFI T ANP1
2023 3120 FFY S
2024 6276 FFUT D1
2025 5512 FGFI T AKL
2026 5513 FFY C
2027 1276 FADD D1
2030 6276 FFUT D1 /D1=C*(A(K)+S*(A(512+K))
2031 5512 FGFI T AKL
2032 3120 FFY S
2033 6301 FFUT E1
2034 5513 FGFI T ANP1
2035 3115 FFY C
2036 2301 FSUB E1
2037 6301 FFUT E1 /E1=C*(A(512+K)-S*(A(K))
2040 5510 FGFI T AJL
2041 2276 FSUB D1
2042 6512 FFUT I AKL /A(K)=A(J)-D1
2043 5511 FGFI T ANP1
2044 2301 FSUB E1
2045 4511 FFUT I ANP1 /A(512+K)=A(512+J)-E1
2046 5510 FGFI T AJL
2047 1276 FADD D1
2050 6510 FFUT I AJL /A(J)=A(J)+D1
2051 5511 FGFI T ANP1
2052 1301 FADD E1
2053 6511 FFUT I ANP1 /A(512+J)=A(512+J)+F1
2054 0000 FFI
2055 1106 TAD J
2056 1366 TAD DJ /J=J+DJ
2057 3106 DCA J
2060 1106 TAD J
2061 7041 CIA
2062 1104 TAD PS11
2063 7200 CIA CLA
2064 5215 JMF JLOOP
2065 2371 ISZ MINL /NO-NEXT PASS OF PART.
2066 5272 JMF HUREL /YES-ALL INT. S DONE
2067 2367 ISZ MINF /YES-ALL INT. S DONE
2070 5206 JMF I HUREL /NO-NEXT INTERMEDIATE
2071 5609 JMF I HATH /YES-EXIT
2072 4340 MOREL JMS SETOF /SET UP FOR NEXT PARTIAL
2073 5213 JMF LLOOP
2074 1663 SETOF SETLOC
2075 5171 FFI
2076 0000 D1 0
2077 0000 0
2080 0000 0
2081 0000 0
2082 0000 0
2083 0000 0
2084 0000 0
2085 0000 0
2086 0000 0
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2100 0000 0
2101 0000 0
2102 0000 0
2103 0000 0
2104 0000 0
2105 1366 TAD DJ
2106 3114 DCA H /H=DJ
2107 1366 TAD DJ
2110 7171 CIL KAL
2111 5366 DCA DJ /DJ=2*DJ
2112 1114 TAD H
2115 4543 JMS I FLOAT
2116 4407 FENTER
2117 6372 FFUT R4
2118 5100 FGFI PI
2119 4375 FFUT R4 /R4=PI/H
2120 6372 FFUT R4
2121 0004 COS
2122 6123 FFUT T /T=COS(R4)
2123 5372 FGFI T4
2124 0003 SIN
2125 6126 FFUT V /V=SIN(R4)
2126 0000 FFI
2127 7201 CLA IAC
2128 3115 DCA C
2129 7132 SIL R1R
2130 3116 DCA C1
2131 3117 DCA C2 /C=1
2132 3120 DCA S
2133 3121 DCA S1
2134 5172 DCA S2
2135 5704 JMF I LOOP1 /EXIT
2137 5704 /SUBROUTINE TO SET UP FOR INTERMEDIATE PARTIAL
/TRANSFORM EXCEPT FOR FIRST ONE
SETUP1 0
2140 0000 FENTER
2142 5115 FGFI T5
2143 6377 FFUT R4 /R4=C
2144 5170 FGFI S
2145 3126 FFY V
2146 6115 FFUT L
2147 5372 FGFI R4
2148 5123 FFY T
2149 2115 FSUB C
2150 6115 FFUT C /C=R4*T-S*V
2151 5170 FGFI S
2152 3123 FFY T

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2115 6120 FPUT S
2136 5121 FGET R4
2157 4116 FMEY V
2160 1120 FADD S
2161 6110 FPUT S
2162 0000 FEXT
2163 2370 ISZ L
2164 7000 NOP
2165 7240 JMP I SEUFL /EXIT
2166 0000 DJN 0
2167 0000 MINP, 0
2170 0000 L, 0
2171 0000 MINL, 0
2172 0000 R4, 0
2173 0000
2174 0000
2200 0000 /SUBROUTINE TO UN-SCRAMBLE COEFFICIENTS
UNSCRM, 0
2201 4730 JMS SETUF /DO A(0)+A(512)
2202 1105 TAB F512 /SET UP REST
2203 7110 CII RAR
2204 7041 CIA
2205 7001 IAL
2206 3114 ICA N /255 ITERATIONS
2207 7201 CIA IAC
2210 3106 DLA J /START WITH J 1
2211 4731 JMS LOOF4 /UN SCRAMBLE
2212 2106 NEXT J /NEXT J
2213 3114 ISZ H /IS SCRAMBLING DONE
2214 5211 JMF UNLOOF /NO NEXT
2215 2361 TS7 FN /YES DO A(256)+A(768)
2216 4775 JMS I SEUFL
2217 4407 FENTER
2220 5010 FNET I AJL
2221 3361 FMEY FN
2222 6510 FPUT I AJL
2223 5511 FGET I ANF JL
2224 3361 FMEY FN
2225 6511 FPUT I ANF JL
2226 0000 FEXT
2227 5600 JMP I UNSCRM /EXIT
2230 0000 /SUBROUTINE TO UN-SCRAMBLE A(0)+A(512)+ AND
SEUFL, 0
2231 1105 TAB F512
2232 4503 JMS I INDEX
2233 3112 ICA ANL /ADDRESS OF A(512)
2234 3020 DOA STORE
2235 1105 TAB F512
2236 4543 JMS I FLOAT
2237 4407 FENILK
2240 6161 FPUT I N INJ
2241 5462 FGET I STORE
2242 6173 FPUT I /T A(0)
2243 1012 FADD I ANL
2244 4361 F2IV FN
2245 6455 FPUT I STORE /A(0)=(A(0)+A(512))/2
2246 5123 FGET I
2247 2512 FSL I ANL
2250 4361 F2IV FN
2251 6512 FPUT I ANL /A(512)=(T-A 512))/512
2252 5100 FGET I I
2253 4361 F2IV FN
2254 6777 FPUT F5 /F5=F I/512
2255 0004 COS
2256 6173 FPUT I /T-COS(F5)
2257 6115 FPUT C /C=T
2260 5172 FGET F5
2261 0003 SIN
2262 6126 FPUT V /V=SIN(F5)
2263 6120 FPUT S /S=V
2264 5351 FGET ZFS
2265 4361 F2IV FN
2266 6361 FPUT FN /FN=1/1024
2267 0000 FEXT
2270 7630 JMF I SEUFL /EXIT
2271 0000 /SUBROUTINE TO DO ACTUAL UN-SCRAMBLING
LOOF4, 0
2272 1106 TAB J
2273 7041 CIA
2274 1105 TAB F512
2275 3107 ICA N /K=512-J
2276 4755 JMS I SEUFL /GET ADDRESSES
2277 4407 FENTER
2280 5511 FGET I ANF JL
2281 5513 FADD I ANF JL
2282 4775 FPUT RS /R5=A(512+J)+A(512+K)
2283 3115 FMEY C
2284 6164 FPUT DS /DS=C#R5
2285 5375 FGET R5#
2286 1120 FMEY S
2287 6375 FPUT RS /R5=S#R5
2290 5510 FGET I AJL
2291 2512 FSUB I ANL
2292 6356 FPUT O
2293 3115 FMEY C
2294 1375 FADD RS
2295 6167 FPUT F5 /E5=C#R+R5
2296 5156 FGET O
2297 3110 FMEY S
2298 2364 FSUB DS
2299 6156 FPUT O /O=S#O-DS
2302 5010 FGET I AJL
2303 1512 FADD I ANL
2304 6512 FPUT I ANL /A(K) A(J)+A(K)
2305 2356 FSUB O
2306 3361 FMEY FN
2307 6510 FPUT I AJL
2308 5512 FGET I ANL
2309 1516 FMEY FN
2310 3361 FPUT I AJL
2311 5512 FGET I ANL
2312 1516 FMEY FN
2313 3361 FPUT I ANL /A(K)-(A(K)+O)#FN
2314 6112 FPUT I ANL /A(K)-(A(K)+O)#FN
2334 5112 FGET I ANF JL
2335 2513 FPUT I ANF JL
2336 1117 FPUT I ANF JL
2337 6167 FADD F5
2338 1117 FMEY I N
2339 6513 FPUT I ANF JL
2340 6513 FPUT I
2341 6513 FPUT I
2342 5367 FGET I
2343 111 FPUT I ANF JL
2344 1361 FMEY FN
2345 6511 FPUT I ANF JL
2346 0000 FEXT
2347 4774 JMS I SEUFL
2350 671 JMS I LOOF4
2351 0000
2352 2000
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210	5100	3100	2667	004	TWENTY	/20
211	0000	0000	2670	2507	FIFTY	/50
/SUBROUTINE 5 TO FLOT OF THE SPECTRUM.						
/INIT OF THE SCALERS.						
212	2455	MS10K1	2672	1643	DELTA	/0.01
213	0006	F45	2673	0022	F18	/0.02
214	2640	0006	2674	0050	F40	/40
215	0000	2640	2675	1130	P600	/600
216	5654	DELTA	2676	4770	F4770	4770
217	5270	JMF I	2677	3070	XCORD1	XCORD1
218	0000	SEFLOT	2700	1100	YCORD1	YCORD1
219	4404	SCALE2	2701	4407	SCALE2	FENIF
220	5674	WRITE	2703	5033	FLOT FHIGH	
221	4512	MS674	2705	1170	F508 S	
222	1140	MS ANSWER	2706	0000	FEXT	
223	1312	TAC ANSWR	2708	1160	TAD HOKD	
224	2640	AL HSTAKT	2709	2700	SMA CLA	
225	114	52A 1LA	2707	5224	JMF SCALE1+4	
226	4477	JMF 4	2710	3150	DCA ASLIX	
227	0006	FENTER	2711	1275	TAD F600	
228	113	DF0	2712	1171	DCA ASCLY	
229	113	FLOT FHIGH	2713	4404	WRITE	/FREQUENCY (CM )*
230	2030	FLOW FLOW	2714	3161	MS675	
231	6154	FLOW	2715	1150	TAD ASCLX	
232	2000	1275	2716	1275	TAD F18	
233	2727	52A	2717	1150	DCA ASCLX	
234	527	LOT	2720	1151	TAD ASCLY	
235	2701	LA TAC	2721	1774	TAD F40	
236	4405	FLOT	2722	3151	DCA ASCLY	
237	0170	0170	2723	7105	C IAC RAL	/2
238	0060	0060	2724	3153	DCA SIZE	
239	2700	LA	2725	4404	WRITE	/*-1*
240	4405	LOT	2726	3171	MS626	
241	0120	0120	2727	7307	CIA CLL IAC RTL	/4
242	4620	4620	2730	3123	DCA SIZE	
243	2700	LA	2731	3120	DCA S	/PLOT AMPL. SCALE
244	4403	FLOT	2732	3121	DCA S+1	
245	544	2544	2733	3122	DCA S+2	
246	4620	4620	2734	4407	FENIF	
247	2700	FLOT	2735	5036	FLOT MAX	
248	4405	2544	2736	2671	F508 I DELT	
249	0060	0060	2737	0000	FEXT	
250	2700	LA	2740	1160	TAD HOKD	
251	4403	FLOT	2741	7700	SMA CLA	/MAX 0.1 *
252	0170	0170	2742	7145	JMF +3	
253	0060	0060	2743	1777	TAD DELT+1	/YES, INTERVAL-0.01
254	0170	0170	2744	7410	SAP	
255	0060	0060	2745	1771	TAD DELT	/NO, INTERVAL-0.1
256	512	DIA TIEFLT	2746	5114	DIA H	/H INTERVAL
257	7177	LA TAC HIL	2747	4407	FENTER	
258	312	LA SIZE	2750	1120	FLOT S	
259	2701	TAC	2751	0000	FEXT	
260	5154	DCA OKIENT	2752	4677	JMS I XCORD1	
261	4407	FENTER	2753	3156	DCA +3	
262	511	FLOT F45	2754	0001	FLOT	/FLOT MARK
263	6117	FEXT	2755	4403	0000	
264	0000	FEXT	2756	0000	0000	
265	1314	TAD DELTA	2757	4670	4620	
266	30	LA STORE	2760	7200	CLA A	
267	4407	FENTER	2761	1156	TAD -3	
268	11	FLOT C	2762	5364	DCA +2	
269	16	FLOW V	2763	4403	FLOT	
270	0030	FEXT	2764	0000	0000	
271	1160	TAD HOKD	2765	4606	4606	
272	2710	LA TAC	2766	7200	LLA	
273	0060	JMF +3	2767	1774	TAD F4770	
274	1477	TAD I STOK	2770	3151	DCA ASCLY	
275	2700	JMF TACAL1	2771	1364	TAD +5	
276	4407	FENTER	2772	3150	DCA ASCLX	
277	1117	FLOT C	2773	4407	FENTER	
278	5664	F508 I DEL	2774	5120	FLOT S	
279	6117	FLOT C	2775	0017	OUTPUT	/PLOT NUMBER
280	0000	FEXT	2776	0037	37	
281	1077	LA TAC	2777	5170	FLOT S	
282	2700	JMF SCL	3000	1714	FADD I H	/ADD INTERVAL
283	5114	DCA H	3001	1170	F508 S	/S SH4
284	5170	DCA S	3002	2036	F508 MAX	
285	5170	DCA +3	3003	0000	FEXT	
286	1407	FENTER	3704	1160	TAD HOKD	
287	110	FLOT S	3005	7750	4FA SMA CLA	/S MAX ?
288	1114	FADD I H	3006	5614	JMF I SCALES	/N7-CONTINUE
289	6170	FEXT	3007	4407	FENTER	/YES
290	0000	FEXT	3010	0007	FENIF	
291	4470	JMS I IM1	3011	0000	FEXT	
292	4470	52A A	3012	5611	JMF I SCLF2	/EXIT
293	501	JMS I Y HOK1	3013	5017	SCLF17	SCLF17
294	511	DCA F4	3014	747	SCLF15	SCLF15
295	2701	TAC	3015	0000	SCLF10	SCLF10
296	4103	FLOT	3016	7700	LA	
297	0170	0170	3017	5177	LA H	
298	0000	0000	3018	3114	DCA H	
299	2700	LA	3019	4476	JMS I INITL	
300	1741	LA +2	3020	4407	FENTER	/COORDINATE
301	1741	LA +2	3021	5510	FLOT I AJL	
302	1741	DCA +3	3024	0000	FEXT	
303	4103	FLOT	3025	4470	JMS I IM1	/BETWEEN FREQ.LIMITS ?
304	0170	0170	3026	7640	52A A	
305	0000	0000	3027	5577	JMF I Y HOK2	/NO, SKIP
306	0000	0000	3030	4100	JMS YCORD1	
307	1774	LA F40	3031	3255	DCA Y HOKD	
308	1774	LA F40	3032	4407	FENTER	
309	1774	LA F40	3033	5036	FLOT MAX	
310	1774	LA F40	3034	5177	F508 I ANL	
311	1774	LA F40	3035	0000	FEXT	
312	1774	LA F40	3036	1160	TAD HOKD	
313	1774	LA F40	3037	7700	SMA CLA	/DATA MAX ?
314	1774	LA F40	3040	5114	JMF F3	
315	1774	LA F40	3041	3114	DCA H	
316	1774	LA F40	3042	5757	JMF YCORD+2	/YES, SNIF
317	1774	LA F40	3043	4407	FENTER	/X COORDINATE
318	1774	LA F40	3044	5517	FLOT I ANL	
319	1774	LA F40	3045	0000	FEXT	
320	1774	LA F40	3046	4770	JMS XCORD	

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304/ 3254 DCA XL0KD
305/ 1114 IAD H
4051 7650 SNA CLA
3052 /001 IAC
3053 4403 FLOT
3054 0000 XLOKD, 0
3055 0000 YCOKD, 0
3056 2114 ISZ H
3057 4477 JMS I LAST /N=512 /
3060 5221 JMP PLTSP /NO; CONTINUE
3061 3157 DCA TYFFLT /SET TYFF-FLOT FLAG
3062 /001 IAI /TO TYPE
3063 4403 FLOT /NEW PAGE
3064 4160 F80+ 4160
3065 0000
3066 /300 CLA CLL
3067 5615 JMF I 5FFLOT
3070 0000 XCOORD, 0
3071 4407 FENTER /X COORDINATE.
3072 4036 FITV MAX
3073 3320 FMY L1300
3074 0000 FEXT
3075 4373 JMS INT
3076 1314 TAD F80
3077 5670 JMF I XCOORD
3100 0000 YCOORD, 0 /CALCULATE
3101 4407 FENTER /Y COORDINATE.
3102 2030 FSUR FLOW
3103 4174 FITV U
3104 3315 FMY C2400
3105 2315 FSUR C2400
3106 0010 NEGATE
3107 0000 FEXT
3110 4323 JMS INT
3111 1313 TAD F48
3112 5700 JMF I YCOORD
3113 0060 P48, 0060
3114 0120 F80+ 4170
3115 0014 C2400, 0014 /2400
3116 2260
3117 0000
3120 0013 C1300, 0013 /1300
3121 2434
3122 0000
/CONVERT C-1 PNT. TO UNSIGNED 5-MILE-FREL 5.0N
/INTEGER.
3123 0000 INT, 0
3124 1160 TAD HORD
3125 7710 SNA CLA / 0
3126 5723 JMP I INT /YES+ -0
3127 1161 TAD HORD /NO
3130 7104 CLL KAL
3131 7200 CLA
3132 1160 TAD HORD
3133 7004 RAL
3134 1160 DCA HORD
3135 1157 TAD EX
3136 1353 TAD M12 / 4095 ?
3137 7510 SPA
3140 5345 JMF 1,5 /YES
3141 7640 SZA CLA /NO; 4095 ?
3142 5723 JMP I INT /YES+ 0
3143 1160 TAD HORD /NO; 4095
3144 5723 JMF I INT
3145 5117 DCA EXP
3146 1160 TAD HORD
3147 7110 CLL KAL
3150 2117 ISZ EXP
3151 5347 JMF -2
3152 5723 JMF I INT
3153 7764 M12, 7764
/TEXT-4TRIMS, FOR FLOT-ROUTINES.
3154 4040 MFS620, 4040
3155 2205 /RF
3156 3340 /S
3157 7600 /
3160 /400 MFS621, 7400 /
3161 0622 MFS625, 0622 /FR
3162 0521 0521 /EQ
3163 7505 2505 /UE
3164 1603 1603 /M
3165 3340 3340 /L
3166 7003 7003 /L
3167 1540 1540 /M
3170 5100 5100 /
3171 5764 MFS626, 5764 / 1
3172 0000
/SUBROUTINE TO SMOOTH THE SEFTALK.
/PRINT SMOOTHING WITH POLYNOMIAL OF DEGREE 3.
SMOOTH, J
3200 0000
3201 7200
3202 1107 DCA A
3203 4114 DCA H
3204 2107 I-Z H
3205 1107 TAD K
3206 4503 JMS I INDEXL
3207 3114 DCA ANF JL
3210 2107 I-Z H
3211 4407 FENTER
3212 5514 FGET I H
3213 6120 FEUT S /S A(0)
3214 5511 FGET I ANF JL /T-A(1)
3215 6123 FEUT T
3216 0000 FEXT
3217 1107 SMTH, TAD A
3220 4503 JMS I INDEXL
3221 3112 DCA ANL
3222 1107 TAD A
3223 7001 IAC
3224 4503 JMS I INDEXL
3225 1110 DCA A JL
3226 7324 CLA 511 KTL

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3227 1107 TAD K
3228 4503 JMS I INDEXL
3229 1113 DCA ANL JL
3232 4407 FENTER
3233 5170 FEUT S
3234 6115 FEUT C
3235 1373 IGT T /C S
3236 6120 IPUT S /S-T
3237 1112 FGET I ANL
3240 6123 FEUT T /T-A(1)
3241 2110 FSUR S
3242 5710 SUBS I AJL
3243 6126 FEUT U /U-T-S A(K+1)
3244 1126 IADU V
3245 1123 FADU T
3246 6126 FEUT U /U U+U+T
3247 1126 FARD V
3248 1115 FARD C
3251 1513 FARD I ANPAL
3252 6126 FEUT V /U U+U+T A(K+2)
3253 0000 FEXT
3254 1374 STA I11 KAL
3255 1107 TAD K
3256 7640 STA I1A
3257 1121 JMS SMTH2 /NO
3260 4407 FENTER /YES
3261 3327 FMY U1
3262 1115 FAL C
3263 6514 FEUT I H /A(0) C 014785/18U
3264 5176 FEUT V
3265 1662 FMY U1
3266 1170 FALD S
3267 6513 FEUT I ANF H /A(1) 54105/147868U
3270 0000 FEXT
3271 4407 FENTER
3272 5176 FEUT V
3273 3335 FMY U1
3274 1173 FAL T
3275 6513 FEUT I ANL /A(K) C 085/14788U
3276 0000 FEXT
3277 1107 FEUT K
3280 1104 TAD F511
3281 7041 CLA
3282 1107 TAD K
3283 7640 STA I1A
3284 2117 JMS SMTH /NO CONTINUE
3285 1107 TAD K
3286 4503 JMS I INDEXL
3287 1114 DCA H
3288 2107 I-Z H
3289 1107 TAD K
3292 4503 JMS I INDEXL
3293 5311 DCA ANF JL
3294 4407 FENTER
3295 5126 FEUT V
3296 3332 FMY U1
3297 1514 FARD I H
3298 6514 FEUT I H /A(11) A(11)+
3299 5126 FEUT U / 05/14 188U
3302 3372 FMY U1
3303 1111 FARD I ANF JL
3304 6511 FEUT I ANF JL /A(1) A(1)
3305 0000 FEXT / 014785/18U
3306 4407 JMF I SMTH2 /EXIT
3307 7772 / 014785/1
3308 4257 4257
3309 462 462
3311 7774 7774 / 7571476
3312 5320 5320
3313 4520 4520
3314 5203 5203
3315 7275 7275 / 08571429
3316 5203 5203
3317 5103 5103
/DATE BETWEEN FREQUENTLY PRINTS ?
3340 0000 LIM, 0
3341 4407 FENTER
3342 6142 FEUT I 541
3343 2050 FEUT I DM
3344 0000 FEXT
3345 1160 TAD 130D
3346 7510 SFA
3347 7400 JMS I LIM
3348 4407 FENTER
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3830 7510 SFA
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3832 4407 FENTER
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3836 1160 TAD HORD
3837 7510 SFA
3838 7400 JMS I LIM
3839 4407 FENTER
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3842 0000 FEXT
3843 1160 TAD HORD
3844 7510 SFA
3845 7400 JMS I LIM
3846 4407 FENTER
3847 5013 FEUT I 5013
3848 2162 FEUT I 2162
3849 0000 FEXT
3850 1160 TAD HORD
3851 7510 SFA
3852 7400 JMS I LIM
3853 4407 FENTER
3854 5013 FEUT I 5013
3855 2162 FEUT I 2162
3856 0000 FEXT
3857 1160 TAD HORD
3858 7510 SFA
3859 7400 JMS I LIM
3860 4407 FENTER
3861 5013 FEUT I 5013
3862 2162 FEUT I 2162
3863 0000 FEXT
3864 1160 TAD HORD
3865 7510 SFA
3866 7400 JMS I LIM
3867 4407 FENTER
3868 5013 FEUT I 5013
3869 2162 FEUT I 2162
3870 0000 FEXT
3871 1160 TAD HORD
3872 7510 SFA
3873 7400 JMS I LIM
3874 4407 FENTER
3875 5013 FEUT I 5013
3876 2162 FEUT I 2162
3877 0000 FEXT
3878 1160 TAD HORD
3879 7510 SFA
3880 7400 JMS I LIM
3881 4407 FENTER
3882 5013 FEUT I 5013
3883 2162 FEUT I 2162
3884 0000 FEXT
3885 1160 TAD HORD
3886 7510 SFA
3887 7400 JMS I LIM
3888 4407 FENTER
3889 5013 FEUT I 5013
3890 2162 FEUT I 2162
3891 0000 FEXT
3892 1160 TAD HORD
3893 7510 SFA
3894 7400 JMS I LIM
3895 4407 FENTER
3896 5013 FEUT I 5013
3897 2162 FEUT I 2162
3898 0000 FEXT
3899 1160 TAD HORD
3900 7510 SFA
3901 7400 JMS I LIM
3902 4407 FENTER
3903 5013 FEUT I 5013
3904 2162 FEUT I 2162
3905 0000 FEXT
3906 1160 TAD HORD
3907 7510 SFA
3908 7400 JMS I LIM
3909 4407 FENTER
3910 5013 FEUT I 5013
3911 2162 FEUT I 2162
3912 0000 FEXT
3913 1160 TAD HORD
3914 7510 SFA
3915 7400 JMS I LIM
3916 4407 FENTER
3917 5013 FEUT I 5013
3918 2162 FEUT I 2162
3919 0000 FEXT
3920 1160 TAD HORD
3921 7510 SFA
3922 7400 JMS I LIM
3923 4407 FENTER
3924 5013 FEUT I 5013
3925 2162 FEUT I 2162
3926 0000 FEXT
3927 1160 TAD HORD
3928 7510 SFA
3929 7400 JMS I LIM
3930 4407 FENTER
3931 5013 FEUT I 5013
3932 2162 FEUT I 2162
3933 0000 FEXT
3934 1160 TAD HORD
3935 7510 SFA
3936 7400 JMS I LIM
3937 4407 FENTER
3938 5013 FEUT I 5013
3939 2162 FEUT I 2162
3940 0000 FEXT
3941 1160 TAD HORD
3942 7510 SFA
3943 7400 JMS I LIM
3944 4407 FENTER
3945 5013 FEUT I 5013
3946 2162 FEUT I 2162
3947 0000 FEXT
3948 1160 TAD HORD
3949 7510 SFA
3950 7400 JMS I LIM
3951 4407 FENTER
3952 5013 FEUT I 5013
3953 2162 FEUT I 2162
3954 0000 FEXT
3955 1160 TAD HORD
3956 7510 SFA
3957 7400 JMS I LIM
3958 4407 FENTER
3959 5013 FEUT I 5013
3960 2162 FEUT I 2162
3961 0000 FEXT
3962 1160 TAD HORD
3963 7510 SFA
3964 7400 JMS I LIM
3965 4407 FENTER
3966 5013 FEUT I 5013
3967 2162 FEUT I 2162
3968 0000 FEXT
3969 1160 TAD HORD
3970 7510 SFA
3971 7400 JMS I LIM
3972 4407 FENTER
3973 5013 FEUT I 5013
3974 2162 FEUT I 2162
3975 0000 FEXT
3976 1160 TAD HORD
3977 7510 SFA
3978 7400 JMS I LIM
3979 4407 FENTER
3980 5013 FEUT I 5013
3981 2162 FEUT I 2162
3982 0000 FEXT
3983 1160 TAD HORD
3984 7510 SFA
3985 7400 JMS I LIM
3986 4407 FENTER
3987 5013 FEUT I 5013
3988 2162 FEUT I 2162
3989 0000 FEXT
3990 1160 TAD HORD
3991 7510 SFA
3992 7400 JMS I LIM
3993 4407 FENTER
3994 5013 FEUT I 5013
3995 2162 FEUT I 2162
3996 0000 FEXT
3997 1160 TAD HORD
3998 7510 SFA
3999 740
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5406	3151	DLA A011Y	3566	4040	4040	
5407	4407	FENTER	3567	4040	4040	
5410	5564	FHEI NK	3570	4040	4040	
5411	0031	OUTH UR	3571	4040	4040	
5417	0057	S2	3572	4040	4040	
5418	0000	FEXT	3573	4040	4040	
5414	4404	WRITE	3574	4040	4040	
5415	3154	ME5G20	3575	4040	4040	
5416	7201	CLA TAL	3576	4040	4040	
5417	4401	FLOT	3577	4040	4040	
5420	5550	3550	3600	1617	1617	/NO
5421	5440	5440	3601	5617	5617	/D
5422	7000	CLA	3602	0640	0640	/F
5423	4403	FLOT	3603	0401	0401	/DA
5424	4700	2670	3604	2401	2401	/TA
5425	3440	HE51+	3605	4020	4020	/P
5426	4407	FENTER	3606	1711	1711	/OI
5427	5047	FHEI HE5	3607	1674	1674	/NT
5430	4176	FHEI V	3610	3777	3777	/S
5431	5261	ME5 M2400	3611	0000	0000	
5432	0030	FEXT	3612	4040	ME5G12,	4040
5433	4543	JMS I F11X	3613	4040	4040	
5434	1171	TAD RE51	3614	7305	2305	/SE
5435	1241	CLA +4	3617	1673	1673	/MS
5436	7001	FAE	3618	1124	1124	/IT
5437	4403	FLOT	3617	1176	1176	/IV
5440	3550	3550	3620	1124	1124	/T
5441	0000	0	3621	3172	3172	/Y1
5442	7000	CLA	3622	4000	4000	
5443	1741	TAD +	3623	4016	4016	/N
5444	3747	DLA +4	3624	0116	0116	/AN
5445	4401	FLOT	3625	1700	1700	/U
5446	5676	2670	3626	4017	4017	/M
5447	6400	0	3627	1103	1103	/IC
5450	7000	CLA	3630	7217	2217	/RO
5451	1247	TAD +	3631	0000	0000	
5452	5151	DLA A011Y	3632	4015	4015	/M
5453	4404	5404	3633	1114	1114	/L1
5454	1160	ME5011	3634	1411	1411	/L1
5455	6600	JHE I NKRE5	3635	0000	0000	
5456	0010	FH	3636	7617	MESG16,	2617
5457	7600	17600,	3637	1474	1424	/LT
5460	4000	14700,	3640	3700	3700	
5461	0014	M 400,	3641	3716	3716	/N
5462	3570	3570	3642	0577	0527	/E
5463	0000	0000	3643	4015	4015	/M
5464	0000	NK,	3644	0130	0130	/AX
5465	0000	0	3645	1115	1115	/TH
5466	0000	0	3646	7515	2515	/UM
5467	0000	0	3647	4077	4077	/P
5467	3737	ME511,	3650	4040	4040	
5468	3711	3711	3651	4000	4000	
5469	1674	1674	3652	3717	MESG18,	3717
5470	0577	0577	3653	0150	0150	/AX
5471	0607	0607	3654	5011	5011	/L1
5472	1717	1717	3655	1640	1640	/N
5473	1705	1705	3656	4571	4751	/X
5474	7407	7407	3657	7240	7240	/I
5475	7740	7740	3660	4000	4000	
5480	4075	5075	3661	3737	MESG19,	3737
5481	0015	2015	3662	1017	1017	/MO
5482	1157	5157	3663	2740	2740	/NO
5483	0000	0000	3664	1401	1501	/HA
5484	1301	ME567,	3665	1631	1631	/NY
5485	1740	1740	3666	4023	4023	/S
5486	4440	4440	3667	2007	2007	/FE
5487	0701	0701	3670	0374	0324	/CT
5490	0311	0311	3671	7201	2201	/KA
5491	0777	6777	3672	7740	7740	/P
5492	4077	4077	3673	4000	4000	
5493	4000	4000	3674	3737	MESG24,	3737
5494	1677	ME567,	3675	0524	0524	/S
5495	7740	7740	3676	4070	4020	/F
5496	0000	0000	3677	0716	0516	/EN
5497	1716	ME514,	3700	5440	5440	/I
5498	0700	0700	3701	2410	2410	/TH
5499	1477	ME567,	3702	0516	0516	/EN
5500	1700	1700	3703	4024	4024	/T
5501	5507	ME566,	3704	3120	3120	/YF
5502	1714	1714	3705	0540	0440	/F
5503	1715	1715	3706	4223	4223	/MS
5504	0574	0574	3707	4240	4240	/F
5505	0777	0777	3708	7417	2417	/TQ
5506	4075	4075	3711	4023	4023	/S
5507	5173	3123	3712	2401	2401	/TA
5508	7407	7407	3713	7774	7774	/RT
5509	1717	1717	3714	7200	7200	
5510	1717	1717	3715	3716	MESG27,	3716
5511	1717	1717	3716	0577	0577	/F
5512	1411	1411	3717	4006	4006	/EW
5513	1607	1607	3720	2205	2205	/RL
5514	4011	4011	3721	2525	2525	/BU
5515	1674	1674	3722	0566	0566	/EN
5516	0527	0527	3723	0331	0331	/CY
5517	2601	7601	3724	4014	4014	/L
5518	1477	1477	3725	1755	1755	/TH
5519	0000	0000	3726	1124	1124	/IT
5520	4017	ME509,	3727	7340	7340	/S
5521	1173	1103	3730	7740	7740	/P
5522	7717	7717	3731	4000	4000	
5523	1717	1617	3732	3706	MESG28,	3706
5524	0030	0000	3733	1417	1417	/F
5525	3733	3733	3734	2777	2777	/MI
5526	0117	0117	3735	4000	4000	
5527	2014	2014	3736	3706	ME5629,	3706
5528	0572	0772	3737	0711	0711	/HE
5529	4000	4000	3740	0710	0710	/GH
5530	3702	ME5010,	3741	7240	7240	/I
5531	0103	0103	3742	0000	0000	
5532	1307	1307	3743	1723	MESG30,	3723
5533	2727	227	3744	1517	1517	/S
5534	4000	4000	3745	1724	1724	/OI
5535	4040	ME5011,	3746	1007	1007	/HE

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1747 4077 4077 / ?
1750 4000 4000
3751 3715 MESG31, 3/15 / M
1752 1722 1722 /OR
1753 0540 0540 /E
1754 1116 1116 /TN
3755 2405 2405 /TE
3756 2006 2006 /RF
3757 0522 0522 /EK
3760 1707 1707 /OG
1761 3201 3201 /KA
3762 1523 1523 /PS
3763 7740 7740
3764 4000 4000
3765 1517 1517 /MD
3766 0425 0425 /DU
3767 1401 1401 /LA
3770 2411 2411 /LI
3771 1716 1716 /DM
3772 4001 4001 /A
3773 1520 1520 /MF
3774 1411 1411 /LI
3775 2425 2425 /TU
3776 0405 0405 /HC
3777 7200 7200 /?
      #4000
/Routine TO PUT PROGRAM ON "FIR-DATA" TAFF.
4000 /201 LIA IAC
4001 3247 DCA TFIAL /FIELD=0
4003 1207 TAD F440 /FIRST TAFF BLOCK
400J 3261 DCA TFIAL
4004 1210 TAD MB
4005 4235 JMS WRITE /WRITE 8 BLOCKS ON TAFF
4006 /402 HLT
4007 0440 P440: 440
4010 7770 MR: 7770 /-B
      /FFLAG 4047 /FIRI TIL
      TPBLN=4061
      WRTITT: 4035
/ - - - - -

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1747 4077 4077 / ?
1750 4000 4000
3751 3715 MESG31, 3/15 / M
1752 1722 1722 /OR
1753 0540 0540 /E
1754 1116 1116 /TN
3755 2405 2405 /TE
3756 2006 2006 /RF
3757 0522 0522 /EK
3760 1707 1707 /OG
1761 3201 3201 /KA
3762 1523 1523 /PS
3763 7740 7740
3764 4000 4000
3765 1517 1517 /MD
3766 0425 0425 /DU
3767 1401 1401 /LA
3770 2411 2411 /LI
3771 1716 1716 /DM
3772 4001 4001 /A
3773 1520 1520 /MF
3774 1411 1411 /LI
3775 2425 2425 /TU
3776 0405 0405 /HC
3777 7200 7200 /?
      #4000
/Routine TO PUT PROGRAM ON "FIR-DATA" TAFF.
4000 /201 LIA IAC
4001 3247 DCA TFIAL /FIELD=0
4003 1207 TAD F440 /FIRST TAFF BLOCK
400J 3261 DCA TFIAL
4004 1210 TAD MB
4005 4235 JMS WRITE /WRITE 8 BLOCKS ON TAFF
4006 /402 HLT
4007 0440 P440: 440
4010 7770 MR: 7770 /-B
      /FFLAG 4047 /FIRI TIL
      TPBLN=4061
      WRTITT: 4035
/ - - - - -

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820  
/      "FIRRTL"

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0176 0000 QUOL,
0177 0000 FLAG,
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      #4.00
/TELETYPE INPUT SUBROUTINE.
4200 0000 KARR, 0
4201 6032 KCC
4202 6031 KSF
4203 5202 JMF -1
4204 6036 KRR
4205 5600 JMF I KRBX
4206 0000 ANS1, 0
4207 4200 JMS KRBX
4210 3140 DCA ANSWR
4211 1140 TAD ANSWR
4212 4405 TYPE
4213 4200 JMS KRBX
4214 3732 DCA RETRN
4215 1232 TAD RETRN
4216 1233 TAD MATRN
4217 7650 SNA CLA
4218 5230 JMF -110
4219 1732 TAD RETRN
4222 1234 TAD MKUB
4223 7640 SZA CLA
4224 5713 JMF ANS
4225 1235 TAD BCKSL
4226 4405 TYPE
4227 5207 JMF ANS+11
4230 4546 JMS I FLRLP
4231 5606 JMF I ANS1
4237 0000 RETRN, 0
4238 7563 MKTRN, -715
4239 7401 MKUB, 377
4240 0334 BCKSL, 374
/INTEGER INPUT AND CONVERSION.
4236 0000 SICONV, 0
4237 7300 CLA CLL
4240 3305 DCA HOLD
4241 5263 JMF SINFUT
4242 3136 DCA STORI
4243 1136 TAD STOR1
4244 1303 TAD M'AO
4245 7510 SFA
4246 5270 JMF SICTRL
4247 1304 TAD SICTRL
4250 7740 SNA SZA CLA
4251 5270 JMF SICTRL
4252 1605 TAD HOLD
4253 7106 CLL KTL
4254 1305 TAD HOLD
4255 7004 KAL
4256 3305 DCA HOLD
4257 1136 TAD STORI
4260 0307 ANH SIMASH
4261 1305 TAD HOLD
4262 3305 DCA HOLD
4263 6031 SINFUT, KSF
4264 5741 JMF -1
4265 6036 KRR
4266 6046 KRS
4267 5242 JMF SICONV+4
4270 7200 SICTRL, CLA
4271 1136 TAD STOR1
4272 1233 TAD MKTRN
4273 7650 SNA CLA
4274 5300 JMF +4
4275 1235 TAD BCKSL
4276 4405 TYPE
4277 5277 JMF SICONV+1
4280 1305 TAD HOLD
4301 5636 JMF I SICONV
4302 0017 SIMASH, 17
4303 7520 M260, -260
4304 7767 SIM'71, -11
4305 0000 HOLD, 0
/FAITH FOR FL.PNT. INTERPRETER.
/FALL INDIRECT COMMANDS TO FIELD DEFINED BY FL.DF
4306 0000 FDF0, 0
4307 7300 CLA CLL
4310 1344 TAD DFO
4311 3321 DCA INDIA+1
4312 5706 JMF I FDF0
4313 0000 FINF1, 0
4314 7300 CLA CLL
4315 1345 TAD DFI
4316 3371 DCA INDJR+1
4317 5713 JMF I DFI
4320 0000 INDIR, 0
4321 0000 DFI, 0
4322 1321 TAD I
4323 3332 DCA STORE1+1
4324 5720 JMF I INDIR
4325 0000 RESET, 0
4326 7006 RTL
4327 6201 CDF 0
4330 5725 JMF I RESET
4331 0000 STOR1, 0
4332 0000 0
4333 7731 JMF I STORE1
4334 0000 NEXT1, 0
4335 7300 CLA CLL
4336 6201 CDF 0
4337 5734 JMF I NEXT1
4340 1344 DIR, 0
4341 3332 DCA STORF+1
4342 7741 JMF I +1
4343 6630 LOOK-01
4344 6201 CDF 0
4345 6711 DFI, CDF 10
4346 0000 PATCH, 0
4347 1141 TAD LOAD
4350 7646 SZA CLA
4351 5746 JMF I PATCH
4352 1140 TAD MKRD
4353 7510 SFA
4354 7041 CIA
4355 7700 SMA CLA
4356 5746 JMF I PATCH
4357 2157 ISZ EXP
4360 7000 NDF
4361 1160 TAD MKRD
4362 7130 CTL KAK
4363 3160 DCA MKRD
4364 5746 JMF I PATCH
/ERROR SUBROUTINE.
4365 0000 RUB1, 0
4366 4404 WRITF
4367 4373 MSG
4370 1365 TAD BUG1
4371 7701 HLT
4372 5765 JMF I BUG1
4373 3707 3707
4374 2222 7722
4375 1732 1732
4376 3700 3700
4400 0000 /SUBROUTINE TO FLOT ASCII CHARACTERS.
4401 0314 FLTASC, 0
4402 1354 DCA MASK77
4403 3317 TAD ASCII
4404 1150 DCA ASL
4405 3717 TAD ASFIX
4406 1151 DLA 15
4407 3214 TAD ASLIY
4408 7001 DCA +4
4409 4403 IAC
4410 7001 FLOT
4411 4403 0
4412 0000 0
4413 0000 0
4414 7700 CLA CLL
4415 1715 TAD I ASC
4416 3315 DCA ASC
4417 1315 TAD ASC
4418 0316 AND MASKNUM
4419 7106 CLL KTL
4420 7006 RTL
4421 7041 CIA
4422 3317 DCA FLCNT
4423 7040 FMA
4424 1715 TAD ASC
4425 0730 AND MASKADR
4426 1717 TAD CHAR
4427 3011 TAD INEXC
4428 1411 TAD INEXC
4429 3322 DCA MATRIX
4430 7344 STA CIL KAL
4431 3323 DLA MKRXN
4432 1327 TAD MATRIX
4433 7012 KTK
4434 7012 KTK
4435 0314 FLAS1, 0
4436 3344 AND MASK7
4437 1354 DCA MATRIX
4438 1324 TAD MATRIX
4439 0326 AND MASKX
4440 7110 CLL KAK
4441 7017 KTK
4442 0731 DCA I XLKDI
4443 1774 TAD MASKX
4444 0357 AND MASKY
4445 3717 DLA I YCRDI
4446 4733 JMS I SIZE1
4447 1731 TAD I XCRDI
4448 1150 TAD ASLIX
4449 3270 DCA +11
4450 1732 TAD I YKDI
4451 1151 TAD ASLIY
4452 3717 DCA +7
4453 1124 TAD MATRIX
4454 0327 AND MASKFN
4455 7640 SZA CLA
4456 7001 IAC
4457 4403 FLOT
4458 0000 0
4459 0000 0
4460 1322 TAD MATRIX
4461 7511 ISZ MKRXN
4462 5242 JMF FLASC7
4463 7300 CIA CIL
4464 2317 ISZ FLCNT
4465 5737 JMF FLASC1
4466 0717 FLU
4467 1330 TAD NEXT
4468 3731 DCA I XCRDI
4469 3732 DCA I YKDI
4470 4733 JMS I SIZE1
4471 1731 TAD I XCRDI
4472 1150 TAD ASLIX
4473 1732 TAD I YKDI
4474 1151 TAD ASLIY
4475 3151 DCA ASLIX
4476 5600 JMF I FLTASC
4477 0077 MASK77, 0077
4478 0000 ASC, 0
4479 7000 MASKNUM, 7000
4480 0000 PLINT, 0
4481 0777 MASKADR, 0777
4482 4721 CHAR, CHARAL
4483 0000 MATRIX, 0
4484 0000 MKRXN, 0
4485 0000 MASKX, 0
4486 0007 MASKY, 0007
4487 0030 MASKX, 0030
4488 0046 MASKFN, 0046
4489 0005 NEXT, 0005
4490 4717 XCRDI, XLKDI
4491 4701 YKDI, YKDI

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4131	4637	FSIZF1	SIZEF1	4714	5704	JMP I SIZE2	
4134	4533	ASCII	+1	4715	0000	SIZE3	0
4135	1000		/0	4716	0000	SZECNT	0
4136	4001		/1A	4717	0000	XCRB	0
4137	5005		/2B	4720	0000	YCRB	0
4140	4012		/3C	4721	4040	CHARAC	4040
4141	3016		/4D	4722	0215		0215
4142	3021		/5E	4723	2532		2532
4143	2024		/6F	4724	3042		3042
4144	4026		/7B	4725	3272		3272
4145	3034		/10H	4726	0525		0525
4146	3037		/11I	4727	3423		3423
4147	3042		/12J	4730	0363		0363
4150	3045		/13K	4731	3231		3231
4151	2040		/14L	4732	2000		2000
4152	1042		/15M	4733	7120		7120
4153	1044		/16N	4734	1001		1001
4154	6056		/17O	4735	0445		0445
4155	3064		/20P	4736	3534		3534
4156	4067		/21Q	4737	0525		0525
4157	4075		/22R	4740	3431		3431
4160	6101		/23S	4741	2000		2000
4161	1107		/24T	4742	7000		7000
4162	3111		/25U	4743	0535		0535
4163	2114		/26V	4744	4323		4323
4164	5116		/27W	4745	0535		0535
4165	2121		/30X	4746	4323		4323
4166	1123		/31Y	4747	7120		7120
4167	7130		/32Z	4750	1001		1001
4170	1000		/33	4751	0415		0415
4171	1000		/34	4752	2534		2534
4172	1000		/35	4753	6232		6232
4173	1000		/36	4754	3070		3070
4174	1000		/37	4755	0643		0643
4175	1231		/40+SFACE	4756	3375		3375
4176	3140		/41+	4757	3070		3070
4177	2242		/42+	4760	5030		5030
4180	1000		/43	4761	6025		6025
4181	1000		/44	4762	3535		3535
4182	6112		/45Z	4763	4110		4110
4183	1000		/46K	4764	2031		2031
4184	000		/47	4765	3575		3575
4185	110		/50+	4766	0525		0525
4186	2222		/51+	4767	0253		0253
4187	1000		/52R	4770	3070		3070
4188	1114		/53+	4771	4300		4300
4189	1126		/54+	4772	3070		3070
4191	1127		/55+	4773	0513		0513
4193	1230		/56+	4774	3540		3540
4194	1231		/57+	4775	0532		0532
4195	5132		/60+0	4776	7530		7530
4196	3137		/61+	4777	5001		5001
4197	4142		/62+	5000	0415		0415
4198	6146		/63+	5001	2534		2534
4199	3154		/64+	5002	3120		3120
4202	1157		/65+	5003	1044		1044
4203	6164		/66+	5004	3575		3575
4204	2122		/67+	5005	0525		0525
4205	7174		/70+8	5006	1433		1433
4206	6705		/71+9	5007	2202		2202
4207	2152		/72+	5010	5001		5001
4210	2144		/73+	5011	0415		0415
4211	2144		/74	5012	2534		2534
4212	2216		/75+	5013	3120		3120
4213	2246		/76+	5014	1041		1041
4214	5250		/77+	5015	3070		3070
4215	0000	SIZE1	0	5016	0525		0525
4216	1117	TAD XLAB	/1A ARB	5017	3433		3433
4217	4104	IMF SIZE	/ORIENTATION	5020	2202		2202
4240	1117	DCA XCRD		5021	6230		6230
4241	1110	TAD YCRD		5022	4110		4110
4242	4304	IMF SIZE		5023	2031		2031
4243	1110	DCA YCRD		5024	3223		3223
4244	1154	TAD ORIENT	/ORIENTATION	5025	1304		1304
4245	2450	CMA		5026	1525		1525
4246	5435	IMF I SIZE1		5027	3474		3474
4247	2700	SMA LIA		5030	6025		6025
4250	5260	IMF I10		5031	4535		4535
4251	1117	TAD XCRD		5032	4501		4501
4252	7041	CIA		5033	1020		1020
4253	5317	DCA XCRD		5034	3135		3135
4254	1110	TAD YCRD		5035	4510		4510
4255	7041	CIA		5036	3575		3575
4256	5320	DCA YCRD		5037	4500		4500
4257	1635	IMF I SIZE1		5040	1320		1320
4260	7040	CMA		5041	3575		3575
4261	1114	TAD ORIENT		5042	3545		3545
4262	7540	SMA LIA		5043	3070		3070
4263	1274	IMF +11		5044	4110		4110
4264	1420	TAD YCRD		5045	2031		2031
4265	1115	DCA SIZE3		5046	3545		3545
4266	1117	TAD XCRD		5047	0312		0312
4267	7041	CIA		5050	2233		2233
4270	3120	DCA YCRD		5051	4735		4735
4271	1115	TAD SIZE3		5052	0030		0030
4272	3417	DCA XCRD		5053	5001		5001
4273	5635	IMF I SIZE1		5054	0415		0415
4274	1517	TAD XCRD		5055	2534		2534
4275	3315	DCA SIZE3		5056	3120		3120
4276	1320	TAD YCRD		5057	1050		1050
4277	7041	CIA		5060	5425		5425
4700	3317	DCA XCRD		5061	2050		2050
4701	1315	TAD SIZE3		5062	3070		3070
4702	3320	DCA YCRD		5063	4415		4415
4703	5635	IMF I SIZE1		5064	2534		2534
4704	0000	SIZE2	0	5065	3300		3300
4705	1315	DCA SIZE3		5066	3070		3070
4706	1153	TAD SIZE	/SIZE	5067	4415		4415
4707	7041	CIA		5070	2534		2534
4710	3316	DCA SZECNT		5071	2313		2313
4711	115	TAD SIZE1		5072	6332		6332
4712	2316	IS1 SZECNT		5073	3120		3120
4713	5111	IMF 2		5074	1001		1001

5075	6502	6502	/4	5253	5263	152 FLAG2	/NOT FOUND
5076	3275	3275		5254	4555	JMS I RUG	/0
5077	3070	3070		5255	4206	JMS RNMWD	/READ AGAIN
5100	4110	4110	/5	5256	4406	READ	
5101	2031	2031		5257	4217	JMS FORKRD	
5102	3223	3223		5260	4406	READ	
5103	0305	0305		5261	3136	DCA STOR1	
5104	3575	3575		5262	5734	JMF CONV1 ?	
5105	7425	7425	/6	5263	0000	FLAG2	
5106	1504	1504		5264	0000	DIGIT	0
5107	0110	0110		5265	3000	IBM	0
5110	2031	2031		5266	5501	IMHBL	+113
5111	3223	3223		5267	7747	7747	/IBM 9
5112	1302	1302		5270	7770	7770	/0
5113	4535	4535	//	5271	7771	7771	/-7
5114	1050	1050		5272	7752	7752	/6
5115	5302	5302	/8	5273	7753	7753	/5
5116	0110	0110		5274	7774	7774	/4
5117	2031	2031		5275	7755	7755	/-3
5120	3223	3223		5276	7776	7776	/1
5121	1304	1304		5277	7777	7777	/-1
5122	1525	1525		5300	7746	7746	/0
5123	3423	3423		5301	7766	MINIO	7766
5124	4110	4110	/9			/SUBROUTINE TO TYPE 16 FLOT AS LII	
5125	2031	2031				/CHARACTERS	
5126	3425	3425		5302	0000	TYPE1	0
5127	1504	1504		5303	1136	DLI STOR1	
5130	0312	0312		5304	52	TAD TYPE1	/16 OK FLOT
5131	2233	2233		5305	524	SZA CLA	
5132	7070	7070	/SPACE	5306	5315	JMF 17	/FLOT
5133	3555	3555	/%	5307	1136	TAD STOR1	/TYPE
5134	0504	0704		5310	6041	ISZ	
5135	1415	1415		5311	5310	JMF -1	
5136	4131	4131		5312	6046	1135	
5137	3020	3020		5313	7300	CLA CLA	
5140	2161	2161		5314	5702	JMF I TYPE1	
5141	6011	6011	/1	5315	1136	TAD STOR1	
5142	1425	1425		5316	4230	JMS I FLOISC	
5143	5021	5021	/3	5317	5702	JMF I TYPE1	
5144	2415	2415		5310	4400	FL15L	CLTASC
5145	4232	4232	/4			/SUBROUTINE TO TYPE ON FLOT FXT.	
5146	6024	6024		5321	0000	TYPEX	0
5147	7021	7021	/5	5322	7300	CLA CLA	
5150	4232	4232	/6	5323	7321	TAD I TYPEX	/GET FJINTER
5151	6020	6020	/7	5324	3337	DCA TYPEX	/AND SAVE IT TO ALLY
5152	3575	3575	//	5325	3321	ISZ TYPEX	
5153	6020	6020	/1	5326	1737	TAD I TYPEX	/GET LEFT HAND
5154	6424	6424	/1	5327	7012	RTR	/CHARACTER
5155	6021	6021		5328	7012	RTR	
5156	6474	6424		5331	7017	RTR	
5157	4131	4131	/8	5332	4340	JMS TYPE	/CONVERT AND TYPE
5160	7303	7303		5333	1737	TAD I TYPEX	/RIGHT HAND CHAR.
5161	6721	6020	/1	5334	1337	ISZ TYPEX	/POINT TO NEXT WORK
5162	6020	6020		5335	4340	JMS TYPE	/CONVERT AND TYPE
5163	5425	5425	/9	5336	5476	JMF .10	/CONTINUE UNTIL DONE
5164	6475	6435		5337	0000	TYPEX	0
5165	7302	7302	/8	5340	0000	TYPEX	0
5166	1171	3171		5341	0177	AND TR77	
5167	4332	4132	/	5347	7400	SMA	
5170	0141	0141		5348	7231	JMF I TYPEX	/EXIT II TERMINATION
5171	4304	4304	/7	5349	1360	TAD TRN37	
5172	1525	1525		5345	7440	SZA	/TEST FOR CH LI
5173	3443	3443		5346	5357	JMF +4	/NOT A 37
5174	2221	2221		5347	1361	TAD TR75	/TYPE A FR
5175	6020	6020		5350	4405	TYPE	
		5200		5351	1362	TAD TRN125	/CONVERTS TO A LF
				5352	7510	9EA	/TEST RANGE
				5353	1363	TAD TR100	/RANGE IS 301 156
				5354	1364	TAD TR237	/RANGE IS 740 777
				5355	4405	TYPE	/TYPE CHARACTER
				5356	5740	JMF I TYPE	
				5357	0077	TR77	77
				5360	7741	TRN77	37
				5361	0215	TR215	215
				5362	7653	TRN125	125
				5363	0100	TR100	100
				5364	0137	TR237	237
						/ARRANGE RETURN LINE FEED.	
				5365	0000	CLFL	0
				5366	7100	CLA CLA	
				5367	1361	TAD TR215	
				5370	4405	TYPE	
				5371	1774	TAD IF	
				5372	4405	TYPE	
				5373	5765	JMF I CLFL	
				5374	0212	LF	12
						/LINE FEED	
						/INCREMENTAL PLOTTER SUBROUTINE.	
				5400	0000	FLOTX	0
				5401	7510	SPA	/MOVE THE PEN
				5402	5740	JMF FLOTA	/NO:CONTINUE
				5403	1364	TAD FLOTH	/AND F-N STATUS
				5404	112	CLA RTR	
				5405	7710	SFA CLA	/ANY CHANGE?
				5406	5227	JMF F 171	/NO:UN INUE
				5407	7670	SNL CLA	
				5410	5214	JMF +4	/LOWER THE PEN
				5411	1364	DCA FLOT 4	/RAISE THE PEN
				5412	5214	CLA	
				5413	5214	JMF +5	
				5414	1464	ISZ FLOTPN	/LOWER THE PEN
				5415	6524	FLOTPN	
				5416	4373	JMS FLOTPN	/WAIT FOR FLAG
				5417	5227	JMF FLOTPN	/CONTINUE
				5420	7700	FLOTA	CLA
				5421	604	FLOTA	
				5422	3364	DCA FLOTH	/RAISE THE PEN
				5423	3365	DCA FLOTH	/O TO X TO ORD.
				5424	3366	DCA FLOTH	/O TO Y TO ORD.
				5425	4373	JMS FLOTH	
				5426	5600	JMF I FLOTX	
				5427	1465	FLOTH	/SET A PREVIOUS
				5430	7141	CLA CLA	/X CO-ORDINATE.

5200	0000	READ1	0				
5201	7200	CLA					
5202	6011	RSF					
5203	5202	JMF .1					
5204	6016	JMF I READ1					
5205	5600						
5206	0000	READER					
5207	6011	RSF					
5210	5707	JMF .1					
5211	6632	RBA					
5212	4406	READ					
5213	4406	READ					
5214	4406	READ					
5215	7200	CLA					
5216	5606	JMF I BLANK					
5217	0000	READER					
5220	6011	RSF					
5221	5220	JMF .-1					
5222	6634	RFO					
5223	4406	READ					
5224	4406	READ					
5225	4406	READ					
5226	7200	CLA					
5227	5617	JMF I FORWD					
5230	0000	HECH					
5231	3136	DCA STOR1					
5232	7240	STA					
5233	7263	DCA FLAG2					
5234	1101	TAD MINIO					
5235	3164	DCA DIGIT					
5236	1264	TAD DIGIT					
5237	1266	TAD IMHBL					
5240	1265	DCA IBM					
5241	1136	TAD STOR1					
5242	1665	TAD I IBM					
5243	7640	SZA CLA					
5244	5751	JMF +5					
5245	7001	TAD					
5246	1264	TAD DIGIT					
5247	7041	LIA					
5250	5630	JMF I CONV					
5251	2264	ISZ DIGIT					
5252	5236	JMF CONV1					

5431	1600	TAD I FLOTX	/FORM NX-NFX	5610	1216	TAD C27	
5432	7420	SNL	/L-01 NX NFX	5611	3157	DLA EXP	
5433	7041	CIA	/ABSOLUTE VALUE	5612	4407	FENTER	
5434	3367	DLA FLOTDX	/OF DIFFERENCE	5613	7000	FNOR	/NORMALIZE F.P.
5435	7004			5614	0000	FEIT	/NUMBER
5436	1377	DLA FLOTMV	/SAVE SIGN BIT	5615	5600	JMF I FLINTP	
5437	1600	TAD I PLOTX	/SE' NEW	5616	0027	0027	
5440	3367	DLA FLOTNX	/PREVIOUS X			/DOUBLE PRECISION DECIMAL-BINARY	
5441	2200	ISZ FLOTX	/INCR. POINTER			/INPUT AND CONVERSION	
5447	1366	TAD FLOTNV	/FETCH PREVIOUS	5617	0000	BECONV, 0	
5443	7141	CIA CLL	/Y CO ORDNATE.	5620	7300	CLA CLL	/INITIALIZE
5444	1600	TAD I FLOTX	/FORM NY-NPV	5621	3160	DLA HORD	/MANTISSA
5445	7420	SNL	/+01 NPV NY	5622	3161	DLA LORD	
5446	7041	CIA	/ABSOLUTE VALUE	5623	4406	READ	/READ DIGIT
5447	7141	DLA FLOTDY	/OF DIFFERENCE	5624	4547	JMS I CONVRT	/18N-BINARY
5450	1372	TAD FLOTMV	/SAVE SIGN BIT	5625	3136	DLA STOR1	
5451	7004	RAI		5626	4232	JMS MULT10	
5452	1372	DLA FLOTMV		5627	2156	ISZ COUNT	
5453	1600	TAD I FLOTX	/SET NEW	5630	5213	JMF -5	
5454	3166	DLA FLOTNV	/PREVIOUS Y	5631	5617	JMP I DECONV	/NEXT DIGIT
5455	2200	ISZ PLOTX	/INCR. POINTER	5632	0000	MULT10, 0	/MULTIPLY DOUBLE
5456	1367	TAD FLOTDX		5633	1161	TAD LORD	/PRECISION WORD
5457	7141	CIA CLL	/L=01	5634	3252	DLA LTEMP	/BY TEN(DECIMAL)
5460	1370	TAD FLOTDY	/DELTA Y DELTA X	5635	1160	TAD HORD	
5461	7670	SNL CLA		5636	3253	DLA HTEMP	
5462	5275	JMF FLOTX		5637	3254	DLA REMAIN	
5463	1367	TAD FLOTDX	/REVERSE NUMBERS	5640	4255	JMS MULT2	/MULTIPLY BY TWO
5464	3371	DLA -LOTHA		5641	4257	JMS MULT2	
5465	1370	TAD FLOTDY		5642	4271	JMS DUBLAD	/CALL DOUBLE ADD
5466	3367	DLA FLOTDX		5643	4257	JMS MULT2	
5467	1371	TAD FLOTDY		5644	1136	TAD STOR1	/LAST DIGIT
5470	3370	DLA FLOTDY		5645	3252	DLA LTEMP	/RECEIVED
5471	7001	IAL	/SET MAJOR MOT	5646	3253	DLA HTEMP	
5472	0372	AND PLOTMV	/ION INSTRUCTION	5647	4271	JMS DUBLAD	
5473	1345	TAD FLOTT1		5650	1254	TAD REMAIN	/EXIT WITH RE-
5474	5100	JMF -14		5651	5632	JMP I MULT10	/MAINDER IN AC.
5475	1372	TAD FLOTMV		5652	0000	LTEMP, 0	/DOUBLE PRECIS-
5476	7110	CIA RAK		5653	0000	HTEMP, 0	/ION WORD.
5477	1350	TAD FLOTT2		5654	0000	REMAIN, 0	
5500	3371	DLA FLOTHA		5655	0000	MULT2, 0	/MULTIPLY HORD.
5501	1771	TAD I FLOTHA		5656	7300	CLA CLL	/LORD BY TWO.
5503	1343	DLA FLOTHA		5657	1161	TAD LORD	
5503	1372	TAD FLOTHV	/SET COMB.MOTION	5660	7004	RAI	
5504	1353	TAD FLOTT3		5661	3161	DLA LORD	
5505	3372	DLA FLOTHV		5662	1160	TAD HORD	
5506	1372	TAD I FLOTHV		5663	7004	RAI	
5507	1314	DLA FLOTHV		5664	3160	DLA HORD	
5510	1367	TAD FLOTHV	/INITIALIZE PLOTHA	5665	1254	TAD REMAIN	
5511	7110	CLL RAK		5666	7004	RAI	
5512	1371	DLA FLOTHA		5667	3254	DLA REMAIN	
5513	1367	TAD FLOTDX	/SET STEP COUNTER	5670	5655	JMP I MULT2	
5514	7040	CMA		5671	9000	DUBLAD, 0	/DOUBLE PRECIS-
5515	1372	DLA PLOTHV		5672	7300	LLA CLL	/ION ADDITION.
5516	1372	FLOTT3, ISZ PLOTHV	/FLOT ALGORITHM	5673	1161	TAD LORD	
5517	7410	SKP		5674	1252	TAD LTEMP	
5520	5600	JMF I FLOTX	/ALL DONE	5675	3161	DLA LORD	
5521	7100	CLL		5676	7004	RAI	
5522	1571	TAD PLOTHA		5677	1160	TAD HORD	
5523	1370	TAD FLOTDY		5700	1253	TAD HTEMP	
5524	3171	DLA PLOTHA		5701	3160	DLA HORD	
5525	7430	SZL	/PLOTHA 100007	5702	7004	RAI	
5526	3134	JMF FLOTHV	/YES	5703	1254	TAD REMAIN	
5527	1371	TAD PLOTHA	/NO	5704	3254	DLA REMAIN	
5530	7140	CMA CLL		5705	5671	JMP I DUBLAD	
5531	1367	TAD FLOTDX		5706	0000	MSION, 0	/ROUTINE TO FORM
5532	7610	SZL DLA		5707	7300	CLA CLL	/2S COMPLEMENT
5533	3343	JMF FLOTHA	/SINGLE MOTION	5710	2323	ISZ SGN1	/IF MINUS
5534	0000	DLA FLOTHV	/IONB.MOTION	5711	5706	JMF I MSION	/SIGN=0000:EXIT
5535	1367	TAD FLOTDX		5712	1161	TAD LORD	
5536	7041	CIA		5713	7041	CMA IAC	
5537	1371	TAD FLOTHA		5714	3161	DLA LORD	
5540	3371	DLA FLOTHA		5715	1160	TAD HORD	
5541	4373	JMS FLOTHV		5716	7040	CMA	
5542	316	JMF FLOTHV		5717	7430	SZL	
5543	0000	FLOTH4, 0		5720	7001	IAC	
5544	5841	JMF -3		5721	3160	DLA HORD	
5547	5546	FLOTT1, -1		5722	5706	JMF I MSION	
5548	6511	FLFR	/PEN-RIGHT	5723	0000	SGN1, 0	/=0 IF+=7777IF-
5549	6521	FLFI	/PEN-LEFT			THE EXPONENT.	
5550	5751	FLOTT2, -1		5724	0000	FECL, 0	
5551	6512	FLDU	/DRUM-UP	5725	7300	LLA CLL	
5552	6514	FLDD	/DRUM DOWN	5726	1157	TAD EXP	
5553	5554	FLOTT3, -1		5727	7510	SFA	
5554	6513	FLUK	/UP-RIGHT	5730	7061	CMA IAC CML	
5555	6573	FLUL	/UP-LEFT	5731	3157	DLA EXP	
5556	6515	FLUK	/DOWN-RIGHT	5732	1370	TAD C253	
5557	4760	JMS -11	/DOWN-LEFT	5733	7430	SZL	
5560	0000	0		5734	1371	TAD C255	
5561	6514	FLDD		5735	4776	JMS I DGP1	
5562	6571	FLUL		5736	3160	DLA HORD	
5563	5760	JMF I -3		5737	1157	TAD EXP	
5564	0000	FLOTHN, 0		5740	2160	SZL HORD	
5565	0000	FLOTHX, 0		5741	1372	TAD H144	
5566	0000	FLOTHY, 0		5742	7500	SMA	
5567	0000	FLOTHZ, 0		5743	5340	JMP -3	
5568	0000	FLOTHV, 0		5744	1373	TAD C144	
5569	0000	FLOTHW, 0		5745	3157	DLA EXP	
5571	0000	FLOTHA, 0		5746	7040	CMA	
5572	0000	FLOTHV, 0		5747	1160	TAD HORD	
5573	0000	FLOTHW, 0		5750	7440	SZL	
5574	6501	FLSF	/WAIT FOR DONE FLAG	5751	4776	JMS I DGP1	
5575	5374	JMF -1	/NOT YET	5752	3160	DLA HORD	
5576	6502	FLCF	/CLEAR FLAG	5753	1157	TAD EXP	
5577	5773	JMF I FLOTHV	/EXIT	5754	2160	ISZ HORD	
		#5600		5755	1374	TAD H12	
		/FLOATING POINT INPUT FROM PAPERTAPE.		5756	7500	SMA	
5600	0000	FLINTP, 0		5757	5354	JMF -3	
5601	1565	TAD I STOR1		5760	1375	TAD C12	
5602	3136	DLA STOR1		5761	3161	DLA LORD	
5603	1336	TAD I STOR1	/NUMBER OF DIGITS	5762	7240	CLA CMA	
5604	7041	CIA		5763	1160	TAD HORD	
5607	3156	DLA COUNT		5764	4776	JMS I DGP1	
5608	2565	ISZ I 601	/INCR.RETURN ADDR.	5765	1161	TAD LORD	
5609	4217	JMS DECONV					

5766	4776	JMS I TGT	6146	1351	TAD C200
5767	4714	JMF T FEXC	6147	4405	TYPE
5770	2773	C253, .253 260	6148	7405	JMI I OUTDO
5771	0002	1.253, .253-253	6151	0.60	0260
5772	7654	H144, 7674	6152	0000	0
5773	0144	C.44, 0.44	6153	7110	CLL KAR
5774	2766	M12, 2766	6154	3365	DCA STOK
5775	0017	C12, 0012	6155	1160	TAD HORD
5776	6147	DGT, OUTDO	6156	7010	KAR
		66000	6157	1160	DCA HORD
		/FLOATING OUTPUT FROM KAM.	6160	1161	TAD LORD
6000	0000	FLOUT, 0	6161	7010	KAR
6001	1565	TAD I GOI	6162	1161	DCA LORD
6002	3136	DCA STOK	6163	1365	TAD STOK
6003	1536	TAD I STOK	6164	5752	JMI I DIVTUM
6004	0327	AND MASH2	6165	0000	STOK, 0
6005	7010	RAR	6166	7775	TENH, 7777, /10
6006	7110	CLL RAR	6167	4146	3146
6007	7110	CLL KAR	6170	3147	3147
6010	3166	DCA FORMAT	6171	0004	TEN, 0004 /10
6011	1536	TAD I STOK	6172	7400	2400
6012	0130	AND MASH1	6173	0000	0000
6013	3167	DCA FURK		76700	
6014	2765	152 I GOI		/IN THE COMMENTS OF M	
6015	1160	TAD HORD		/ F NUMBER OF DIGITS TO BE OUTPUT	
6016	7710	SPA CLA		/ D = NUMBER OF DECIMAL PLACES	
6017	1337	TAD MINUS		/ F DECIMAL EXPONENT	
6020	1155	TAD SPACE		/ F NUMBER OF PLACES REMAINING TO BE	
6021	4405	TYPE		PRINTED BEFORE THE FINAL POINT	
6022	1334	TAD BEXF	6200	0000	0
6023	7010	DCA INDEX	6201	1166	TAD FORMAT
6024	5237	JMP FOJTCN	6202	7450	SNA
6025	1331	TAD HX	6203	277	KAR
6026	3157	JMS T FEXC	6204	0004	CLA
6027	4714	JMS T FEXD	6205	1167	TAD FORM
6030	5600	JMI I FLOUT	6206	7510	SPA
6031	7200	CLA	6207	14	JMP 145
6032	1337	TAD CHE	6210	7740	CLA LMA
6033	4405	TYPE	6211	1166	TAD I FEXAT
6034	4744	JMS I FEXFPT	6212	3167	DCA FORM
6035	7200	CLA	6213	7040	TAD X
6036	5600	JMI I FLOUT	6214	1177	TAD X
6037	1160	TAD HORD	6215	7500	SNA
6040	7700	SNA CLA	6216	7200	CLA
6041	5245	JMP 14	6217	1177	TAD FORMAT
6042	7040	CMA	6220	7510	SPA
6043	3741	DCA I SMT	6221	5250	ME FRNT-1
6044	4740	JMS T FEXFPT	6222	1354	TAD MING
6045	7240	TAD CHE	6223	7500	SNA
6046	1157	TAD EXP	6224	7000	CLA
6047	3157	DCA EXP	6225	1555	KAR
6050	3331	DCA BEXF	6226	5644	DCA TEMX
6051	1157	TAD EXP	6227	1166	TAD HORD
6052	7200	CLA	6228	1564	TAD TEMX
6053	5266	JMP FG03	6231	1366	DCA FILE
6054	1336	TAD FOUR	6232	1564	TAD TEMX
6055	7700	SNA CLA	6233	7041	CLA
6056	5273	JMP FG04	6234	1366	DCA TEMX
6057	4407	FENTER	6235	7107	CLL IAC KAR
6060	3371	FMFY TEN	6236	7266	152 FILE
6061	0000	FEXT	6237	1766	TAD I FILE
6062	7240	CLA CMA	6240	1375	TAD NIO
6063	1331	TAD BEXF	6241	7110	SPA CLA
6064	3131	DCA BEXF	6242	5754	JMI FRNT-1
6065	5251	JMP FG02	6243	3766	DCA FILE
6066	4407	FENTER	6244	2564	152 TEMX
6067	3366	FMFY TENTH	6245	5313	JMI DEX
6070	0000	FEXT	6246	7266	152 I FILE
6071	2131	152 BEXF	6247	1177	152 EXI
6072	7251	JMP FG02	6250	7200	CLA
6073	3136	DCA STOK	6251	1356	FRNT-1
6074	4743	JMS I FEXPT	6252	3010	DCA INDEX
6075	4742	JMS I FLOFT	6253	1166	TAD FORMAT
6076	7410	SPA	6254	7450	SNA
6077	4152	JMS DIVTUM	6255	5342	JMP FLOP
6100	2157	152 EXP	6256	7041	CLA
6101	5277	JMP 2	6257	3367	DCA FLOUT
6102	7450	SNA	6260	1367	TAD FLOUT
6103	5314	JMP FG07	6261	1177	TAD EXI
6104	3410	DCA I INDEX	6262	7540	SNA J/A
6105	1335	TAD MINUS	6263	5317	JMP XXX
6106	3157	DCA EXP	6264	1167	TAD FORM
6107	4742	JMS I FLOFT	6265	7500	SNA
6110	3410	DCA I INDEX	6266	7200	CLA
6111	2157	152 EXP	6267	7041	CLA
6112	5277	JMP 2	6270	1157	TAD EXI
6113	7251	JMP FG07	6271	7041	CLA
6114	7240	CLA CMA	6272	1364	DCA TEMX
6115	1331	TAD BEXF	6273	1361	TAD M7
6116	3331	DCA BEXF	6274	3365	DCA J COUNT
6117	1160	TAD HORD	6275	1157	TAD EXP
6120	7640	SPA CLA	6276	1564	TAD TEMX
6121	5125	JMP 14	6277	7650	SNA CLA
6122	1161	TAD LORD	6300	5130	JMP DIG
6123	7650	SNA CLA	6301	1564	TAD TEMX
6124	3331	DCA BEXF	6302	7001	IAC
6125	7240	CLA CMA	6303	7710	SPA CLA
6126	5305	JMP FG06	6304	1360	TAD SPACE
6127	0070	MASH2, 0070	6305	4323	JMS OUTX
6130	0007	MASH1, 0007	6306	2364	152 TEMX
6131	0000	REX, 0	6307	5775	JMP RACK
6132	0305	CHE, 305	6310	1367	TAD POINT
6133	7772	MINUS7, 7772	6311	4757	JMS I OFUT
6134	6167	BKST, BUFFER-1	6312	5275	JMI BACK
6135	6200	FXAD, FIX1	6313	7040	DCA
6136	0004	FOUR, 0004	6314	1366	TAD PLACE
6137	0017	SEVEN, 0017	6315	3366	DCA FILE
6140	7640	MSN7, MS16N	6316	5136	JMP RET
6141	5273	SNPT, SGN	6317	7200	XXX, CLA
6142	5631	MI0FT, MULT10	6320	1361	TAD IHX
6143	5675	MCFT, MUL2	6321	4757	JMS OUTX
6144	5724	TEMT, T	6322	5136	JMP 2
6145	0000	OUTDO, 0	6323	0000	OUTX, 0
					/OUTPUT 1 DIGIT.

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6374 4257 JMS I OFUT /FRINT CHARACTER
6325 2367 ISZ FOUNT /# CHARACTERS PRINTED*
6326 5773 JMF I OUIX /NO, RETURN
6327 5600 JMF I FIXI /YES, NUMBER FINISHED
6330 7040 DIG, FMA
6331 1157 IAD EXP /REDUCE E BY 1
6332 3157 DCA EXP
6333 2345 ISZ SCOUNT /# SIG.FIGS.PRINTED*
6334 5340 JMF ,F4 /NO
6335 7040 CMA /YES,
6336 3345 DCA SCOUNT /RESET COUNT TO -1
6337 5305 JMF IN /AND LEAVE C(AC) = 0
6340 1410 TAB I INDEX /TAKE NEXT DIGIT
6341 5705 JMF IN
6342 1324 FLOF, TAB MIN6 /FRINT & DIGITS
6343 3367 DCA FOUNT /AFTER DECIMAL POINT
6344 4757 JMS I OFUT /FRINT "0"
6345 1367 TAB FOINI
6346 4777 JMS I OFUT /FR, " "
6347 7700 ISZ FIXI /NIMEM.RETURN, ATIK.
6350 1410 TAB I INDEX /TAKE NEXT DIGIT
6351 4323 JMS OUIX /FRINT IT
6352 5350 JMF ,Z2 /AND REPEAT
6353 0007 K7, 7
6354 7772 MIN6, 6
6355 7766 N10, 12
6356 6367 BUFFER, BUFFER
6357 6147 OFUT, OUIX
6358 7760 SECT, 77-260
6361 7771 M7, /
6362 7776 FOINI, 776 760
6363 0050 LHX, 330 260
6364 0000 FIFX, 0
6365 0000 SCOUNT, 0
6366 0000 FICL, 0
6367 0000 FOUNT, 0
6370 0000 BUFFER, 0
6371 6372 OUTPUT BUFFER
6372 44400
/FLLOATING POINT SINE.
6400 0000 FFIN, 0
6401 1160 TAB HOKD /X 0?
6402 7740 SNA SIA CLA /NO
6403 5210 JMF ,F5 /YES
6404 1160 TAB HORD
6405 7700 SNA CLA /NO X 0?
6406 4600 JMS I FFIN /YES SIN(0) 0
6407 4363 JMS FNEG /SIN(X) SIN(X)
6410 3147 DCA FNTR /X MODULO 2FI
6411 4407 FENTER
6412 4317 FDIU TWOFI
6413 6310 FDIU XSOK
6414 0000 FEXT
6415 4547 JMS I FIXR
6416 4350 JMS FLOA
6417 4407 FENTER
6420 6325 FDIU X
6421 7710 FDIU XSOK
6422 4325 FDIU X
6423 3117 FDIU TWOFI
6424 6325 FDIU X
6425 7710 FDIU FI /X FI?
6426 0000 FEXT
6427 1160 TAB HOKD
6428 7710 SFA CLA
6429 7741 MF ,F10 /YES
6431 4407 FENTER /NO
6432 6325 FDIU X /SIN(X-FI)=SINX
6433 6325 FDIU X
6434 0000 FEXT
6435 1347 TAB FNTR
6436 7650 SNA CLA
6437 7040 CMA
6440 1347 DCA FNTR
6441 4407 FENTER
6442 5725 FDIU X
6443 7714 FSUB FIOT
6444 0000 FEXT
6445 1160 TAB HORD
6446 7710 SIA CLA
6447 7257 JMF ,F6 /YES
6448 4407 FENTER /NO
6449 5327 FDIU FI /SINX-SIN(FI-X)
6451 2327 FDIU X
6452 6325 FDIU X
6453 0000 FEXT
6454 4407 FENTER
6455 5325 FDIU X
6456 7710 FDIU FIOT
6457 4314 FDIU X
6460 6325 FDIU X
6461 4325 FDIU X
6462 6310 FDIU XSOK
6463 5111 FDIU X
6464 4310 FDIU XSOK
6465 1536 FADD C7
6466 3330 FDIU XSOK
6467 1341 FADD C3
6470 3330 FDIU XSOK
6471 1344 FADD C3
6472 3330 FDIU XSOK
6473 1514 FADD FIOT
6474 3325 FDIU X
6475 0000 FEXT
6476 2347 ISZ FNTR
6477 5600 JMF I FFIN
6500 4363 JMS FNEG
6501 7200 CLA
6502 5600 JMS I FFIN
6503 0000 FCOS, 0
6504 4407 FENTER /COSX=SIN(PI/2-X)
6505 6377 FDIU X
6506 3514 FDIU FIOT
6507 2325 FSUB X
6510 0000 FEXT
6511 1303 TAB FCOS
6512 3700 DCA FFIN
6513 5201 JMF FFINI
/CONSTANTS AND POINTERS.
6514 0001 P101, 0001 /PI/2
6515 3110 3110
6516 3755 3755
6517 0003 TWOFI, 0003 /28PI
6520 3110 3110
6521 3755 3755
6522 0002 PI, 0002 /PI
6523 3110 3110
6524 3755 3755
6525 0000 X, 0
6526 0000 0
6527 0000 0
6530 0000 XSOK, 0
6531 0000 0
6532 0000 0
6533 7764 C9, 7764
6534 2366 2366
6535 5735 5735
6536 7771 C7, 7771
6537 5466 5466
6540 6317 6317
6541 7775 C5, 7775
6542 2431 2431
6543 5053 5053
6544 0000 C3, 0000
6545 5325 5325
6546 0420 0420
6547 0000 FNTR, 0
6550 0000 FLOA, 0
6551 3160 DCA HORD
6552 7300 CLA CL
6553 3161 DCA LORD
6554 1562 TAB C13
6555 3157 DCA EXP
6556 4407 FENTER
6557 7000 FNOR
6558 0000 EXT
6559 5750 JMF I FLOA
6562 0013 C13, 0013
/NEGATION SUBROUTINE.
6563 0000 FNEG, 0
6564 3167 JMS I ACHN /CALL SUBROUTINE
6565 7240 CLA CMA /IN INTERPRETER
6566 5763 JMF I FNEG
6567 7000 ACHN, ACHNS
/LOADING SQUARE SUBROUTINE.
6570 0000 SQUARE, 0
6571 4407 FENTER
6572 6162 FDIU FPAC1
6573 3162 FMFY FPAC1
6574 0000 FEXT
6575 5770 JMF I SQUARE
6576 4400 4400
/FLLOATING POINT ARITHMETIC INTERPRETER.
6600 0000 FFNT, 0
6601 4773 JMS I NEXT2
6602 3173 FDIU DULH
6603 3174 DCA OVER2
6604 1600 TAB I FNT
6605 3254 DCA JUMP
6606 1254 TAB JUMP
6607 0204 AND PADEND
6610 7650 SNA CLA
6611 5214 JMF ,F3 /YES
6612 1262 TAB MASK5
6613 0200 AND FNTR
6614 3257 DCA ADDR
6615 1263 TAB MASK7
6616 0254 AND JUMP
6617 1257 TAB ADDR
6620 3257 DCA ADDR
6621 1265 TAB INDIRT
6622 0254 AND JUMP
6623 7650 SNA CLA
6624 5774 JMF I DIR1 /NO-60 ON
6625 1857 TAB I ADDR /YES DEFER
6626 3257 DCA ADDR
6627 4775 JMS I INDIR1
6630 2200 LOOF 01, ISZ FFNT
6631 1657 TAB I ADDR
6632 3170 DCA EXI
6633 1257 TAB ADDR
6634 3240 DCA SAVE
6635 2240 ISZ SAVE
6636 1640 TAB I SAVE
6637 3171 DCA AC1H /HIGH ORDER MANTISSA
6640 2240 ISZ SAVE
6641 1640 TAB I SAVE
6642 3172 DCA AC1L /LOW ORDER MANTISSA
6643 1254 TAB JUMP
6644 7106 CLL RPL
6645 4776 JMS I RESET1
6646 0261 AND MASK3 /GET BITS 0-2
6647 1266 TAB TABLE /OFFCODE LOOKUP
6650 3255 DCA JUMP2 /IN TABLE
6651 1855 TAB I JUMP2
6652 3255 DCA JUMP2
6653 5655 JMF I JUMP2 /GO THERE
6654 0000 JUMP, 0
6655 0000 JUMP2, 0
6656 0000 SDC, 0
6657 0000 ADDR, 0
6660 0000 SAVE, 0
6661 0017 MASK3, 0017
6662 7600 MASK5, 7600
6663 0177 MASK7, 0177
6664 0200 PADEND, 0200

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6665	0400	INDRECT,	0400	2045	1366	TAD AMOUNT	
6666	6667	TABLE,	+1	2046	1367	TAD TFS1	/EAM EXPONENTS
6667	6742	LXIT	/TABLE USED IN	2047	2710	SFA CLA	/BI JUDGE
6670	6721	FLAD	/INTERPRETING	2050	5256	JMF +46	/YES +6
6671	6720	FLSU	/BITS 0 1 OF	2051	4115	JMS OUTGO	/NO
6672	6761	FLMY	/PSEUDO INSTRUCT	2052	1410	5/1	
6673	7310	FLUV	/ION IF JFCODE 0	2053	1174	TAD TAG	
6674	6677	FLGT	/GO TO EXIT AND	2054	1373	TAD TAG1	
6675	6706	FLPT	/INTERFRET	2055	5342	JMF NOGO	
6676	6737	MORF	/BITS 8-11	2056	4331	JMS OUTGO	
6677	1170	FLGT,	TAR EX1	2057	7420	SAR	/SET UP ADDRESS
6700	3157		DCA EXP	2060	1374	TAD TAG2	
6701	1171		TAD AC1H	2061	1173	TAD TAG1	
6702	3160		DCA HORD	2062	3370	DCA TFS1	
6703	1172		TAD AC1L	2063	1366	TAD AMOUNT	
6704	3161		DCA LORD	2064	1720	TAD I TFS1	
6705	5201		JMF FNTH1	2065	3720	DCA I TFS1	
6706	1157	FLF1,	TAD EXT	2066	2370	ISZ TFS1	
6707	4777		JMS I STORE2	2067	1170	TAD TFS1	
6711	2257		DCA I ADDR	2070	3171	DCA TFS14	
6712	1160		TAD HORD	2071	1371	ISZ TFS14	
6713	3657		DCA I ADDR	2072	1371	TAD TFS14	
6714	2257		ISZ ADDR	2073	3372	DCA TFS15	
6715	1161		TAD LORD	2074	1372	ISZ TFS15	
6716	3657		DCA I ADDR	2075	7740	CLA CMA	/SUBTRACT FROM HIT
6717	5201		JMF FNTH1	2076	1360	TAD AMOUNT	/COUNT OF 45
6720	4741	FLSU,	JMS I OFHINS	2077	5504	DCA SHIF15	/MULTIPLICATION
6721	4771	FLAD,	/FLAD 1 FIRST	100	1771	TAD I TFS14	/LOW ORDER
6722	5201		JMF FNTH1	2101	7421	MAR	
6723	4772		JMS I UNORH	2102	1770	TAD I TFS13	/HIGH ORDER
6724	1173		TAD OVER1	2103	0000	0	
6725	1174		TAD OVER2	2104	0000	SHIF13,	
6726	3174		DCA OVER2	2105	3701	DCA SHIF13	/HIGH ORDER
6727	7004		RAL	2106	7501	MAR	/OVERFLOW
6730	1172		TAD AL1L	2107	3313	DCA UN160	/OVERFLOW
6731	1161		TAD LORD	2110	1366	TAD AMOUNT	/OVERFLOW
6732	3161		DCA LORD	2111	1332	TAD MIN1	
6733	7004		MAR	2112	7510	SFA	
6734	1171		TAD AC1H	2113	3371	JMF TFS15	/LESS THAN 15 SHIFTS
6735	1160		TAD HORD	2114	3371	DCA AC	/OVERFLOW
6736	3160		DCA HORD	2115	1771	TAD I TFS14	/LOW ORDER
6737	4770	NORF,	JMS I NORM	2116	7471	MAR	
6740	5201		JMF FNTH1	2117	1770	TAD I TFS13	/HIGH ORDER
6741	7412	OPHINS,	MINUS2	2118	7412	CLA	
6742	1254	EXIT,	TAD JUMP	2119	0000	0	
6743	6261		AND MASK3	2122	2701	MAR CLA	
6744	7450		SAR	2123	3722	DCA TFS1	/OVERFLOW
6745	5600		JMF I FNTH	2124	1364	TAD MIN1	
6746	1360		TAD AC0M6	2125	3720	DCA TFS13	
6747	3255		DCA JUMP2	2126	1373	TAD OUTGO	
6750	1655		TAD JUMP2	2127	3721	DCA TFS14	
6751	3255		DCA JUMP2	2130	770	ISZ ALGN	
6752	1200		TAD FNTH	2131	5770	JMF I ALGN	
6753	7756		DCA G12	2132	7767	MAR TFS1	
6754	4655		JMS I JUMP2	2133	0000	0	
6755	1256		TAD G02	2134	1170	TAD EX1	/EXAMINE WHICH
6756	5200		DCA FNTH	2135	7041	CLA EX1	/TO 1171
6757	5201		JMF FNTH1	2136	1157	TAD EXF	
6760	7544	AC0M6,	TABLE6-1	2137	7004	RAL	
6761	7201	FLMY,	CLA TAC	2140	2200	CLA	/CLERK ON EX1 EXF
6762	1170		TAD EX1	2141	5233	JMF I OUTGO	
6763	1157		TAD EXF	2142	1170	DCA TFS15	/EAM BE ALGN
6764	3157		DCA EXP	2143	1770	TAD I TFS15	/EAM BE ALGN
6765	4767		JMS I MULT	2144	3157	DCA EXF	/EAM
6766	5201		JMF FNTH1	2145	3720	ISZ TFS15	
6767	7221	MULT,	DMULT	2146	1770	TAD I TFS15	
6770	7600	NORH,	DNORH	2147	3160	DCA HORD	
6771	7020	ALGN,	ALGN	2148	1770	ISZ TFS15	
6772	7564	UNORH,	DUNORH	2151	1770	TAD I TFS14	
6773	4334	NEXT2,	NEXT1	2152	3161	DCA LORD	
6774	4740	DIR1,	DIR	2153	3670	JMF I ALGN	
6775	4720	INDIR1,	INDIR	2154	1170	TAD EX1	/MATH 1170
6776	4325	RESET1,	RESET	2155	3177	DCA EX1	
6777	4331	STORE2,	STORE1	2156	1360	JMF DONE	
			*7000	2157	7240	LESS1,	/SUBTRACT FROM HIT
7000	0000	ACHINS,	0	2160	1366	TAD AMOUNT	/COUNT FOR L
7001	7300		CLL CLA	2161	3371	DCA SHIF13	/OVERFLOW
7002	1174		TAD OVER2	2162	1777	TAD I TFS15	
7003	7041		CMA IAC	2163	7421	MAR	
7004	3174		ICA OVER2	2164	1771	TAD I TFS14	/LOW ORDER
7005	1161		TAD LORD	2165	1320	JMF SHIF13	
7006	7040		CMA	2166	0000	AMOUNT,	
7007	7430		SZL	2167	7720	TFS15,	10
7010	7101		CLL IAC	2170	0000	TFS15,	0
7011	3161		DCA LORD	2171	0000	TFS14,	0
7012	1160		TAD HORD	2172	0000	TFS15,	0
7013	7040		CMA	2173	0157	TAC1,	EXF
7014	7430		SZL	2174	0011	TAC1,	EX1 EXF
7015	7101		CLL IAC	2175	0000	EXF,	0
7016	3160		DCA HORD	2176	5770	JMF EX114	/LUMPY SUBROUTINE
7017	5600		JMF I ALHINS	2200	0000	DEVL,	0
7020	0000	ALIGN,	0	2201	7300	CLA CLL	/SHIFT FOR RIGHT
7021	1160		TAD HORD	2202	1160	TAD HORD	
7022	7640		SZA CLA	2203	7510	SFA	
7023	5222		JMF +4	2204	7101	CLA LML	
7024	1161		TAD LORD	2205	7010	KAR	
7025	7650		SNA CLA	2206	1160	CLA HORD	
7026	5354		JMF NOHERE	2207	1161	TAD LORD	
7027	1171		TAD AC1H	2210	7010	KAR	
7030	7640		SZA CLA	2211	3161	DCA LORD	
7031	5215		JMF +4	2212	1174	TAD OVER1	
7032	1172		TAD AC1L	2213	7010	KAR	
7033	7650		SNA CLA	2214	3174	DCA OVER2	
7034	5620		JMF I ALGN	2215	7100	CLA	
7035	1170		TAD EX1	2216	7527	ISZ EXF	
7036	7041		CMA IAC	2217	0000	NO	
7037	1157		TAD EXP	2218	0000	JMF I DIV	
7040	7430		SNA	2219	0007	IMULT,	0
7041	5330		JMF DONE	2220	7100	CLA CLL	/SAVE FROM TRIPLE
7042	7510		SFA	2221	4130	JMS IGN	/MULTIPLICATION
7043	7041		CMA IAC	2222	1171	TAD EX1	/MULTIPLY 4 TIME
7044	1366		ICA AMOUNT	2223	5230	DCA +45	

7227	74	7	TAD HOKI	7411	7777	GOFF	3///	
7227	74	7	MHI MUY	7412	6000	PHIL	...	/NEGATE D-LAND
7227	74	7		7413	7300			/TRIPLE RALISIM
7227	74	7		7414	...			
7227	74	7		7415	7041	CMA IAC		
7227	74	7		7416	7173	DCA OVERI		
7227	74	7		7417	7172	TAD ACIL		
7227	74	7		7420	7040	CMA		
7227	74	7		7421	7430	SZL		
7227	74	7		7422	7101	CLL IAC		
7227	74	7		7423	7172	DCA ACIL		
7227	74	7		7424	7171	TAD ACIL		
7227	74	7		7425	7040	CMA		
7227	74	7		7426	7430	SZL		
7227	74	7		7427	7101	CLL IAC		
7227	74	7		7430	7171	DCA ACIL		
7227	74	7		7431	5612	JMP I MINUS2		
7227	74	7				/HOURI		
7227	74	7				HURITIV		
7227	74	7		7437	0000			/IFERANDS NORMALIZED
7227	74	7		7438	3153	DCA DVAL		
7227	74	7		7439	1172	TAD ACIL		
7227	74	7		7440	7104	CLL IAC		
7227	74	7		7441	...	CA SVT		
7227	74	7		7442	1171	AF ACIL		
7227	74	7		7443	7004	RAI		
7227	74	7		7444	7276	DCA IAC		
7227	74	7		7445	7276	TAD IAC		
7227	74	7		7446	7250	CA DVAL		
7227	74	7		7447	1161	TAD IAC		
7227	74	7		7448	7431	RAI		
7227	74	7		7449	1160	TAD HOKI		
7227	74	7		7450	7407	DVAL		
7227	74	7		7451	0000	DVAL		
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## SUMMARY

This thesis deals with the investigation of the temperature- and magnetic field-dependence of the photoconductivity spectra of shallow donors and acceptors present in high purity germanium. These spectra lie in the difficult region of the far infrared. To enter this region a Michelson interferometer has been constructed capable to cover the wave number range of  $5\text{-}350\text{ cm}^{-1}$ . This instrument, the relevant detector system (bolometer cooled with liquid helium) and the experimental arrangement for photoconductivity measurements are described in Chapter II. This latter system allows measurements at controlled temperatures between 1 and 15 K.

In the progress of these investigations it was discovered that a substantial deviation from the real spectrum may be caused by radiation which is reflected back into the interferometer due to reflection on the sample. In Chapter III, which starts with the elementary theory of Fourier spectroscopy on which our experimental method is based, the theoretical and experimental investigation of this reflection-effect is discussed.

In Chapter IV the application of the interferometer to the study of the photoconductivity in ultra-pure germanium is treated. Using the technique of photo-thermal ionization spectroscopy the impurities, which are present in extremely low concentrations of  $10^{10}\text{-}10^{11}\text{ atoms/cm}^3$ , are analyzed. The binding energies of the electron/hole of the donor/acceptor impurities were determined very accurately from the temperature dependence of the peak-intensity in the hydrogen-like excitation spectra. The complete temperature dependence of the photo-thermal conductivity signal was thoroughly investigated, both theoretically and experimentally.

From measurements in a magnetic field, the linear and quadratic Zeeman-terms were determined for all lines of the boron acceptor excitation spectrum. The relevant values for these levels are compared with the available values from the literature. Some of the lines showing up in a magnetic field could be identified as states associated with light-hole Landau levels.

The computer programs developed for the processing of the measured interferogram data and executing the Fourier transform to get the frequency spectrum are collected in an Appendix. These programs are suitable for a PDP-12 (or, with slight modifications, for a PDP-8) laboratory computer.



Dit proefschrift behandelt het onderzoek naar de temperatuur- en magnetenveld-afhankelijkheid van de fotogeleidings spektra van ondiep liggende donors en acceptors die aanwezig zijn in ultra-zuiver germanium. Deze spektra liggen in het moeilijke gebied van het verre infrarood. Om door te dringen in dit gebied werd een Michelson interferometer gekonstrueerd die een bereik heeft van golfgetallen tussen 5 en  $350\text{ cm}^{-1}$ . Dit instrument, het bijbehorende detektor systeem (bolometer, gekoeld met vloeibaar helium) en de experimentele opstelling voor foto-geleidings metingen worden beschreven in Hoofdstuk II. Dit laatstgenoemde systeem maakt metingen mogelijk bij een gecontroleerde temperatuur tussen 1 en 15 K.

In de loop van het onderzoek werd ontdekt dat een aanzienlijke afwijking van het werkelijke spectrum kan worden veroorzaakt door straling die ten gevolge van reflectie op het sample wordt teruggekaatst in de interferometer. In Hoofdstuk III, dat begint met de elementaire theorie over Fourier spektroskopie waarop onze experimentele methode is gebaseerd, wordt het theoretische en experimentele onderzoek van dit reflectie-effekt besproken.

In Hoofdstuk IV wordt de toepassing van de interferometer bij het bestuderen van de foto-geleiding in ultra-zuiver germanium behandeld. De onzuiverheden, die aanwezig zijn in extreem lage concentraties van  $10^{10}$ - $10^{11}$  atomen/cm<sup>3</sup>, worden geanalyseerd door gebruikmaking van de techniek van foto-thermische ionisatie spektroskopie. De bindings-energieën van het elektron/gat van de donor/acceptor onzuiverheden werden zeer nauwkeurig bepaald uit de temperatuur-afhankelijkheid van de intensiteit van de pieken in de waterstof-achtige excitatie spektra. De volledige temperatuur-afhankelijkheid van het foto-thermische geleidings-sigitaal werd grondig onderzocht, theoretisch zowel als experimenteel.

Uit metingen in een magnetenveld werden de lineaire en kwadratische Zeeman-termen bepaald voor alle lijnen van het excitatie spectrum van de borium acceptor. De desbetreffende waarden voor deze niveau's worden vergeleken met de beschikbare waarden uit de literatuur. Enige van de lijnen die ontstaan in een magnetenveld konden worden geïdentificeerd als toestanden die verbonden zijn met Landau niveau's van lichte gaten.

De computer programma's die werden ontwikkeld voor het verwerken van de gemeten interferogram data en het uitvoeren van de Fourier transformatie om

het frekwentie spektrum te verkrijgen zijn verzameld in een Appendix. Deze programma's zijn geschikt voor een PDP-12 (of, met geringe wijzigingen, voor een PDP-8) laboratorium computer.

## CURRICULUM VITAE

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# STELLINGEN

## I

De extrinsieke geleiding van halfgeleiders wordt in het algemeen bepaald door de absolute waarde van het verschil van de donor- en acceptor-koncentraties. Onder bepaalde omstandigheden echter dragen donors en acceptors afzonderlijk bij tot de geleiding.

## II

Bij de bestudering van de bandenstructuur van materialen als  $\text{Cd}_{3-2x}\text{As}_x\text{P}_x$  via optische effecten kan de nauwkeurigheid van de analyse sterk worden vergroot door naast reflectie en transmissie-metingen de absorptie te bepalen via de calorische methode.

*M.J. Gelten, A. van Lieshout, C. van Es and F.A.P. Blom,  
J. Phys. C 11, 227 (1978)*

## III

Anders dan bij metalen kunnen bij halfmetalen veranderingen in de effectieve massa van de electronen als gevolg van de electron-fonon interactie niet worden verwaarloosd bij fonon-spektroskopie met behulp van de puntkontakt methode.

*A.G.M. Jansen, F.M. Mueller and P. Wyder, Science 199, 1037 (1978)*

## IV

Door een ongunstige verhouding tussen golflengte en bundel-diameter treden bij verre infrarood apparatuur vaak diffractie-problemen op, die vermeden kunnen worden door bundels met een gaussische intensiteits-verdeling te gebruiken. Een dergelijke intensiteits-verdeling zou kunnen worden verkregen door toepassing van metaal-maas filters met een continu verlopende strip-breedte.

*D.H. Martin and J. Lesurf, Infrared Physics 18, 405 (1978)*

## V

De optische eigenschappen van bepaalde materialen kunnen worden bestudeerd door de spiegels van een Fabry-Perot interferometer met deze materialen te bedekken. De werking van de interferometer kan worden geoptimaliseerd via de bedekkingsgraad.

*J.G. v.M. de Kort, Proefschrift, Nijmegen 1979*

## VI

Smith en Loewenstein geven een uitdrukking voor de efficiency van een beamsplitter, waarbij ook de absorptie in rekening wordt gebracht. Deze formule geeft echter alleen betrouwbare resultaten indien rekening wordt gehouden met de invalshoek-afhankelijkheid van de absorptie term.

*D.R. Smith and E.V. Loewenstein, App. Opt. 14, 2473 (1975)*

## VII

Het is mogelijk verre infrarood straling te detecteren via de gelijkrichtende werking van een puntkontakt tussen normale metalen bij lage temperatuur.

## VIII

Indien camera's in een lichte kleur zouden worden uitgevoerd i.p.v. het gebruikelijke zwart, zou de opwarming ten gevolge van geabsorbeerde zonnestraling en daardoor de kans op afwijkingen in de film veel geringer zijn.

## IX

Daar er geen direct verband bestaat tussen het brandstof-verbruik en de lawaai-productie van de verschillende auto-modellen, zal de inning van de heffing geluidshinder wegverkeer via een prijsverhoging voor de brandstoffen wel een stimulans zijn om te kiezen voor zuinige, maar niet voor geluidsarme auto's.

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